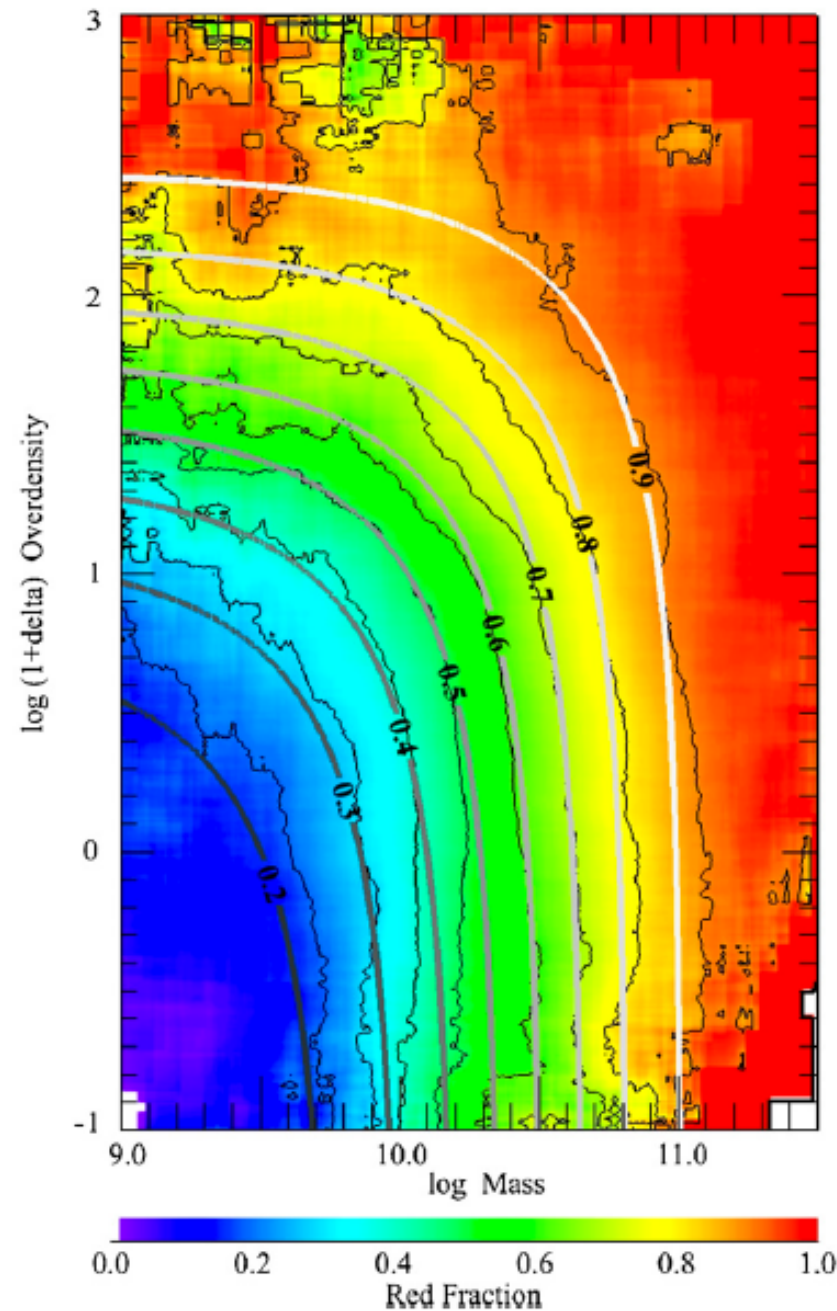


# Environmental effects on massive galaxy formation in most distant clusters

Tao Wang (University of Tokyo)

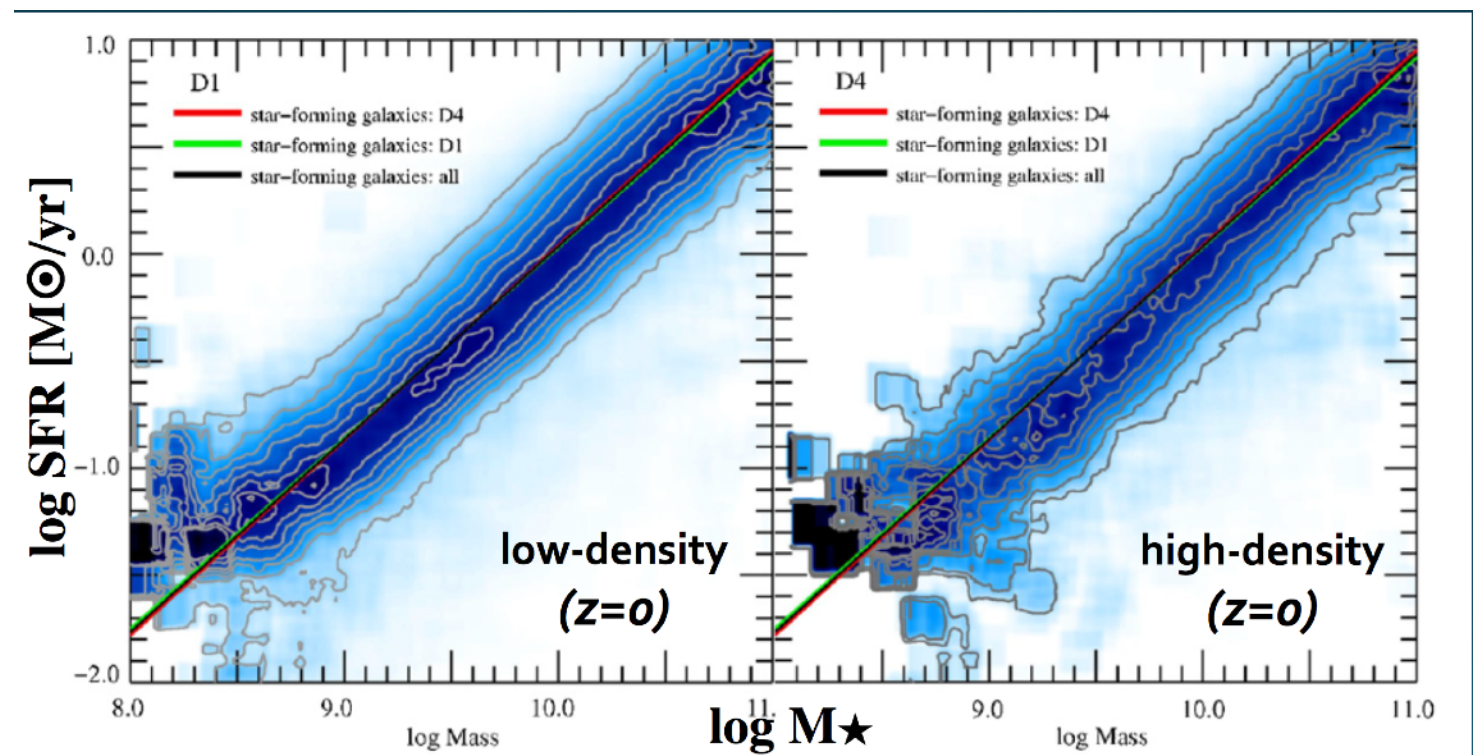


# Environmental effects on massive galaxy formation



Peng+2010

## Invariant star forming main-sequence

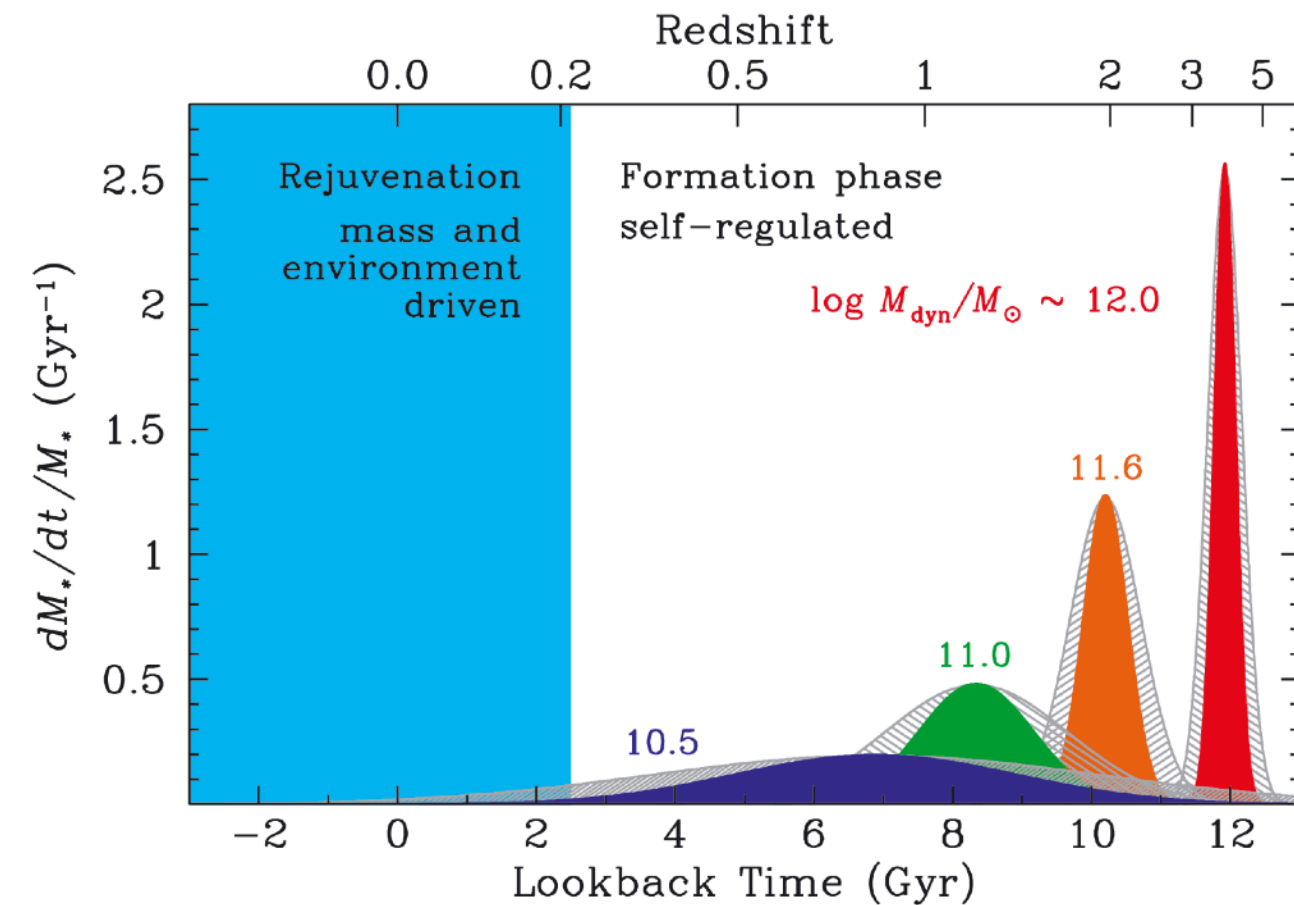


Baldry+2006

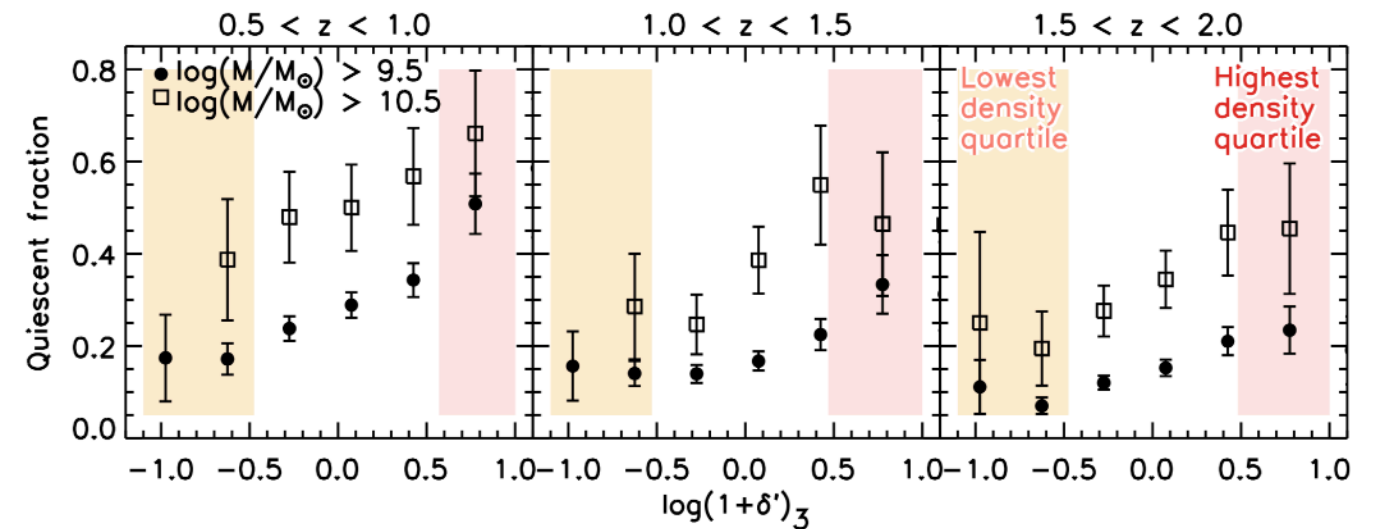
also see Koyama+2013 for high- $z$

The role of environment in massive galaxy formation/quenching remains unclear.

# The early assembly of massive (cluster) galaxies



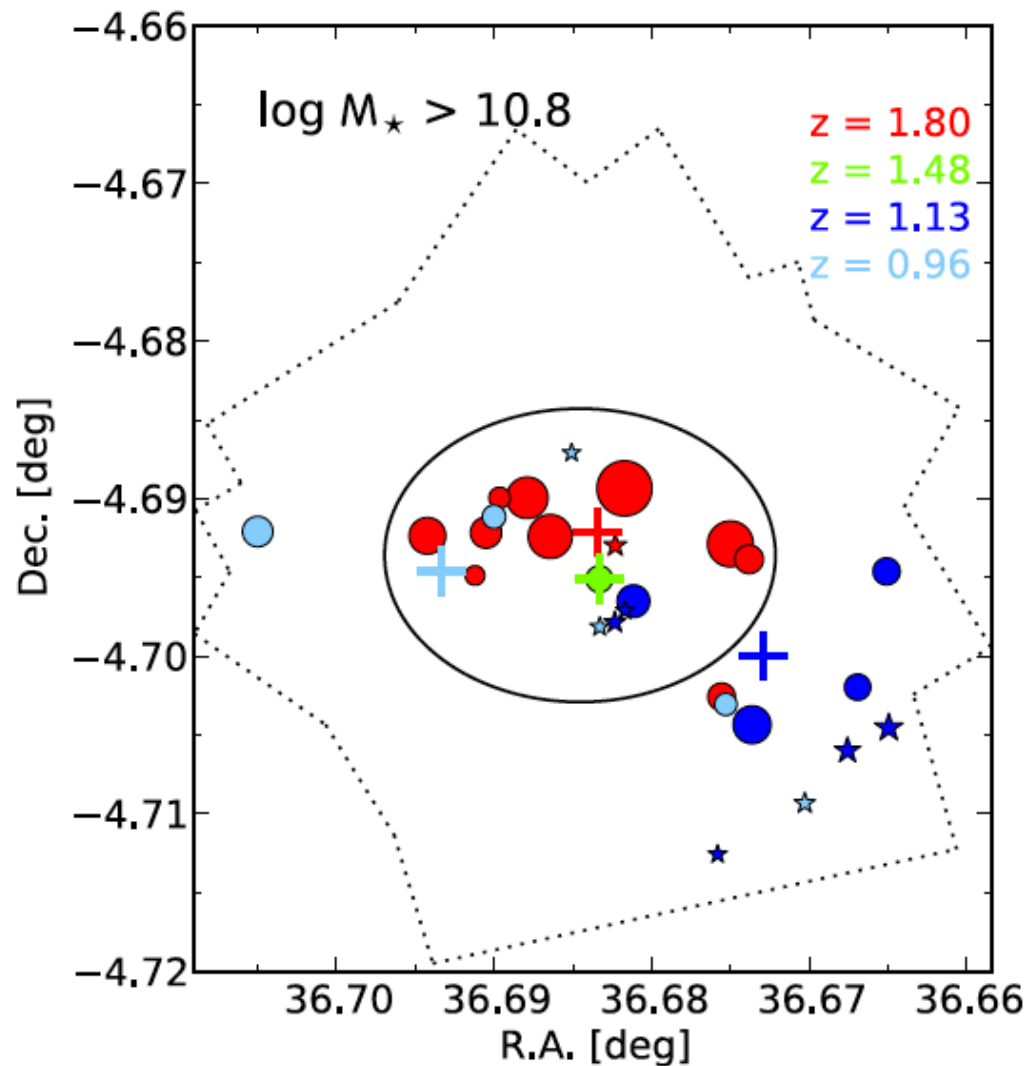
Thomas+2010



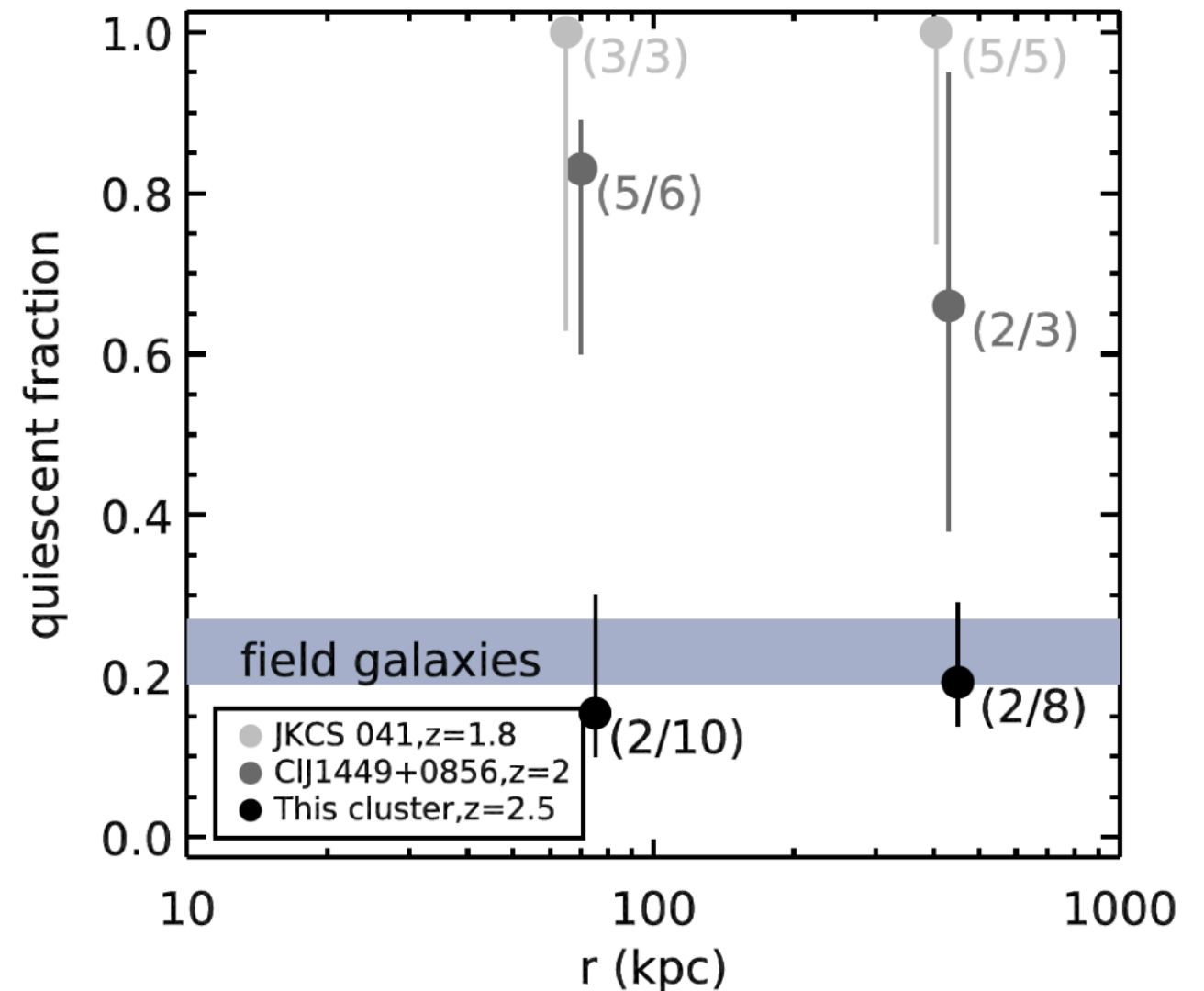
Kawinwanichakij+2017

# The early assembly of massive cluster galaxies:

## The core of the most massive clusters are already dominated by quiescent galaxies at $z \sim 2$



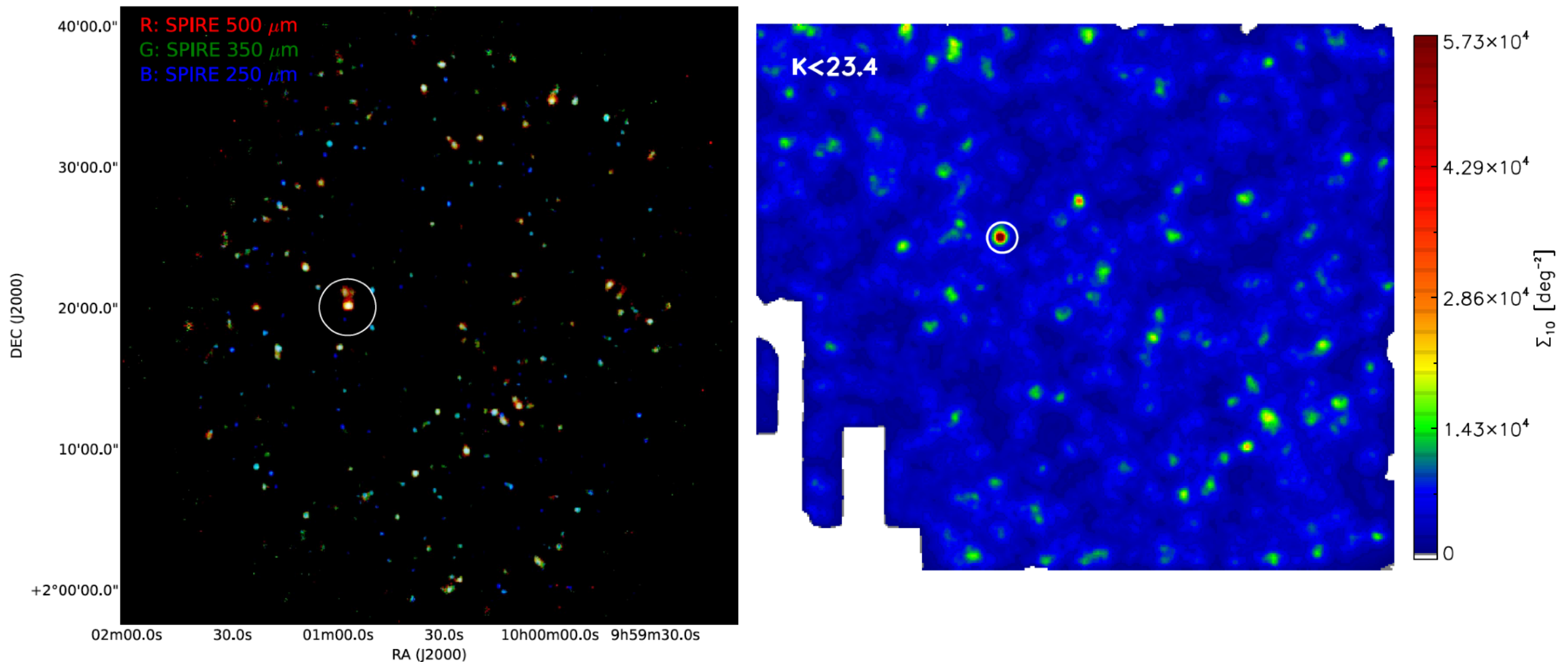
Newman+2014



Strazzullo+2013, Wang+2016

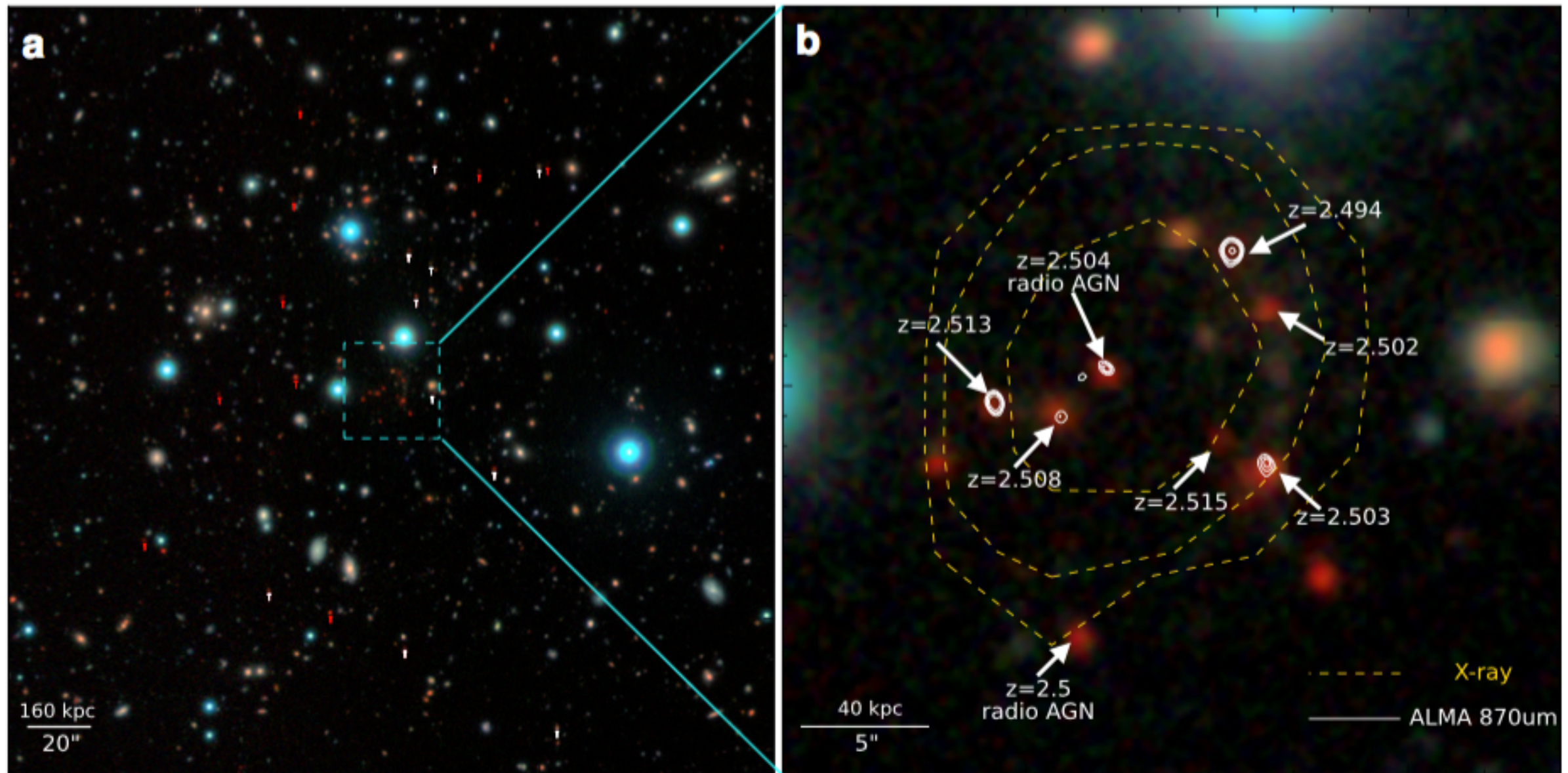


# Hunting for massive galaxy clusters in formation (**starbursting galaxy clusters**)



**Herschel/SPIRE-selected extremely luminous infrared sources**

# “Discovery of a Galaxy Cluster with a Violently Starbursting Core at $z=2.506$ ”

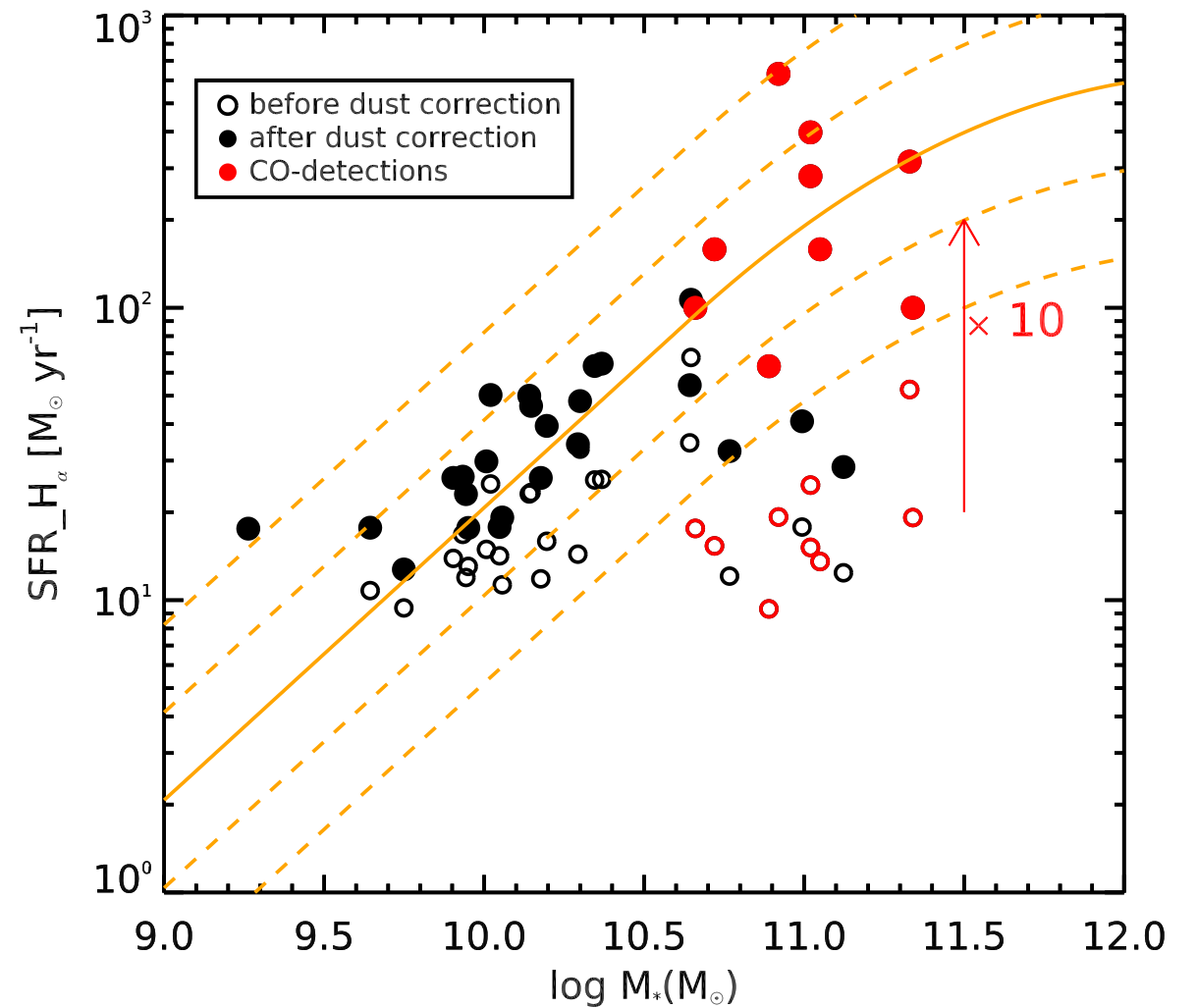
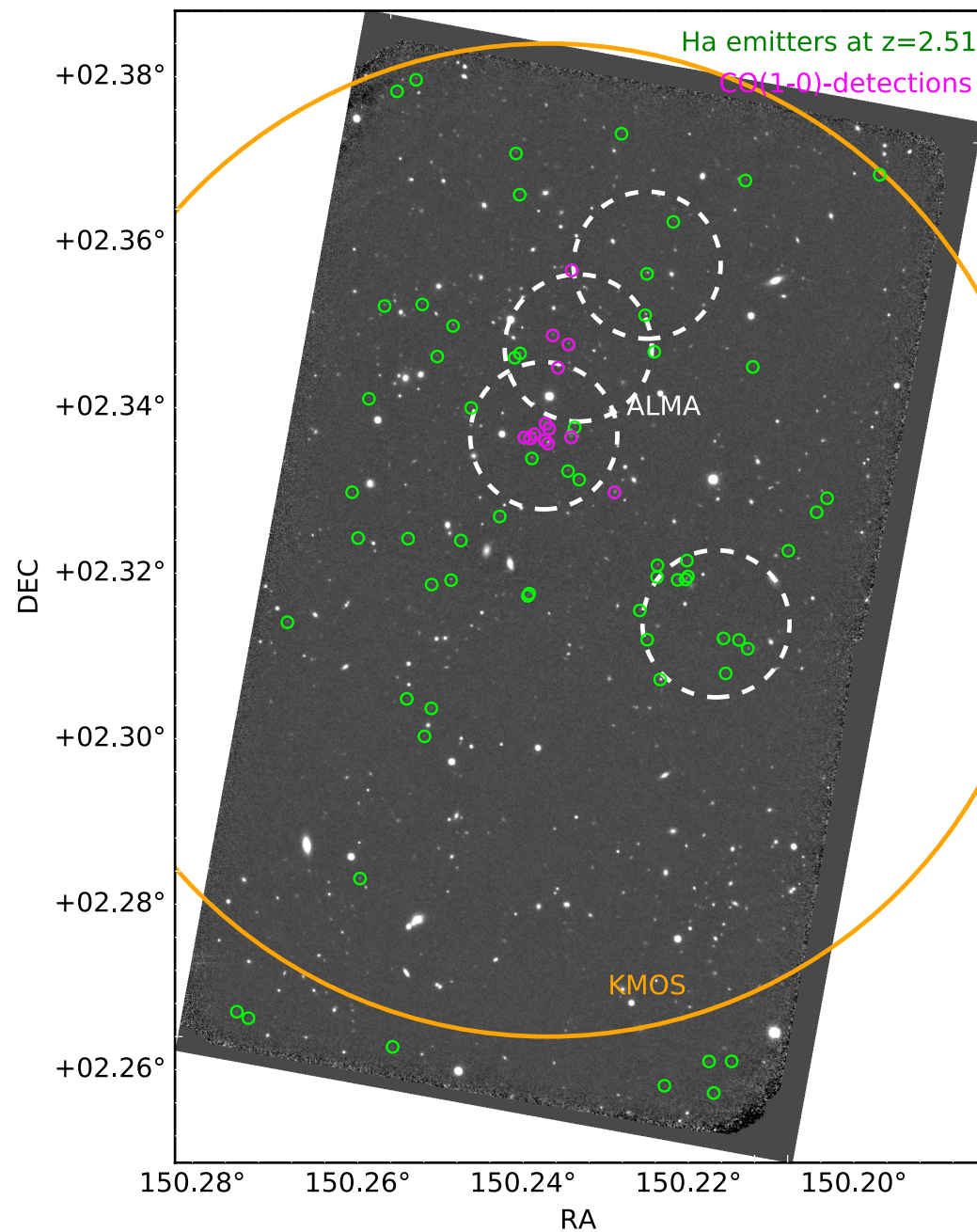


Presence of both **extended X-ray emission** and a **dominant population of massive SFGs**

Wang,T+2016, ApJ, 828,56



# Narrow-band imaging of J1001 with Subaru/MOIRCS: Towards a complete census of star-forming cluster members

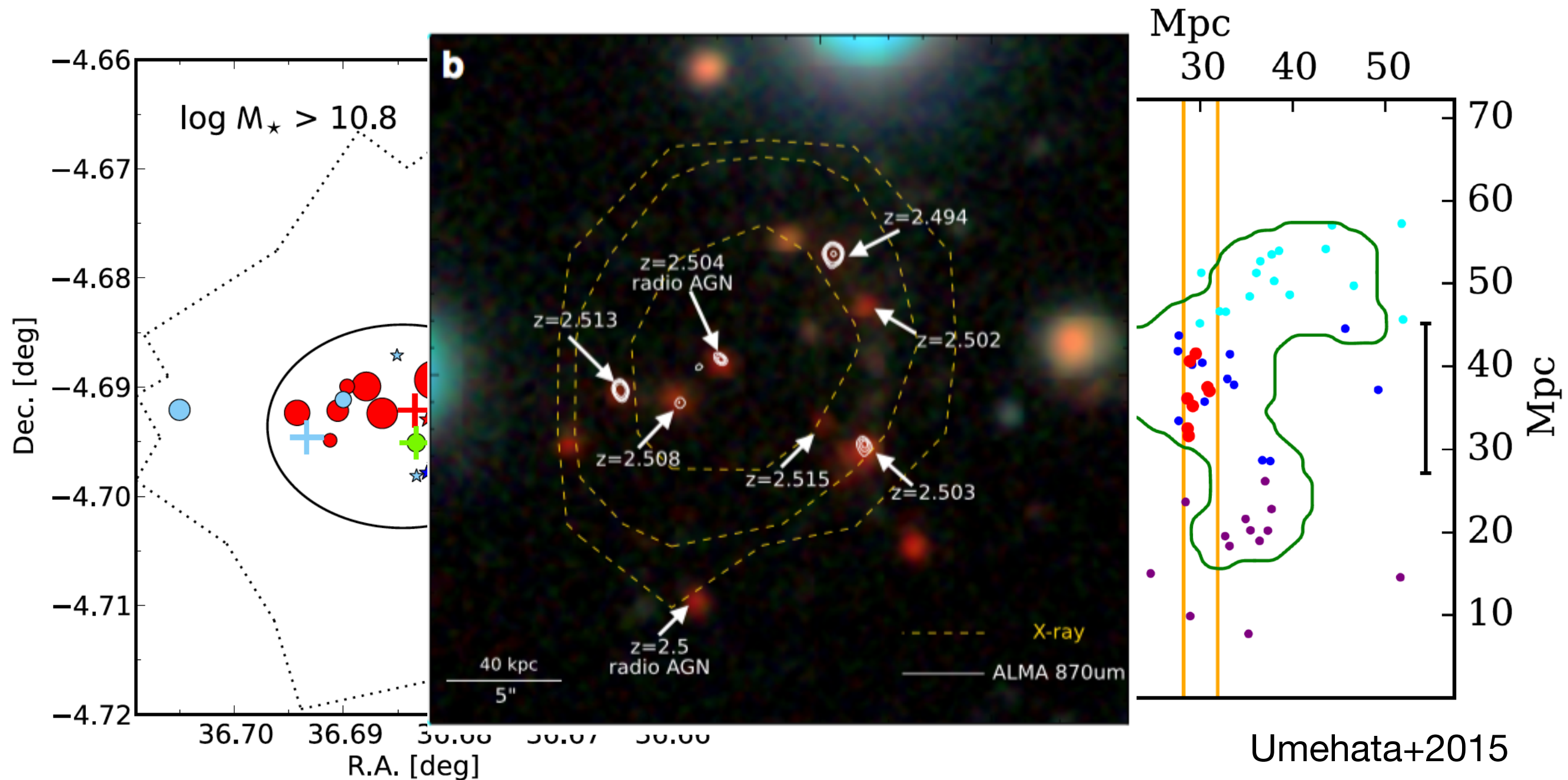


**The massive cluster galaxies are heavily obscured.**

Wang, Kodama, Tanaka, et al. in prep

**Narrow band imaging of J1001 with Subaru/MOIRCS**

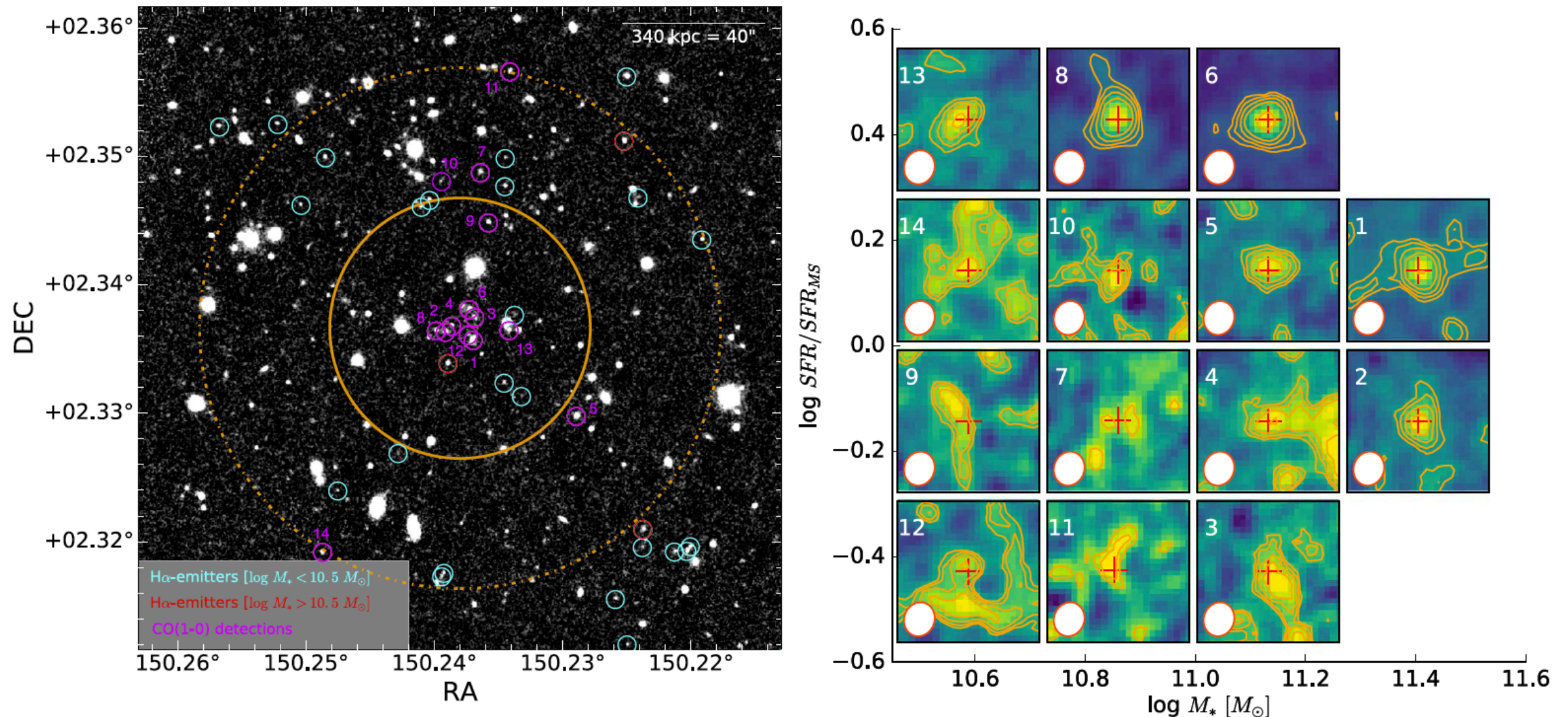
# J1001 bridges the gap from protocluster to mature clusters



**A cluster-size, virialized halo yet a dominant population of massive star-forming galaxies**



# Galaxy Properties in J1001: CO(1-0) observations with VLA



With a single pointing from VLA, we detected 11 cluster members in CO(1-0), which nearly doubles the total number of CO(1-0)-detected normal galaxies at  $z > 2$

# First evidence of environmental dependence of gas content at $z > 2$ :

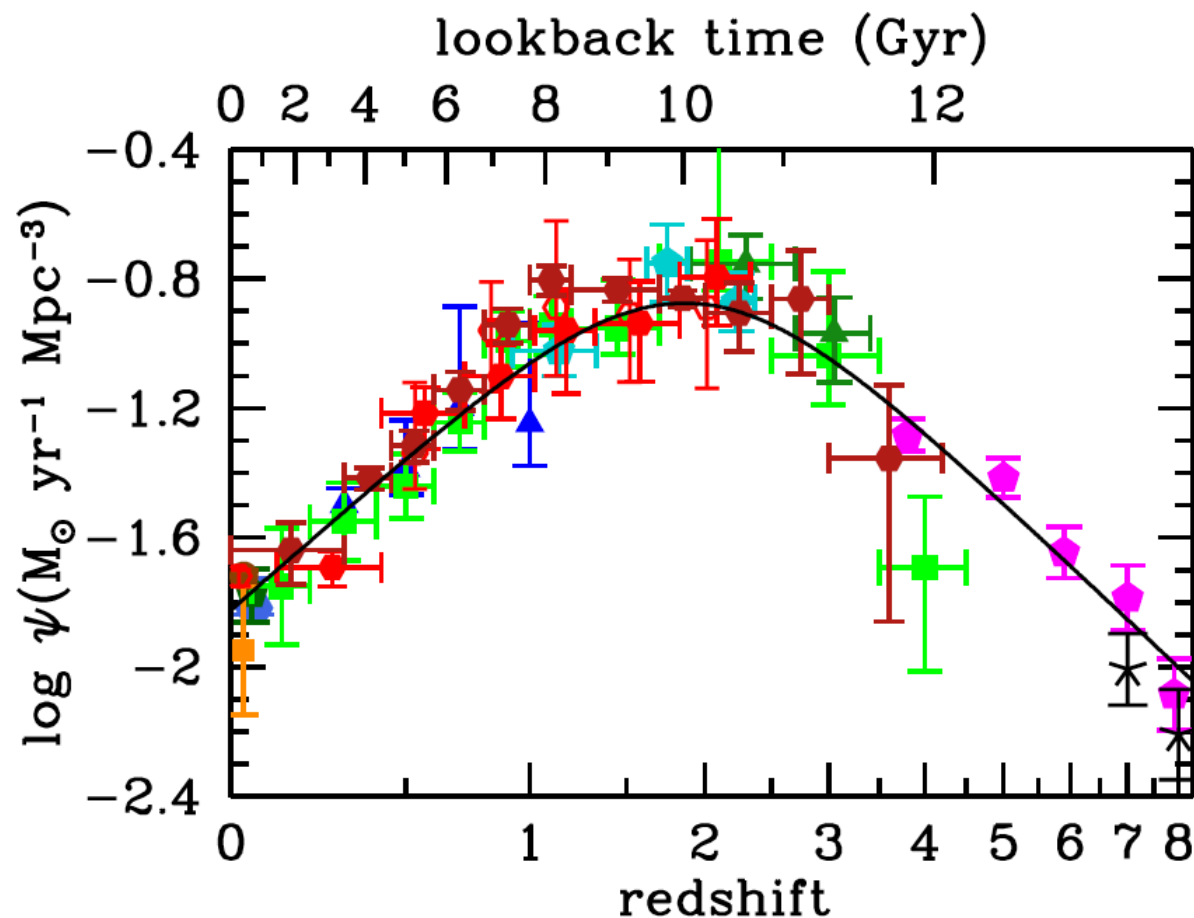
Dec



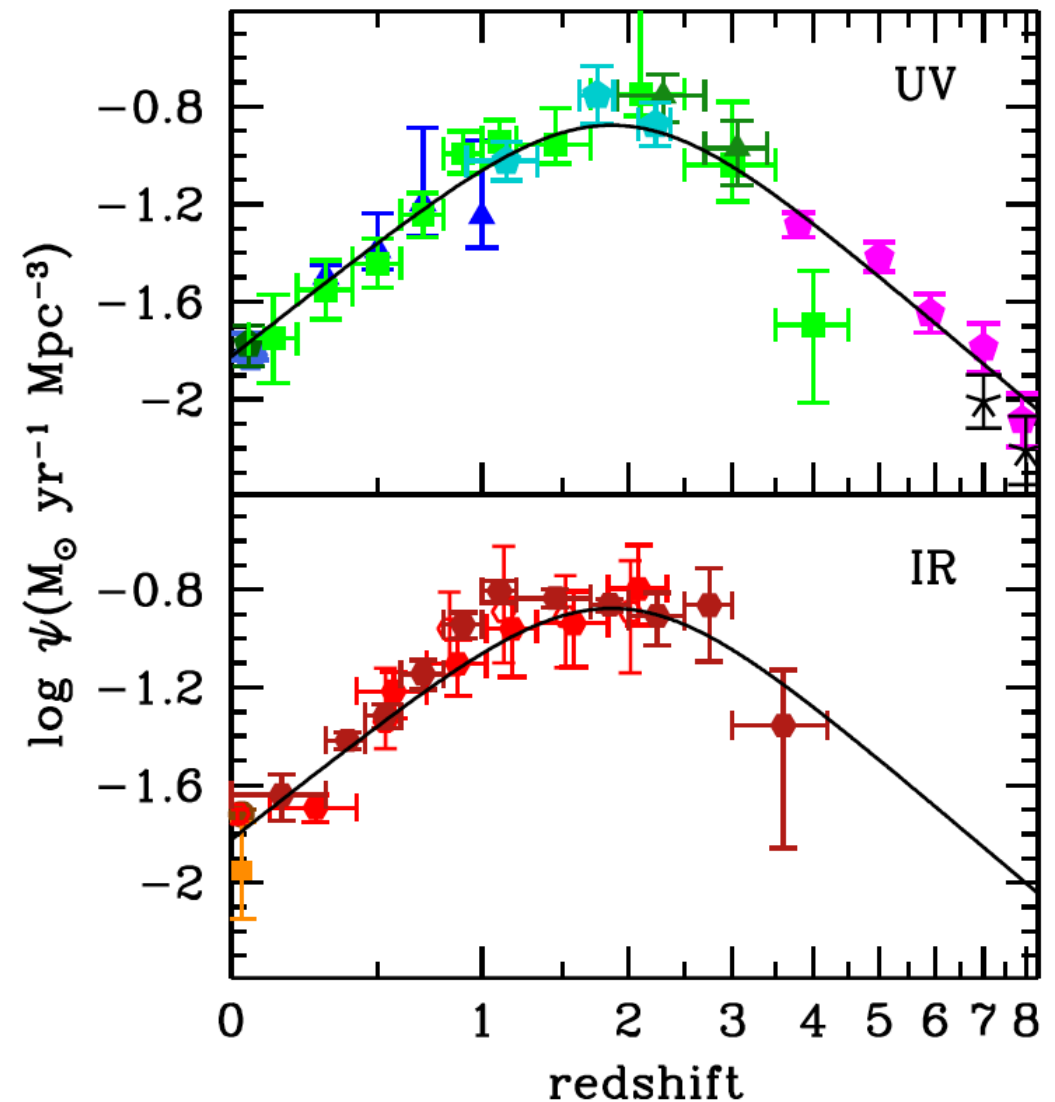
**Key: CO(1-0) observations of **mass-complete** samples of cluster members across different local environment!** also see Tadaki+2019



# Dark galaxy clusters revealed by overdensities of HST-dark galaxies at $z > 3$ :



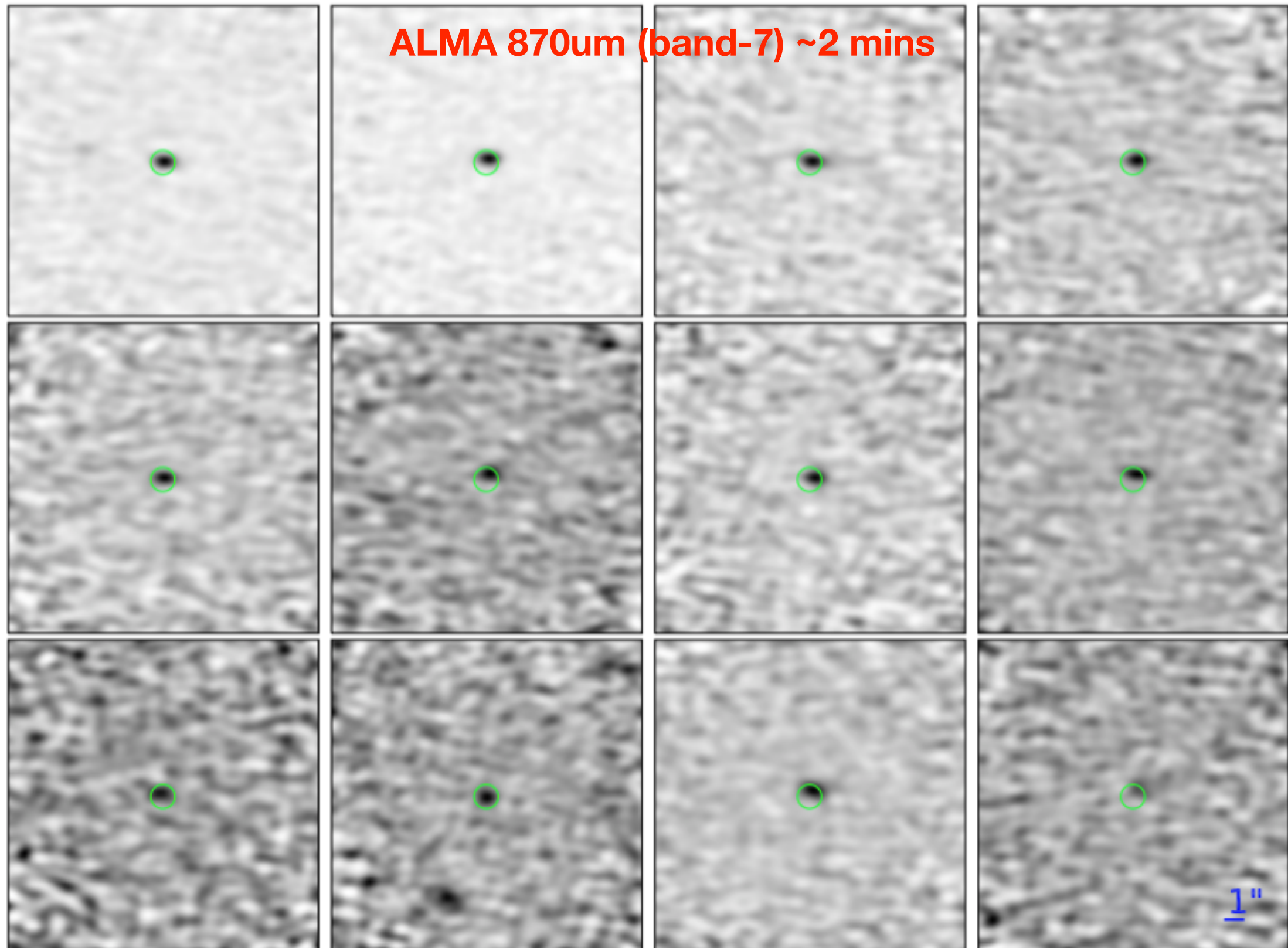
Madau & Dickinson, 2014



**Current determination of the cosmic star formation rate density at  $z > 3$  is solely based on UV-bright galaxies.**

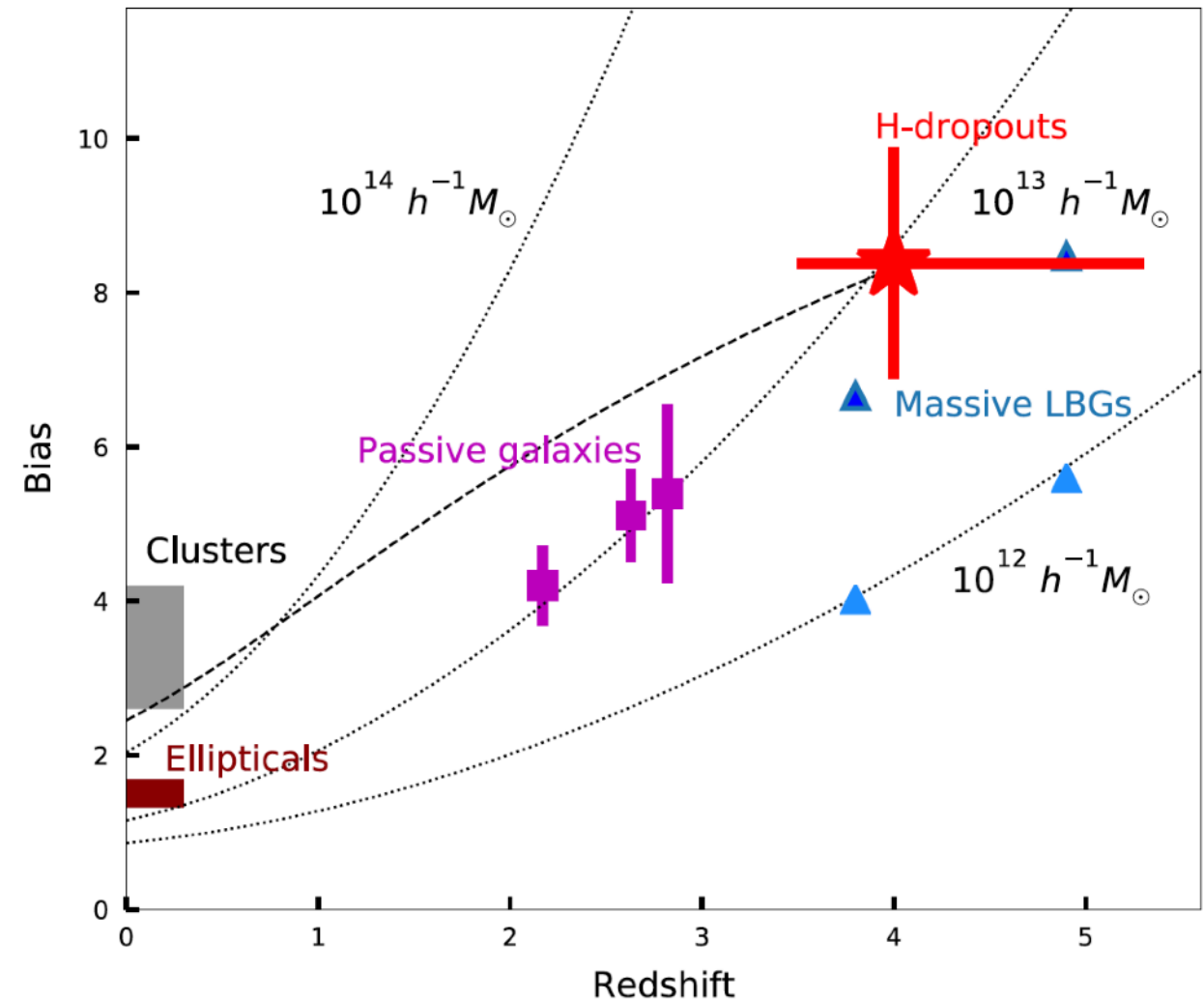
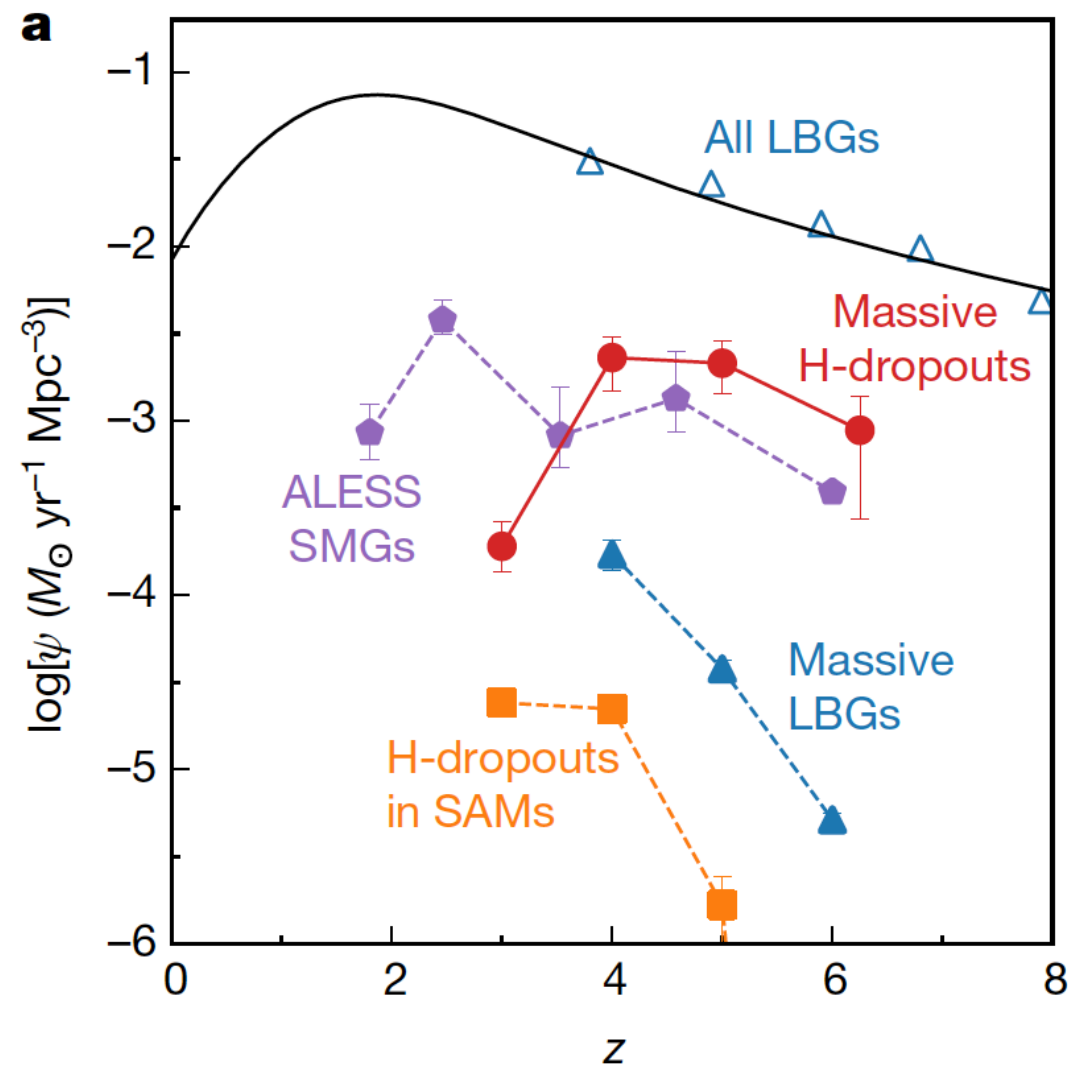
# Are the most massive $z > 4$ galaxies missed by HST?

(A complete H-dropout sample down to  $[4.5] < 24$ )



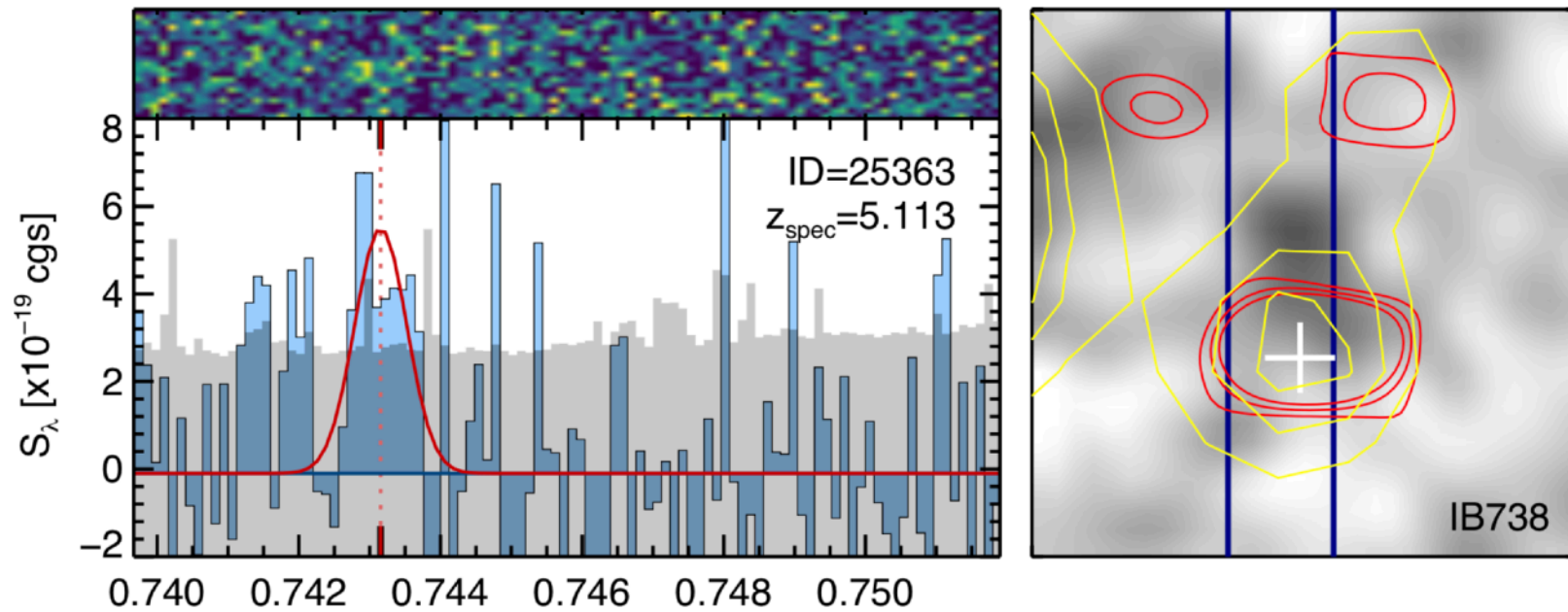


# H-dropouts dominate the cosmic SFR density from massive galaxies, and are likely progenitors of today's most massive galaxies in clusters



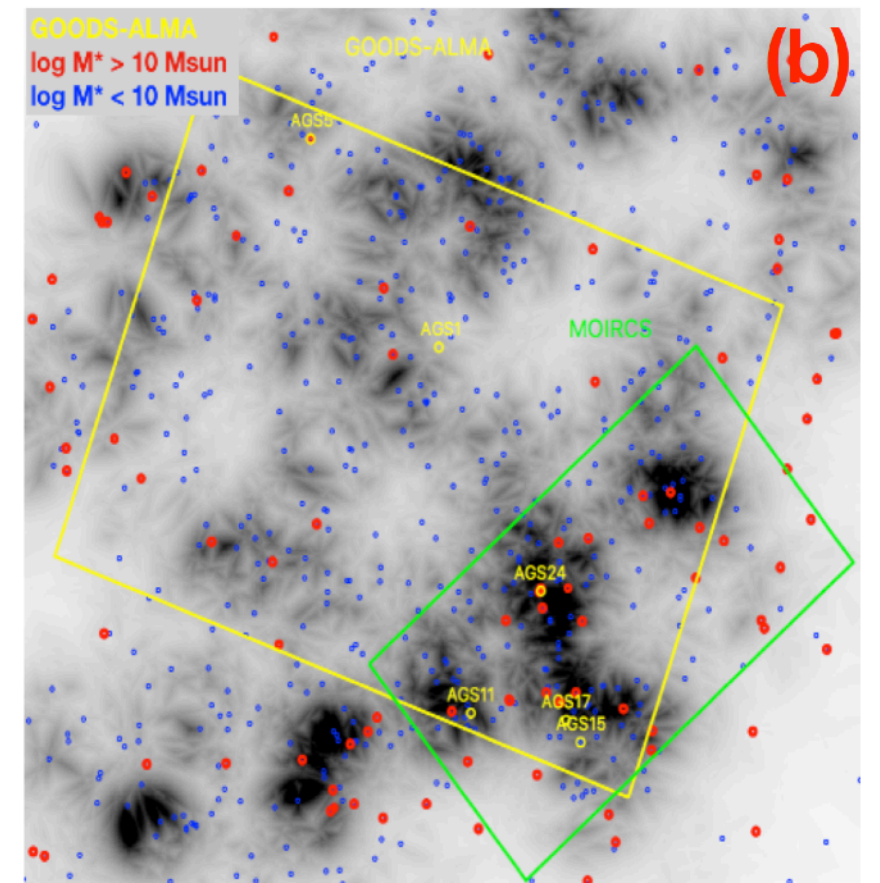
Wang,T+2019, Nature, 572, 211

# Dark galaxy clusters revealed by overdensities of HST-dark galaxies at $z > 3$ :



Subaru archival data!

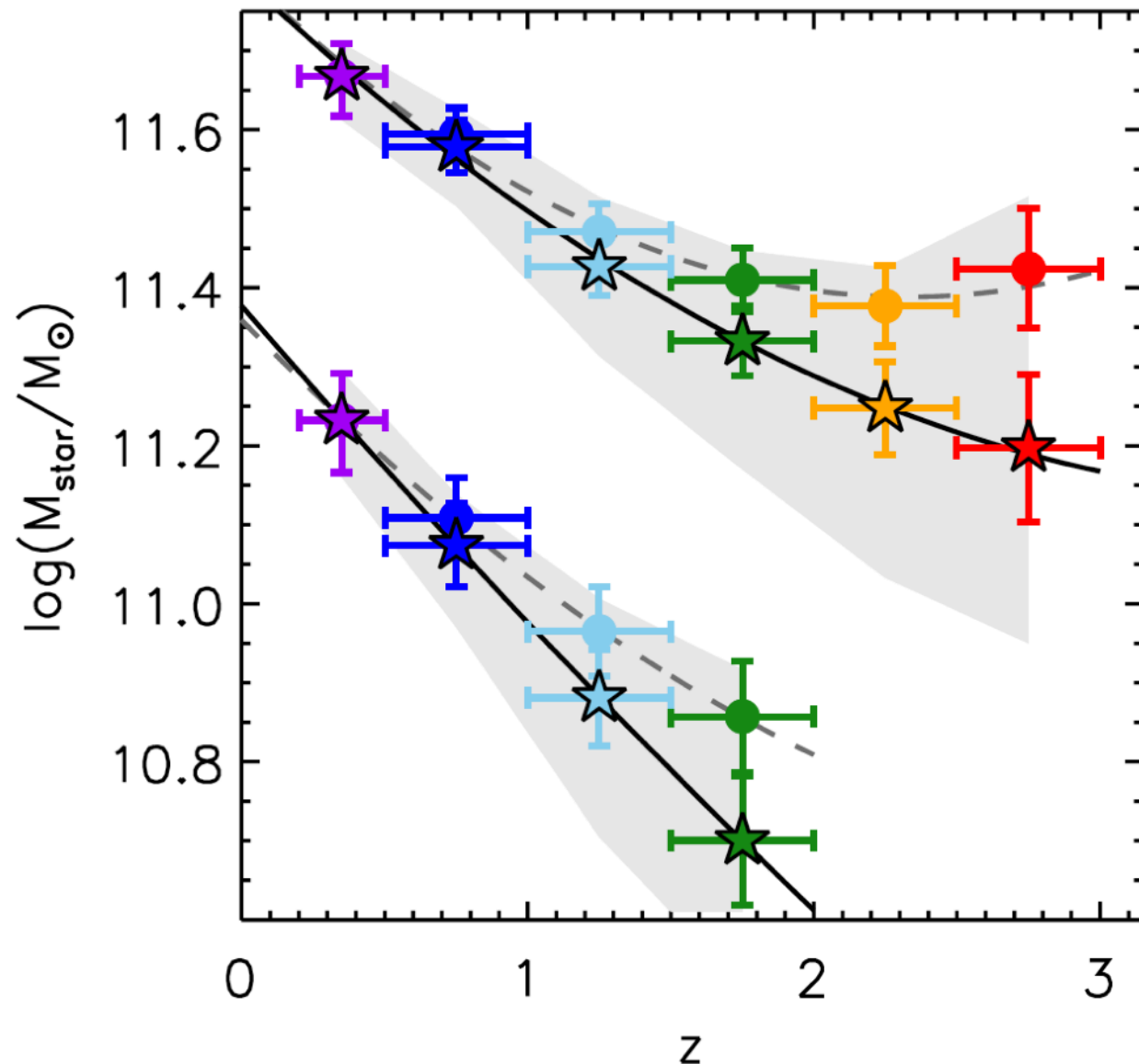
Wang+2019, 2020 (in prep)



Zhou+2019 (to be submitted)  
and GOODS-ALMA

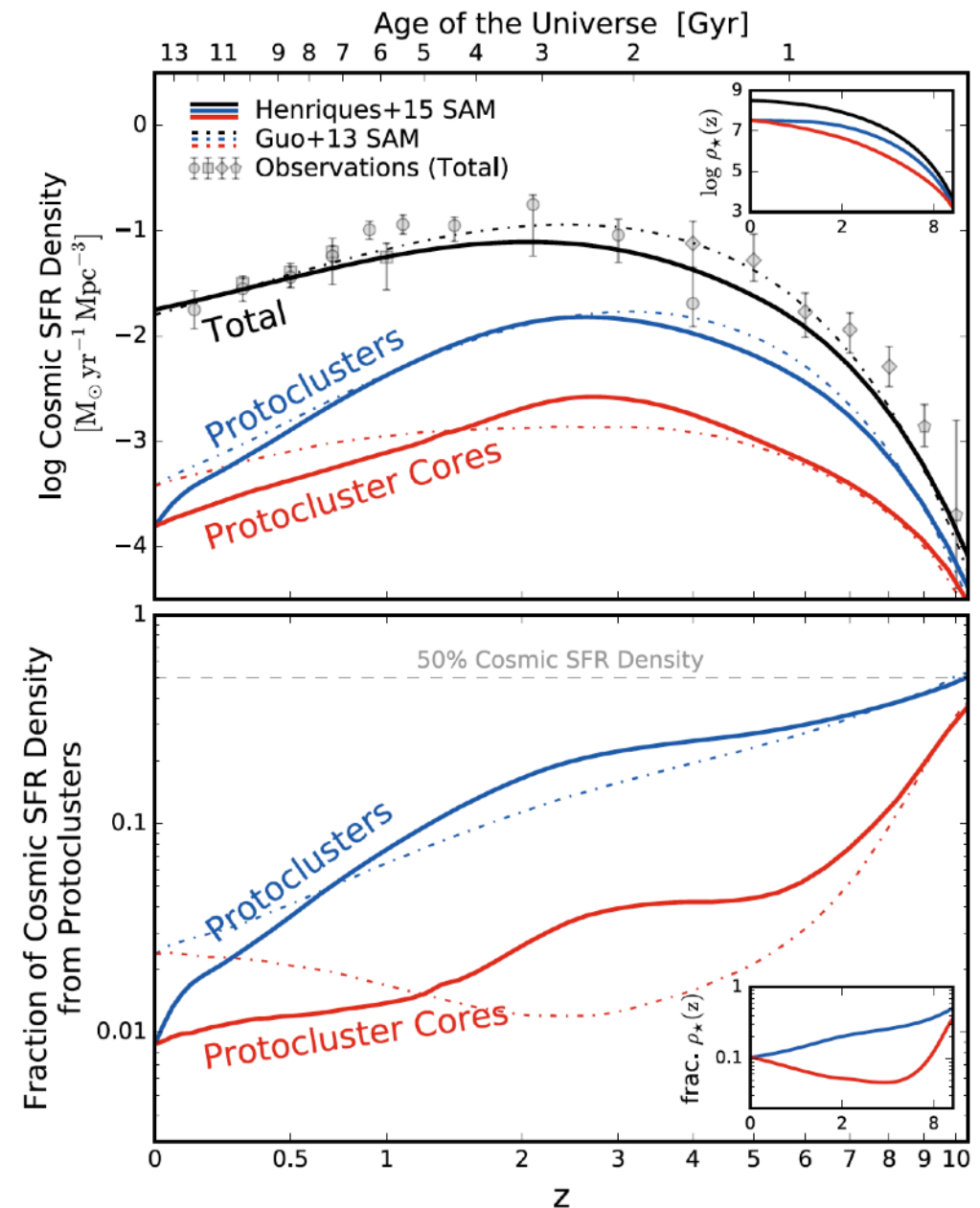


# Massive galaxies at high- $z$ are good tracers of galaxy (proto)clusters



Mass evolution of galaxies at constant number densities

Behroozi+2013



Increasing contribution from protoclusters to the cosmic SFR density at higher redshifts

Chiang+2015

# Summary

- We reveal a novel population of young clusters with a dominate population of massive star-forming galaxies in the core, which bridges the gap between mature clusters and protoclusters.
- While most of the massive cluster galaxies locate on the main-sequence, strong cluster-centric radius dependence of molecular gas content (and SFE) is revealed in J1001, providing direct evidence on environmental dependence of massive galaxy formation at  $z > 2$ .
- Subaru (MOIRCS, HSC) is critical to map the 3-D structures and obtain a complete census of member galaxies of young galaxy clusters in the early universe.