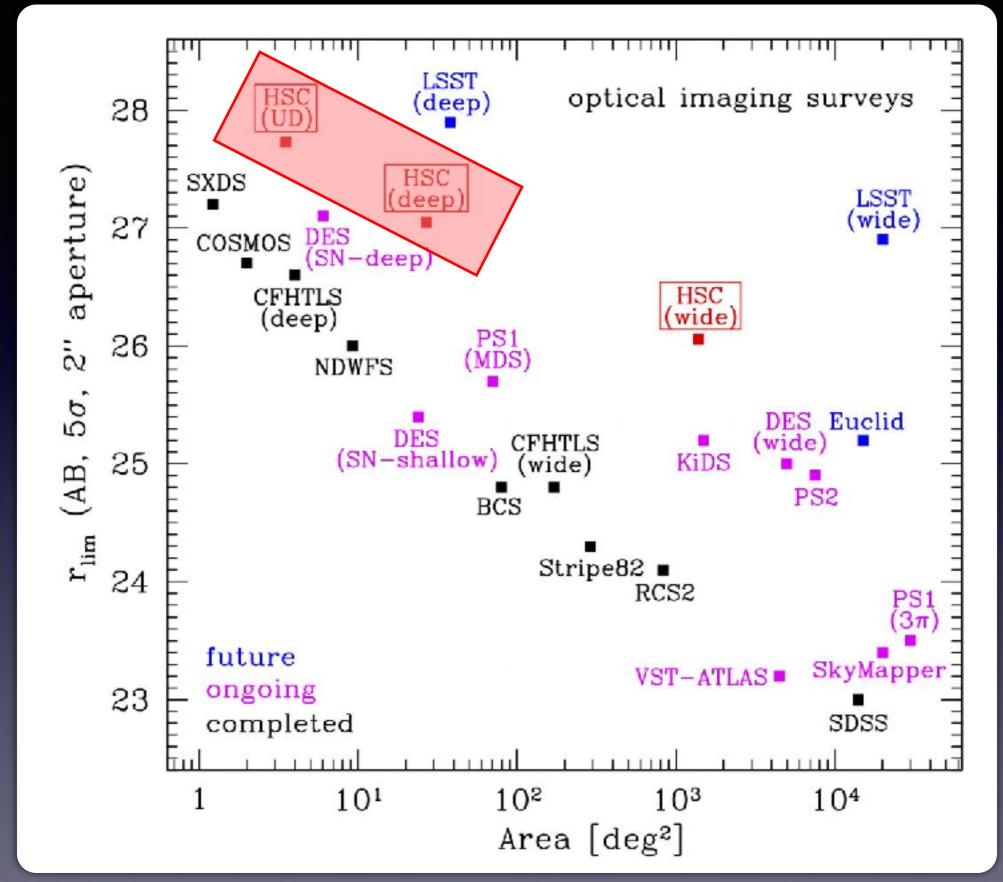
Masaomi Tanaka (NAOJ)

- Transient survey with Subaru/HSC
- Multi-messenger astronomy
- Future



http://hsc.mtk.nao.ac.jp/ssp/science/weak-lensing-cosmology/

#### Masahiro's talk

# SSP Transient survey (2016 Nov-, COSMOS)

HSC transient working group:

Naoki Yasuda, Nozomu Tominaga, Tomoki Morokuma, Nao Suzuki, Ichiro Takahashi, Takashi Moriya, Keiichi Maeda, Masaki Yamaguchi, et al.

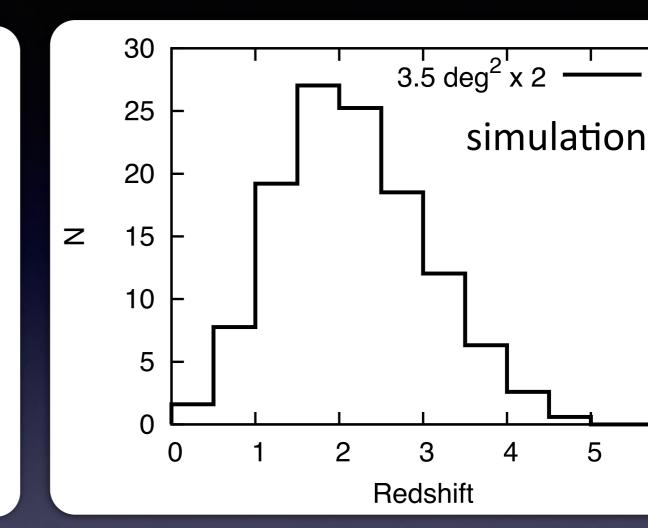
The deepest/widest transient survey as ever

Today

### Type Ia SN cosmology

#### 0.0 simulation SCP Union2 + HSC -0.4

## High-z "superluminous" SNe



- Tidal disruption => Evolution of SMBH

0.4

- SN rate => Success rate of SN/Progenitor system
- New types of transients

0.3

0.1

0.2

-0.6

-0.8

-1.0

-1.2

-1.4

0.0

# Superluminous supernovae

Collaboration with Jeff Cooke (Swinburne) led by Takashi Moriya et al.

Keck 2 nights (2017 Feb, Mar)



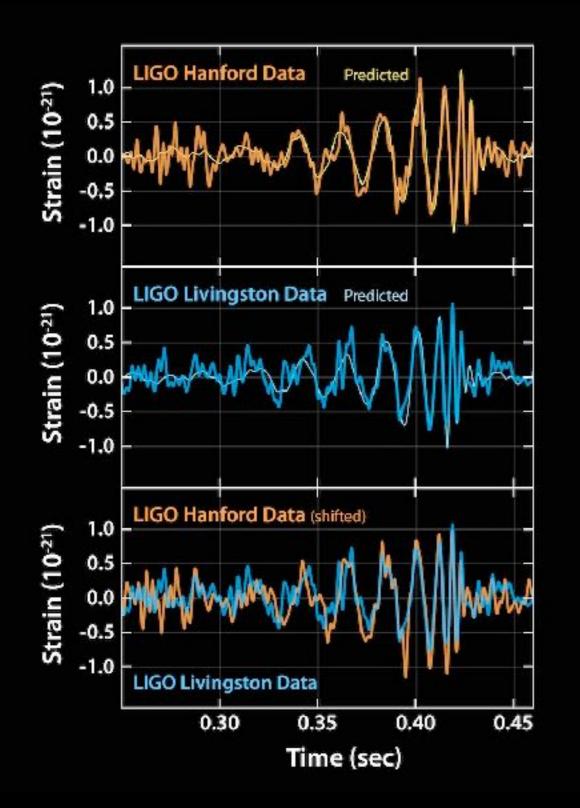
see Jeff's talk

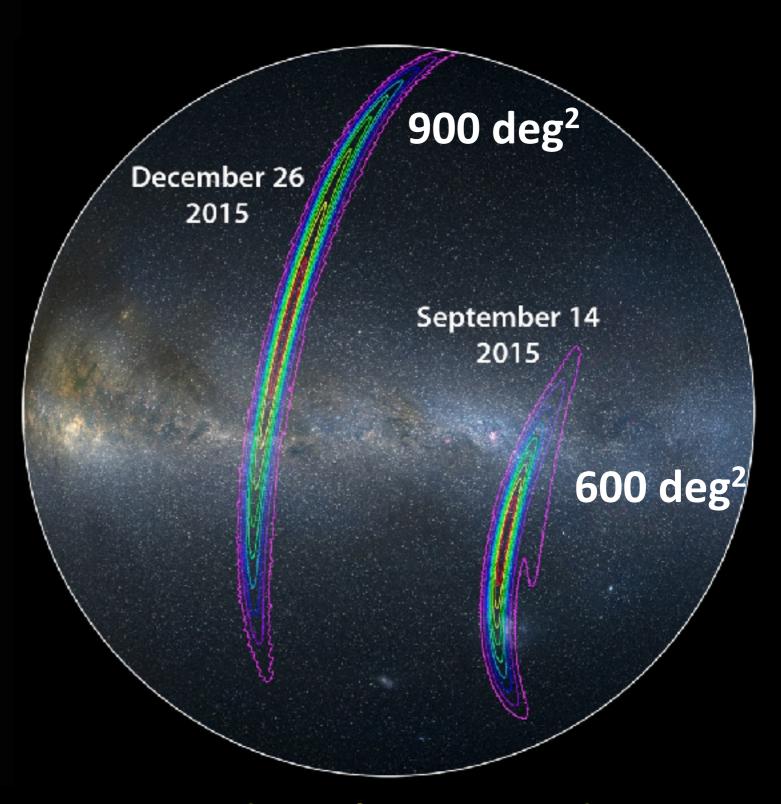
Event rate of 
"superluminous" supernova 
up to z ~ 4 
==> star formation and IMF

- Transient survey with Subaru/HSC
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#### **Gravitational wave astronomy**

Abbott et al. 2016, PRL, 061102

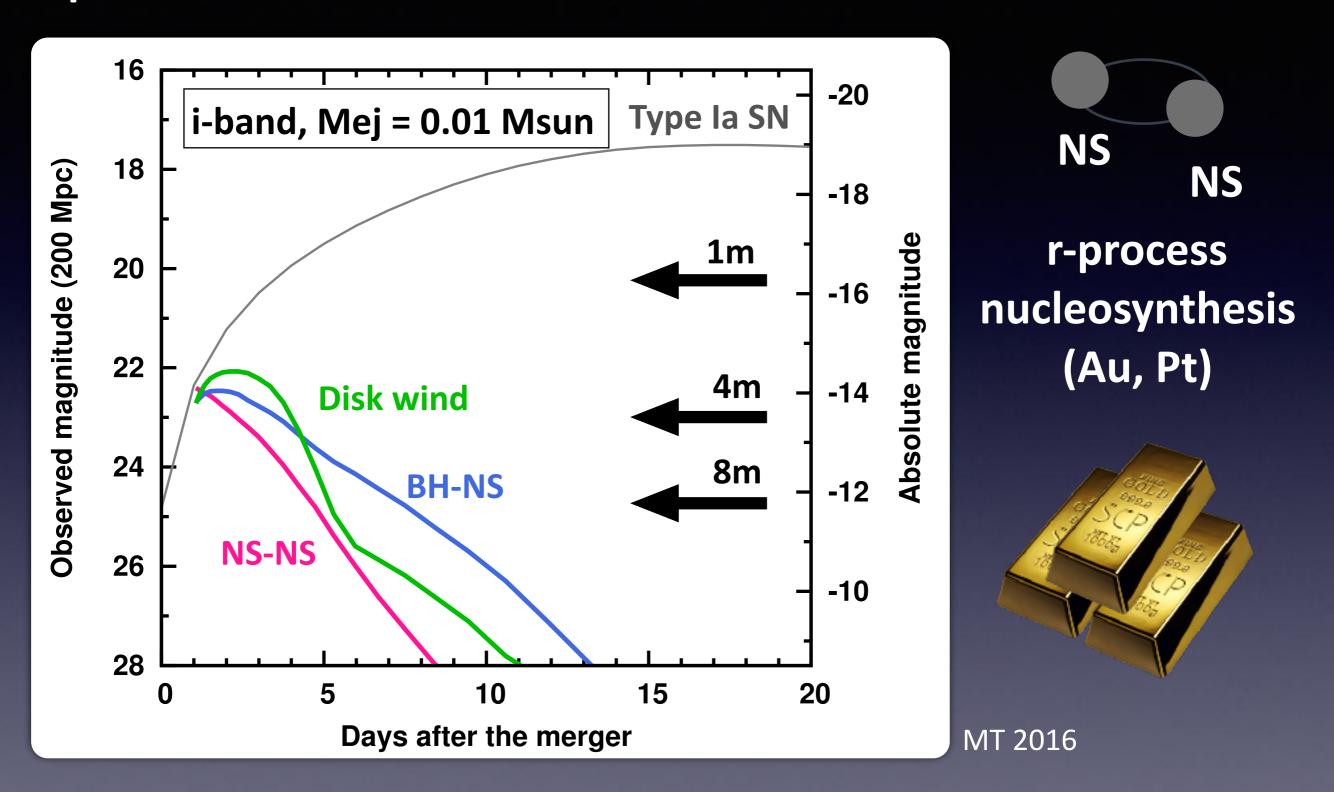




**GW150914** = **BH-BH** merger

=> 10-100 deg2 after Virgo and KAGRA

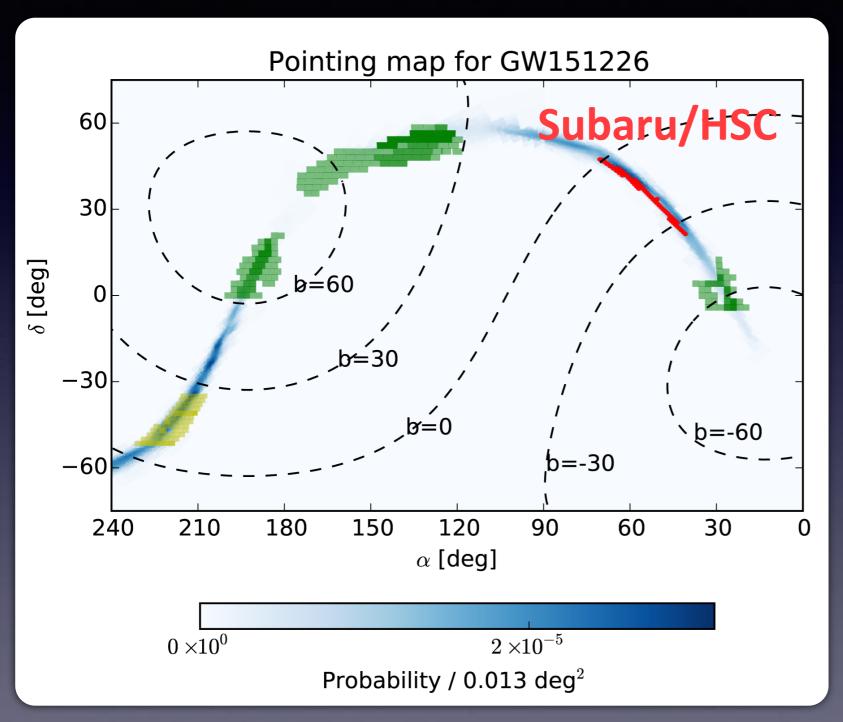
#### Optical emission from GW sources: "kilonova"



Detection of the counterpart => Ejection of r-process elements
Subaru/HSC is the unique instrument for GW follow-up

#### **ToO transient search for GW151226 (BH-BH)**

J-GEM (Japanese collaboration for GW-EM follow-up: Yoshida et al.)



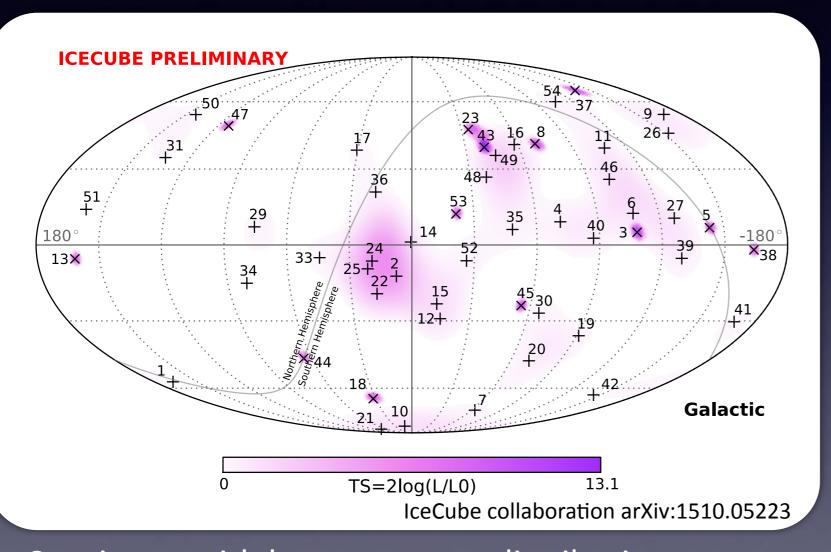
64 deg<sup>2</sup> (7%) in 0.5 night!

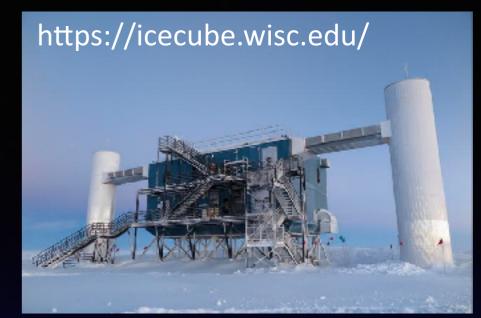
Yoshida, Utsumi, Tominaga et al. 2017, PASJ, 69, 9

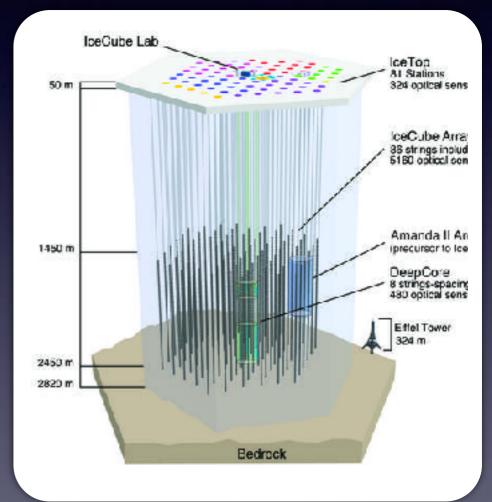
HSC: 1256 candidates and ~60 likely extragalactic transients => Need spectroscopy for the smoking gun

#### More opportunities for multi-messenger

# IceCube: high-E (TeV-PeV) neutrino Real-time alert (2016-)







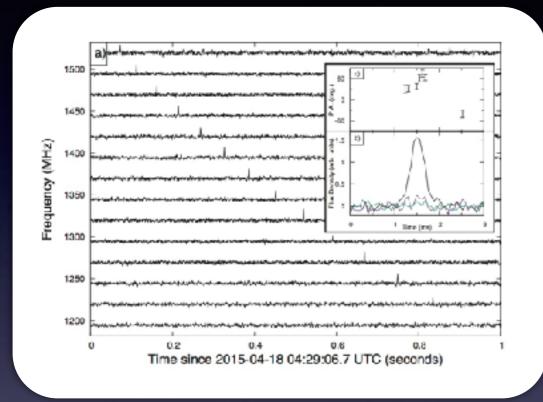
Consistent with homogeneous distribution

No correlation with TeV gamma-ray sources/GRBs

Localization ~ 1 deg => Subaru/HSC

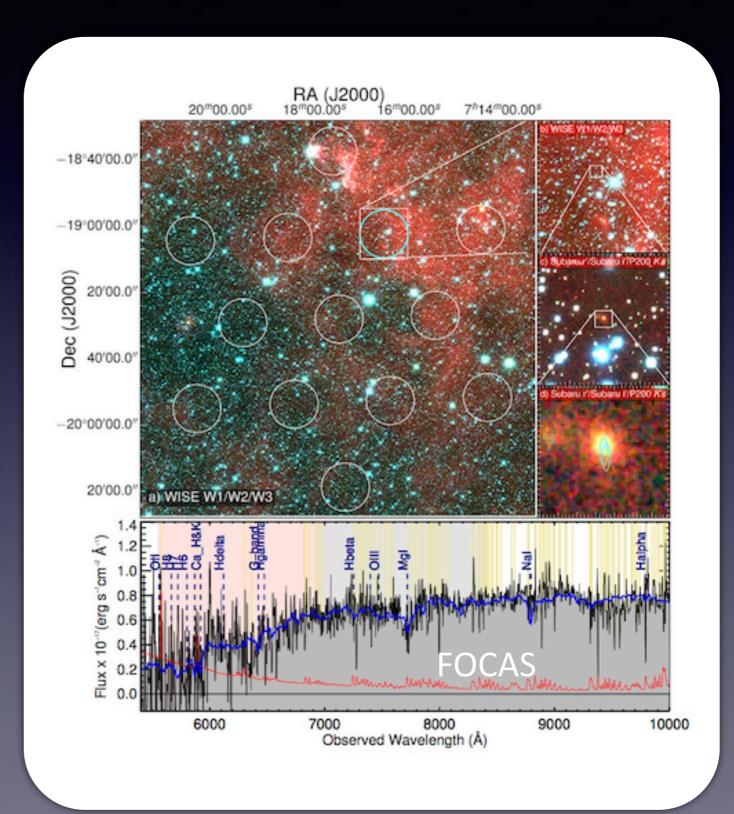
## Fast radio bursts: ToO transient search + follow-up

led by Tomonori Totani et al.



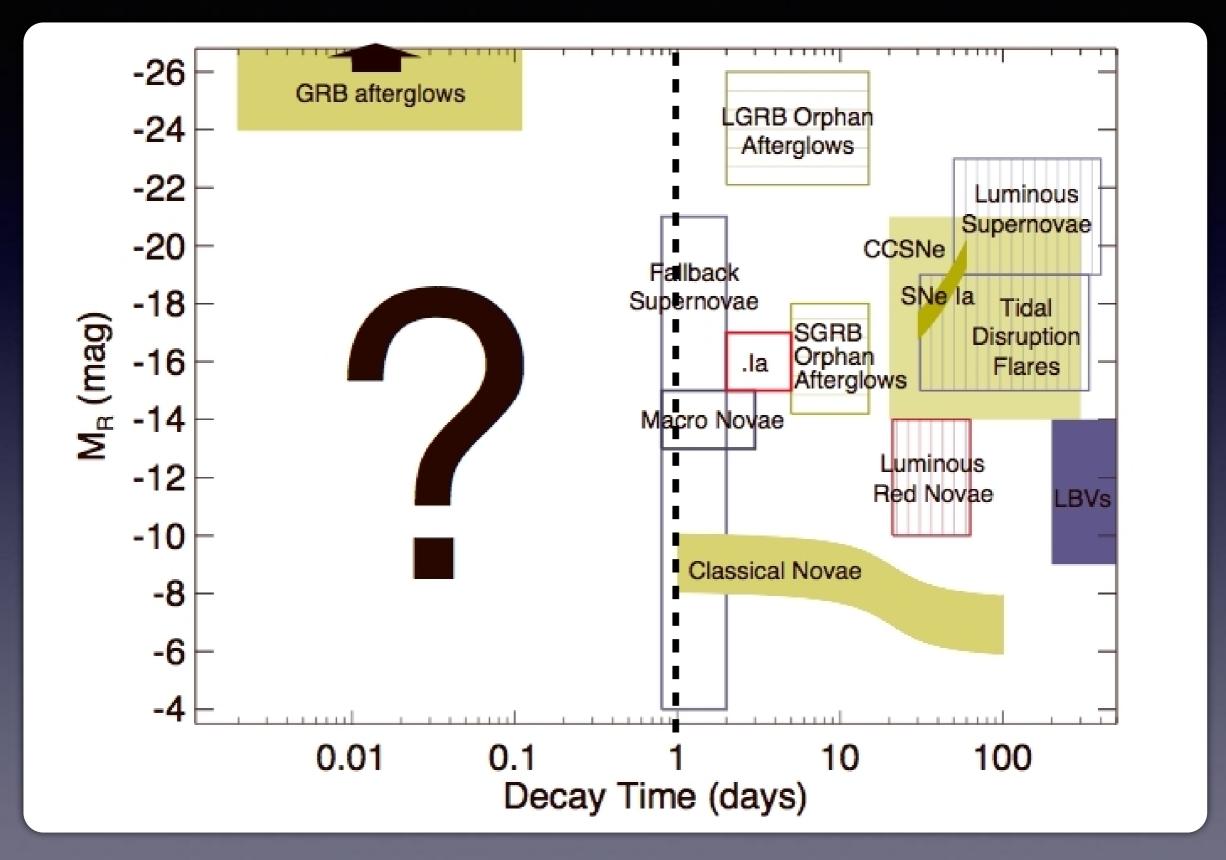
Keane+2016

Parkes beam size
~ 15 arcmin (FWHM)

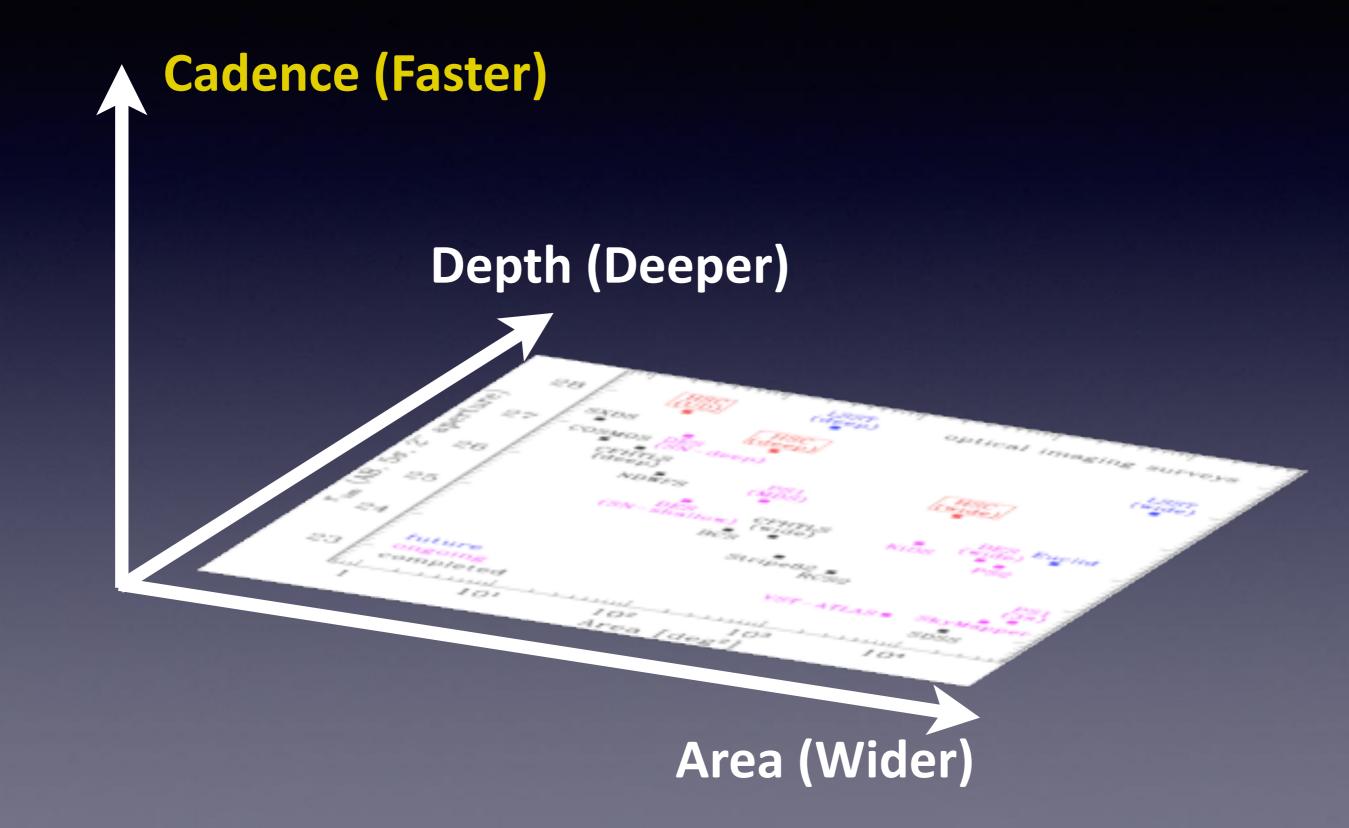


- Transient survey with Subaru/HSC
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# Frontier of transient sky



# 3 key parameters of time-domain survey



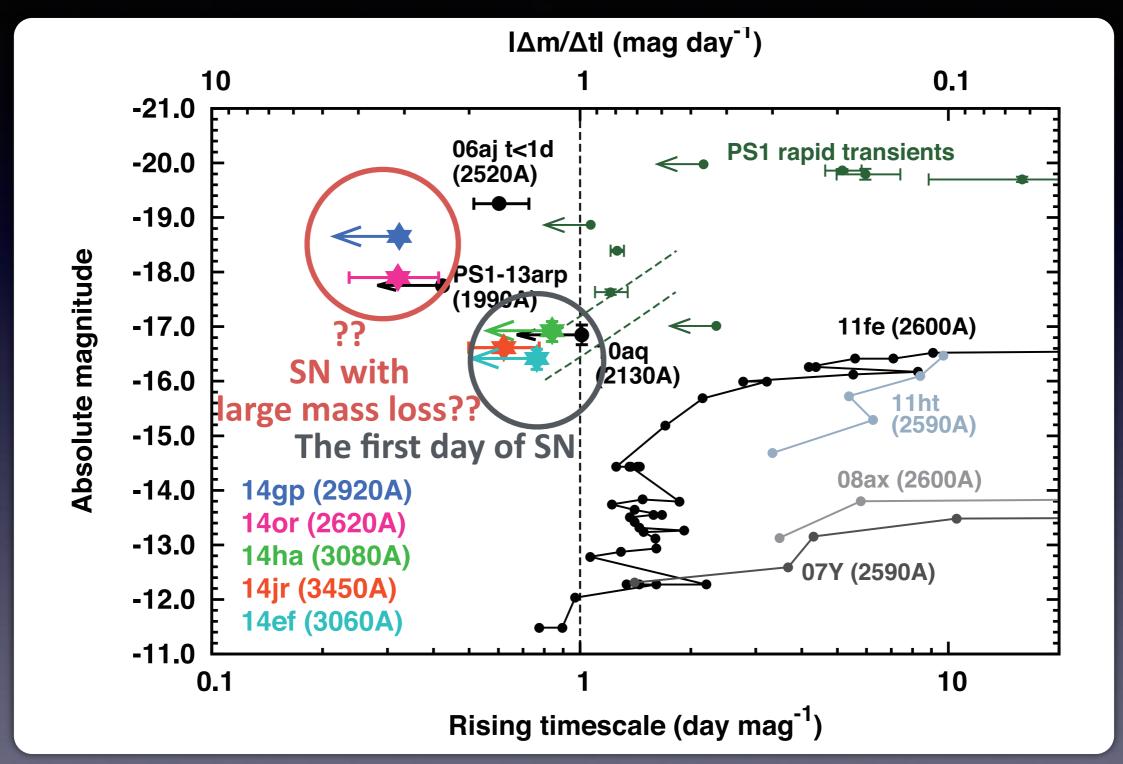
## Fast transients

- First signals from supernovae
  - Shock breakout t ~ a few hours
  - Cooling phase t ~ a few days
- BH-forming SN
  - Disk outflow Mej ~0.01 Msun t ~ a few days
- Accretion induced collapse of WDs
  - Disk outflow Mej ~ 0.01 Msun t ~ 1 day
- Neutron star merger
  - Disk outflow t ~ a few days
  - Free neutron t ~ 1 day
- Unknown unknown...

Mej ~ 0.01 Msun



# Pilot high-cadence survey with Subaru/HSC



MT, Tominaga, Morokuma, Yasuda+2016

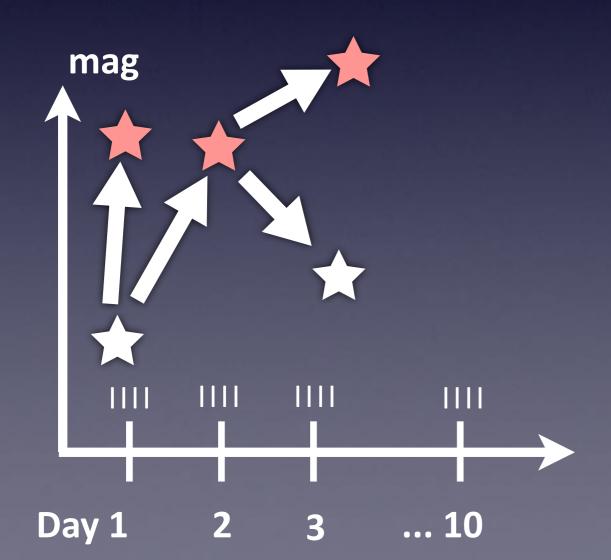
see also Ofek et al.; Gezari et al.

=> Coordinated survey & rapid follow-up are critical

## High-cadence transient survey with Subaru/HSC

- Depth: 25 mag (1 min exposure)
- Cadence: 1 hr + 1 day
- 10 consecutive 0.5 nights (open-use)

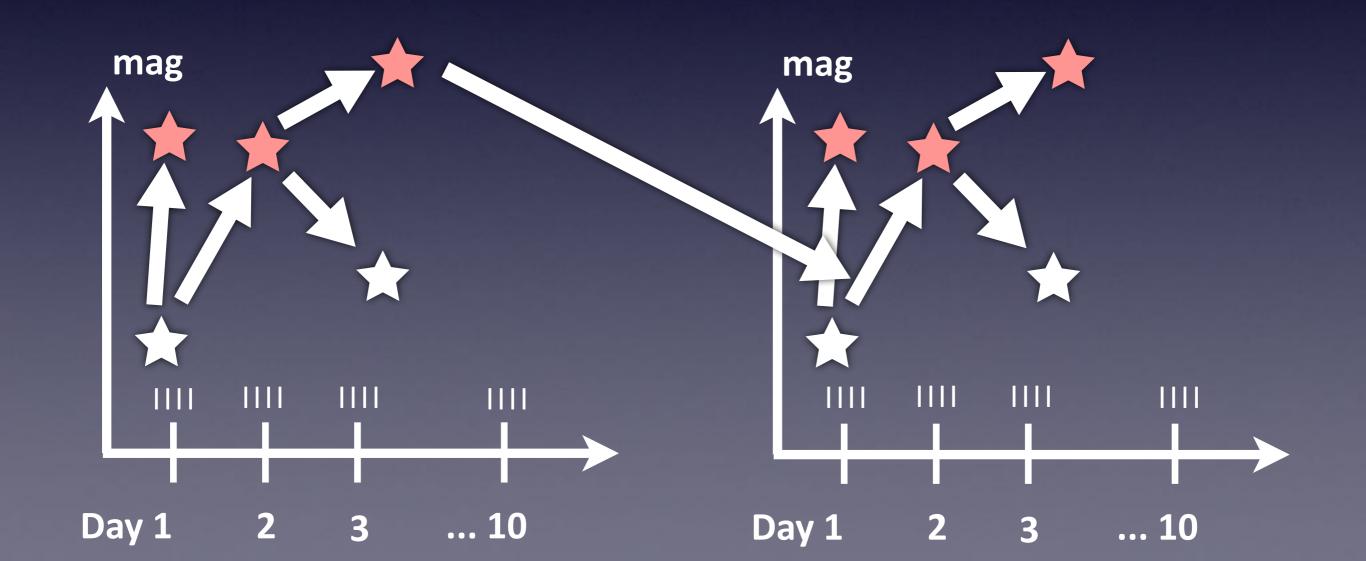
- + Follow-up spectroscopy
- + Coordinated survey (multi-wavelength)



## High-cadence transient survey with Subaru/HSC

- Depth: 25 mag (1 min exposure)
- Cadence: 1 hr + 1 day + 1 month

- + Follow-up spectroscopy
- + Coordinated survey (multi-wavelength)
- 10 consecutive 0.5 nights x 2 (intensive)



- (1. Myself)
   Time-domain, transient, and multi-messenger (GW, v)
- (2. On-going collaboration)
   SSP transient survey + Keck follow-up (J. Cooke)
- (3. Short/Long term collaboration)
   Short: SSP + follow-up (current scheme)
   Long: Dedicated high-cadence transient survey
- (4. Size)
   ~5 HSC nights (normal) to ~10-20 HSC nights (intensive)
   + spectroscopic follow-up + multi-λ coordination
- (5. Instrument)
  HSC, no big demand for new instruments
- (6. Operation)
   Flexible scheduling = cadence observations in queue mode
   Inter-partner ToO