

# W. M. Keck Observatory

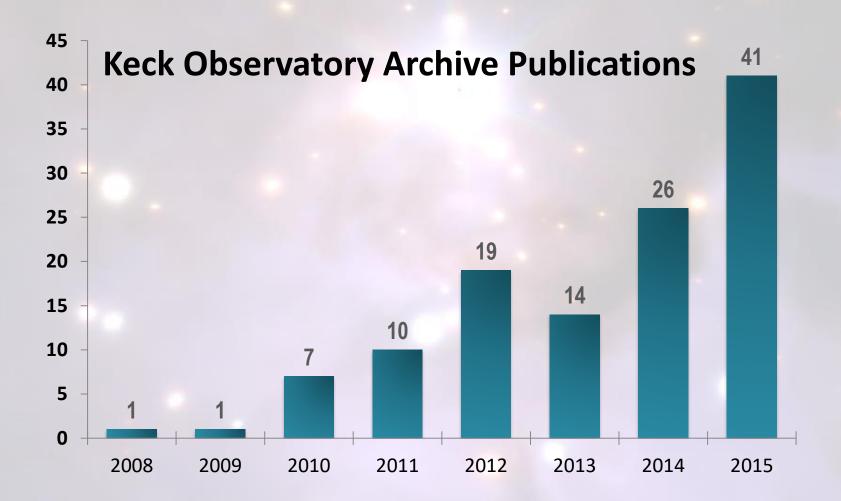


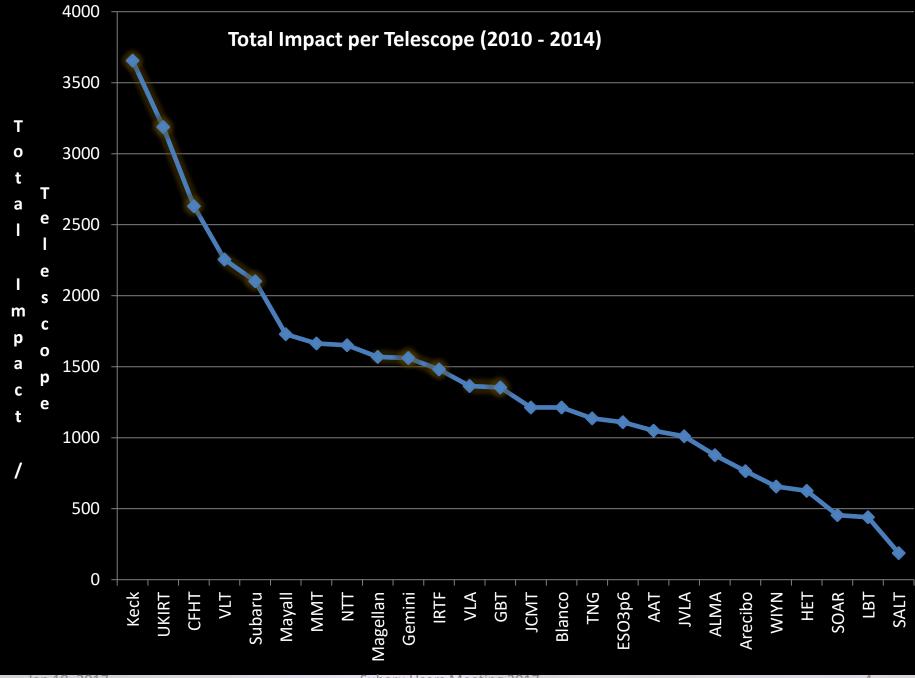
Hilton Lewis
Subaru Users Meeting
10 January 2017

Andrew Richard Hara Photography

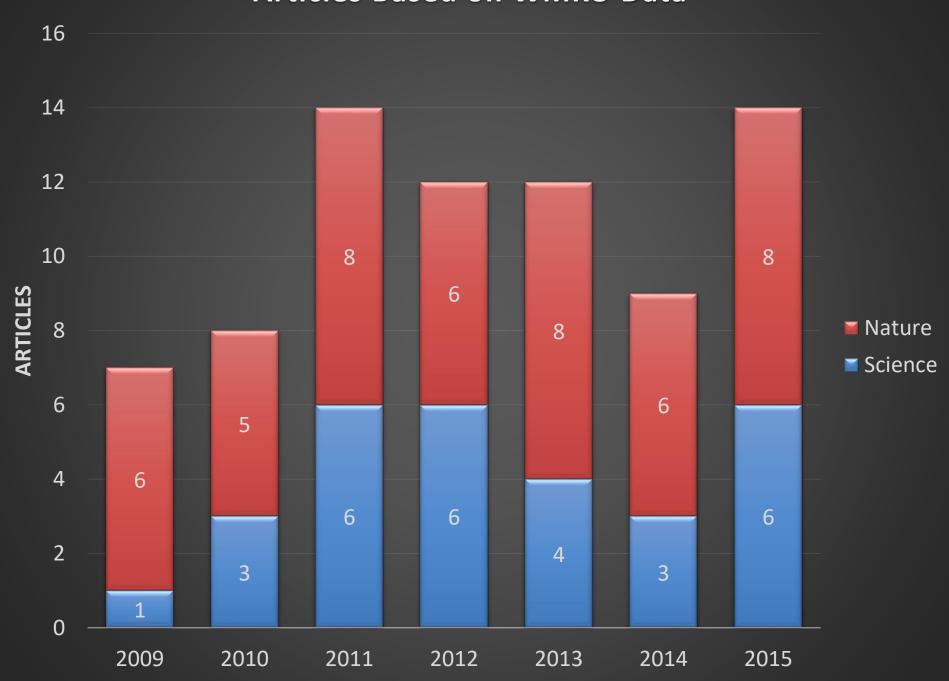
## Refereed Articles: based on WMKO data





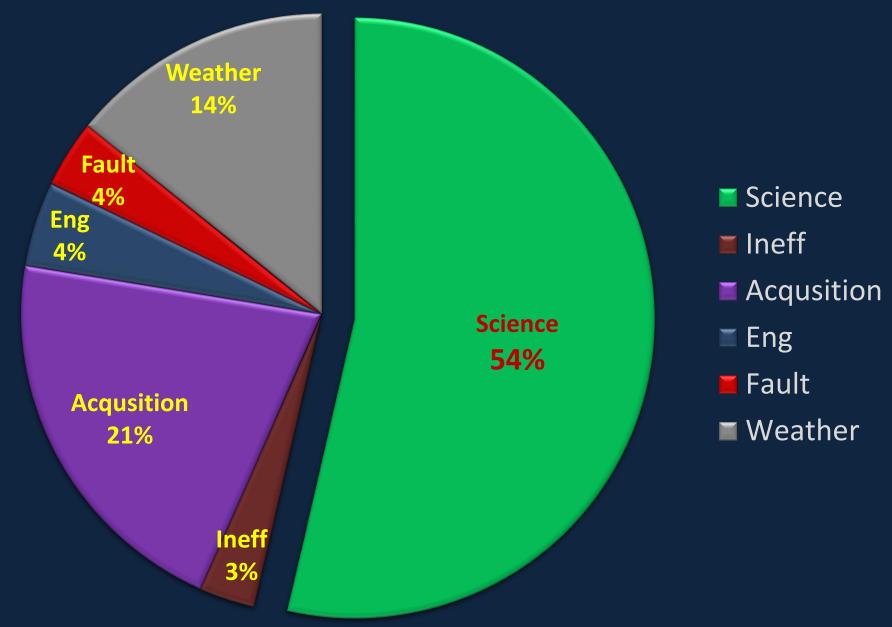


#### **Articles Based on WMKO Data**

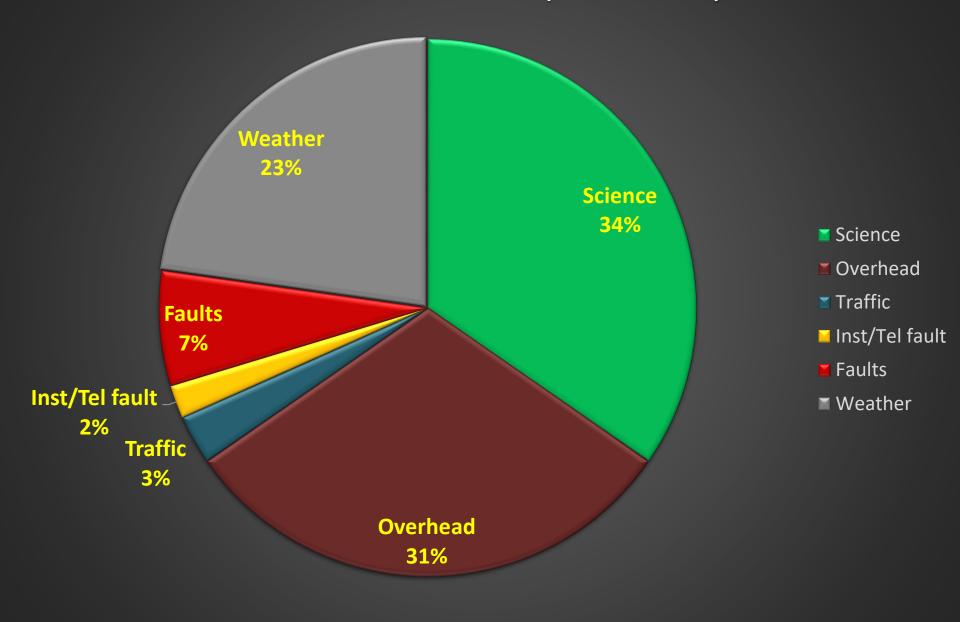


# **OPERATIONS**

#### Observing Time Breakdown: FY16, Both Telescopes



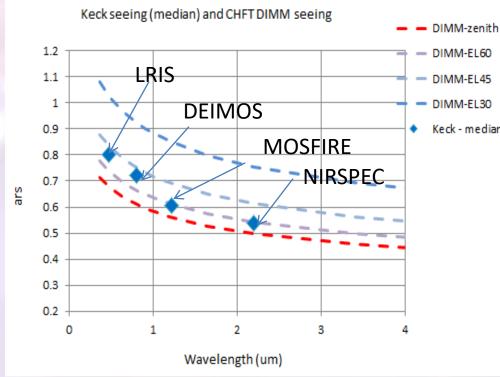
#### LGS AO Time Breakdown (Past 11 Years)



## **Excellent Dome Seeing Conditions At Keck**

- Image FWHM from MIRA vs. CFHT DIMM Seeing
- All instruments, all operating conditions, past 15.5 years
  - 0.55" peak of distribution; 0.67" median
- Median FWHM consistent with DIMM seeing from CFHT

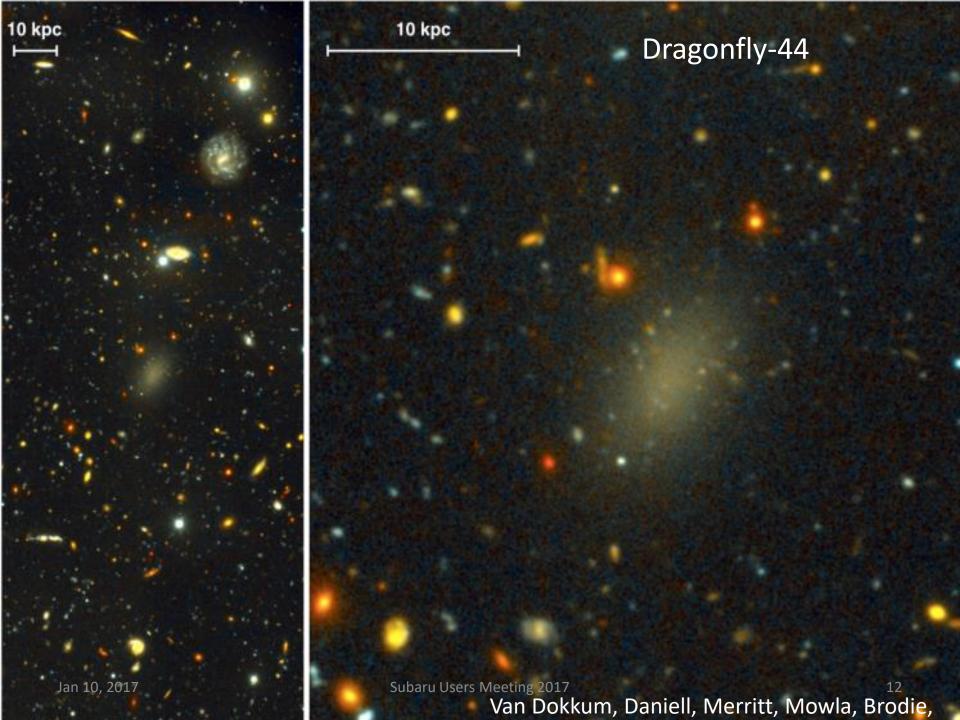




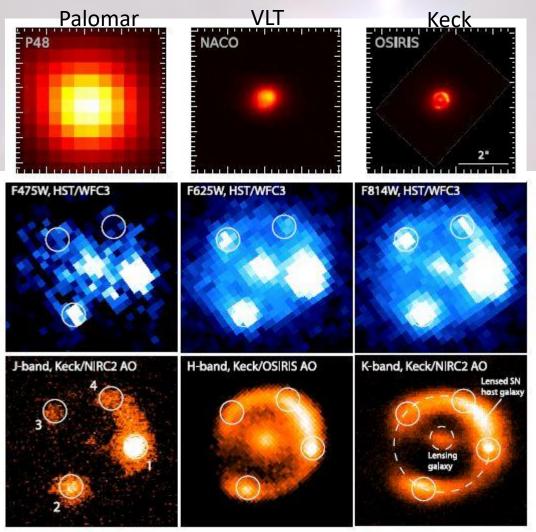
# Time Domain Astronomy

- New rules adopted for time domain astronomy scheduling to accommodate cadence and targets of opportunity
  - 2 year evaluation
- Needed for effective use of deployable tertiary
- Rules apply across the board to all partners
- Designed to minimize impact on short programs (1/2 night or smaller)





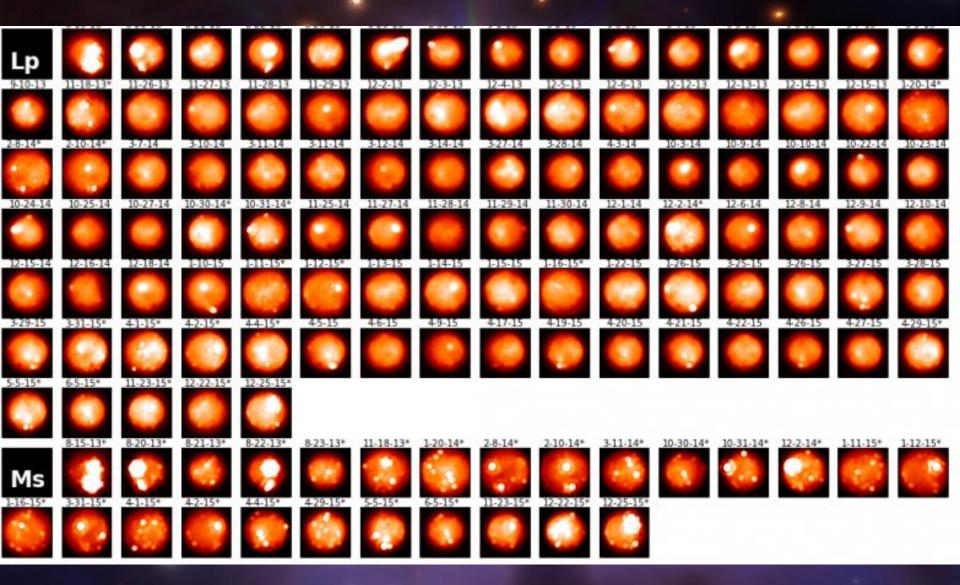
# Lensed Type la Supernova: iPTFgeu



- iPTF-discovered
  - Too bright for distance
  - Suggested lensed
- $Z_{sn} = 0.4$ ;  $Z_i = 0.2$
- Fit multi-color lightcurve of SN
  - Measure magnification
     ~4.4 magnitudes
- First time multiple images of a SN1a have been observed
- Resolved lensing galaxy
  - Small scale clumping

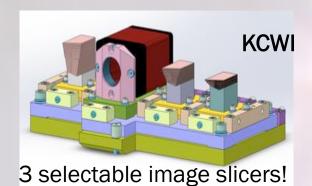
Goobar, A. et al., 2016, astro-ph

## **EVOLUTION OF VOLCANIC ERUPTIONS ON IO**



# FY16 UPDATE

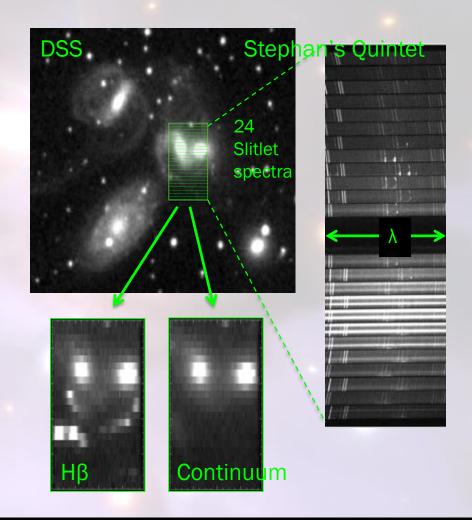
### KCWI Flexible & Powerful 2D Spectroscopy



#### **Strengths**

Jan

- Blue coverage (3500-6000Å)
- Red coverage (5600-10500Å)
- Precision Sky Subtraction
- Flexible observing modes
- High Resolution (R~20,000)



	Slicer	Field of View	<b>Spatial Resolution</b>	Spectral Resolution
	Small	20" x 8"	0.6" x 0.35"	0.23 Å (R 20,000)
	Medium	20" x 16"	0.6" x 0.7"	0.45 Å (R 10,000)
10	Large	20" x 32"	0.6" x 1.4"	0.90 Å (R 5,000)

# Keck Cosmic Web Imager-Blue

- Instrument completed, packed and ready for shipping
- Preliminary results from testing meet / exceed specifications
- Installation between Jan 2017 and Mar 2017
- Commissioning and science verification starts in Mar 2017
- Available for scheduled observing in 2018A



# Segment Repair Project Progress

- Phase II Lab delayed but now complete
- First segments transported to HQ and repair underway
- Schedule
  - Production repairs started Dec 2017 spring 2017
  - Planning to complete 27 segments by end of FY17
  - Project expected to complete late 2019





# **OSIRIS** Upgrade

- Spectrograph detector system upgraded
  - Hawaii-2 replaced with Hawaii-2RG
  - Added focus stage to optimize detector alignment
  - Replaced detector readout system
  - Updated spectrograph computer
  - Successfully commissioned and in routine service
- Preparing to commission new imager detector system in early 2017

## Telescope Control System Upgrade

#### Keck 2

- First successful science demonstration June 2016.
- Excellent pointing achieved, far superior to old system.
   Conservative estimate of time saved through improved pointing is
   5 nights/year, modeled using 1 years' worth of actual acquisitions.
- Blind on-sky performance is 1.2" RMS from 7 months' pointing tests, refitting encoder zero points only.
- Tracking closer to keyhole with improved wind-shake rejection.

#### Keck 1

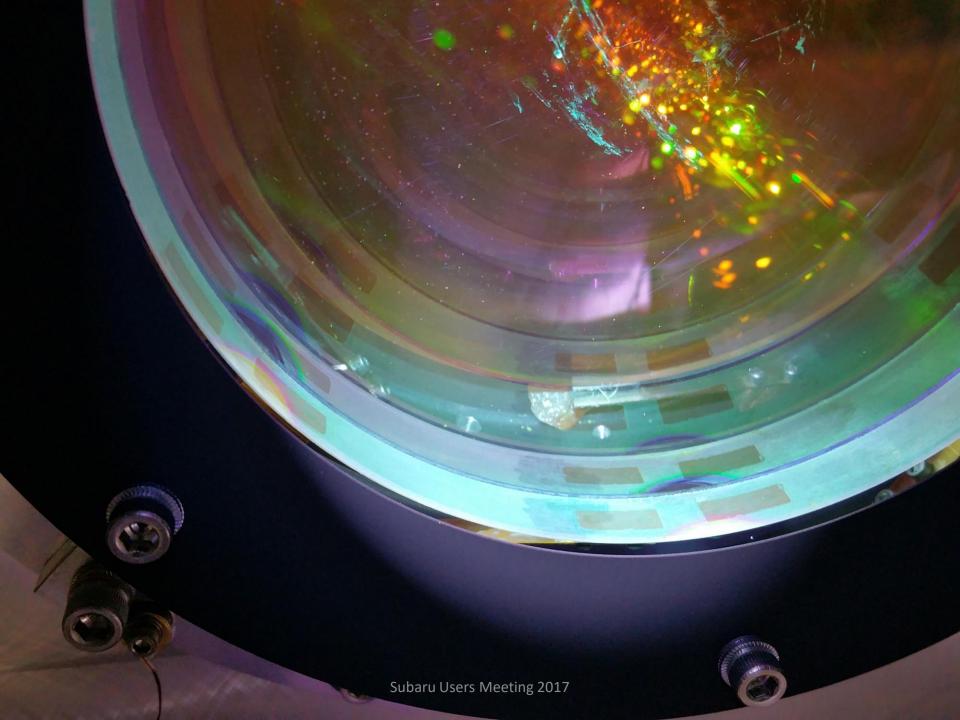
- Excellent pointing much better than old system, though not as good as K2. Suspect due to nature of azimuth journal tilt variations.
- Best model fit ~1.25" RMS

## But Then...

On the point of completion with full transition to science when rotator accident occurred, damaging MOSFIRE

- Reverted all science operations to old control system
- Reviewed rotator subsystem design and servo approach following mishap (with expert outside help)
- Implementing additional safety precautions to prevent the possibility of future incidents
- Completion date dependent on available commissioning time in 2017A





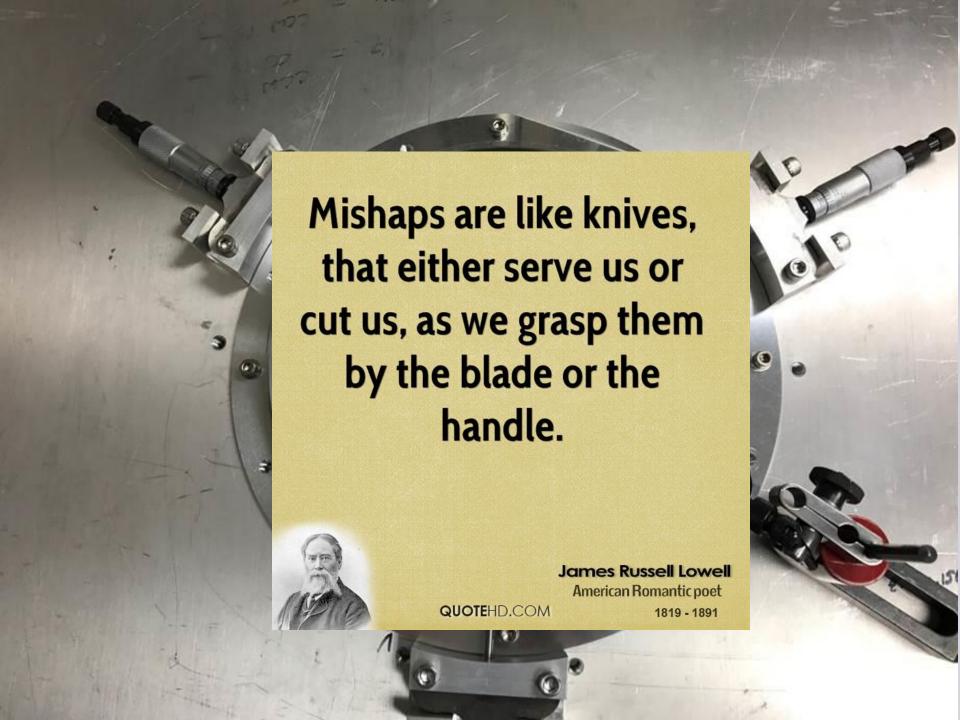






Jan 10, 2017

Subaru Users Meeting 2017



# Unattended Night-Time Operations at Summit

#### Rationale:

Cost savings, reliability improvements, efficiency gains

#### Scope:

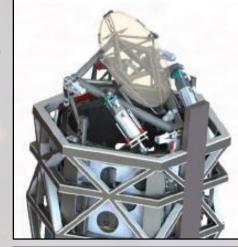
- 1/3<sup>rd</sup>: New tools to support night staff unattended operations needs
- 2/3<sup>rd</sup>: Efficiency and reliability improvements to eliminate need for summit presence for fault detection/recovery

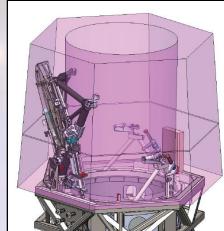
#### Milestones:

- Design reviews: Oct. 2016 (concept), May 2017 (preliminary) and Sep. 2017 (final)
- Efficiency and reliability improvements to existing equipment starting in 2017 and continuing through 2019
- Start of fully unattended operations no sooner than 2019

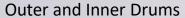
## Keck I Deployable Tertiary (K1DM3)

- Mechanized Tertiary that can be swung in and out in place
- Rapid switch between Nas and Cass instruments (~120 sec)
- Module will remain inside the Tertiary Tower
  - Module and Mirror will be removed for recoating in similar manner as existing tertiary
- Installation and Commissioning scheduled for August 2017
  - Module will be shipped here between June and August
- Key to enabling time-domain astronomy on Keck 1 telescope











Swingarm



Mirror Support Flexure



Whiffle Tree and Dummy Mirror

## **Keck II Laser Status**

- Upgraded old dye laser in use since 2002 to new Toptica fiber laser
- LGS AO science resumed April 2016.
- ~10x signal return but using just 5% input power (3kW).
- Project completed on schedule and under budget!

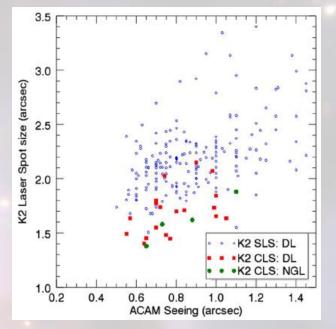


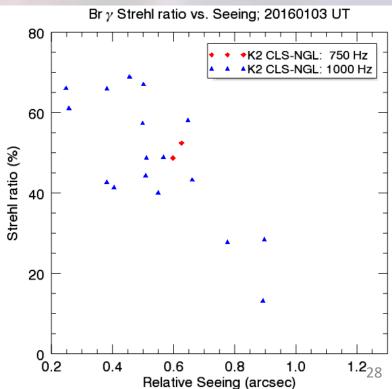




## **K2** Laser Status

- Predicted laser return achieved: R = 7.5 (19x old dye laser)
- Laser spot size ~24% reduced
- Best Strehl ~65% vs. ~53% with dye laser
  - Performance limited by wavefront controller bandwidth; no noticeable improvements operating between 14W and 20W





## **DEIMOS Service Mission II**

- Completed March 2016
- No belt breaks since then
- Can clamp gratings at any rotator angle
- Flexure reduced to historically "good" levels
- No time lost due to mechanical failures since repair. Previously, 50% of DEIMOS problems were due to the grating system

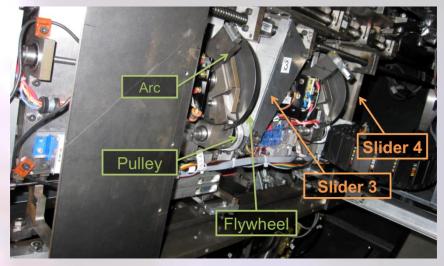


Photo of the DEIMOS grating sub system. Parts labeled green were replaced. There is one set of components for each slider.



# Segment Warping

- Direct slope algorithm for determining warping forces is fully operational
  - Manually tested during 2015 K1 exchange
  - Fully implemented for 2016 K2 exchange
    - Matlab warping GUI, direct interface to PEAS-PCS database, & a new warping fixture with a Python GUI
- Result:  $\theta_{80} = 0.32''$  for both exchanges
  - 15% improvement over previous years



# Programs in Progress

Status	First light	Shared Risk
Testing on sky	June 2015	2017A
Facility rotator re-commissioning	Apr 2015	2017A
Preparing for integration	Jan 2017	2017B
Delivery to observatory	Jan 2017	2018A
Final I&T in lab	Oct 2017	2018A
Detailed design	Jul 2017	2018A
Preliminary design	Feb 2018	2019A
Preliminary design	Mar 2018	2019A
Preliminary design	Mar 2019	2019B
Preliminary design	June 2019	2019B
F C F F	Testing on sky Facility rotator re-commissioning Preparing for integration Delivery to observatory Final I&T in lab Detailed design Preliminary design Preliminary design Preliminary design	Festing on sky  Facility rotator re-commissioning  Preparing for integration  Delivery to observatory  Final I&T in lab  Detailed design  Preliminary design

Oct 2019

2020A

Preliminary design

**Keck Planet Finder** 

## Major Deliverables in FY17

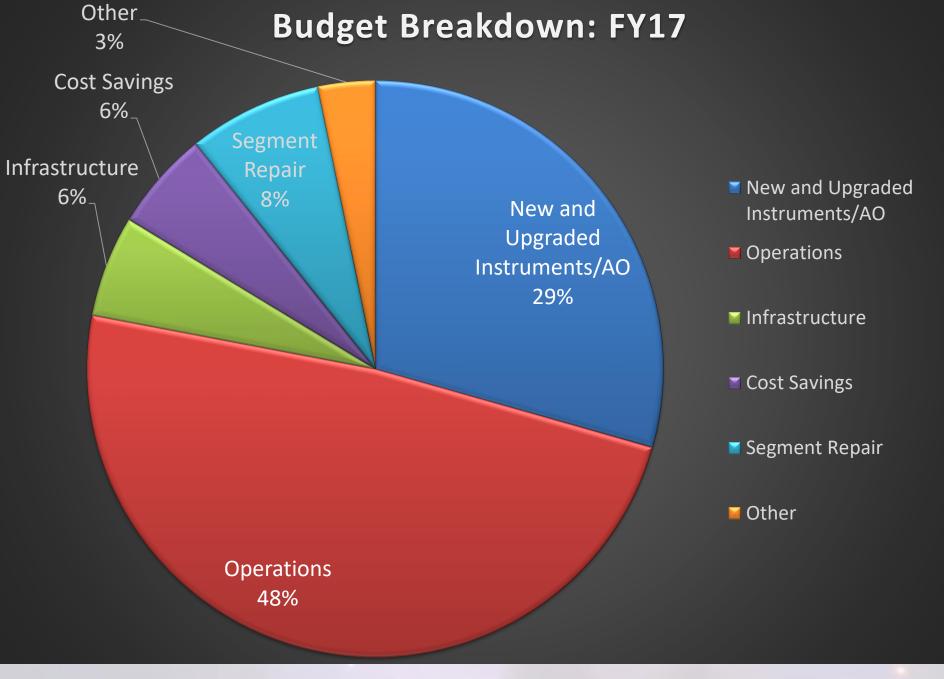
Month	Keck I		Keck I & II	Keck II	
Oct					
Nov	MOSFIRE Repair				
Dec					
Jan	MOSFIRE			KCWI Install	
Feb	Recommissioning	OSIRIS Imager Install		KCVVI IIIStali	
Mar		0010101	TCSU		
Apr	Segment Exchange	OSIRIS Imager Commissioning	Commissioning	KCWI	PCS Camera
May		Commissioning		Commissioning	Upgrade
Jun					ACS Upgrade
Jul	Segment Exchange	K1DM3 Install/Test			Install & Test
		K1DM3		NIRES Install &	ACC Ungrado
Aug	Segment Exchange	Commissioned		Daytime Testing	ACS Upgrade Install & Test
Sep					ilistali & lest
Oct	Segment Exchange			NIRSPEC Upgrade Ins	stall, DEIMOS Grating
Nov				Install, NIRES (	Commissioning

Observing Semester

FY Boundary

## FY17 Highest Priority Projects and 'Must Wins'

Prio	Project	FY17 Deliverables
1	MOSFIRE	Recommission for 17A science & complete Investigation
2	Segment Repair	Complete first 27 segments
3	TCSU	Commission tertiary by Feb. & all before 17B
4	KCWI-B	Commission for 17B science
5	OSIRIS Imager Upgrade	Commission for 17B science
6	K1DM3 Deployable Tertiary	Commission for 17B science
7	KCWI-R	Complete delta PDR + good progress on DDR
8	OSIRIS Spectrograph DRP Fix	DRP release that reduces legacy & current data
9	PCS camera Upgrade	New K2 camera & motor controller operational
10	Un-attended Night Ops	ROSI PDR & start biggest reliability subprojects
11	Spare Secondary	Complete replan based on bids. Place polishing contract
12	Summit PV	Resolve insurance issue & install if approved
13	Enhanced Tip-Tilt Sensor	Commission TRICK for 17A science & E-TRICK for 17B science
14	NIRSPEC Upgrade	Complete design review. Procure & test both detectors at UCLA
15	ACAM Guider Upgrade	Commission for 17B science
16	K2 Laser (NGL)	Service contract in place. Complete spares procurement
17	ACS Upgrade	Final design. In final testing after replacement of all nodeboxes on 1 tel.



# Maunakea Update

- Focus: Develop trusting relationships, educating leaders about astronomy's role in Hawaii and to building astronomy's image
- Message:
  - We seek to engage in a helpful, constructive manner towards a positive future for Astronomy that benefits Hawaii
- Some important findings so far:
  - There is a still a lack of basic knowledge about the astronomy sector in Hawaii among some decision makers
  - The messenger and the approach are as important, sometimes more important, than the message itself.
- Two themes resonate most:
  - Educational and work opportunities for future generations
  - Explaining the advantages for the community

