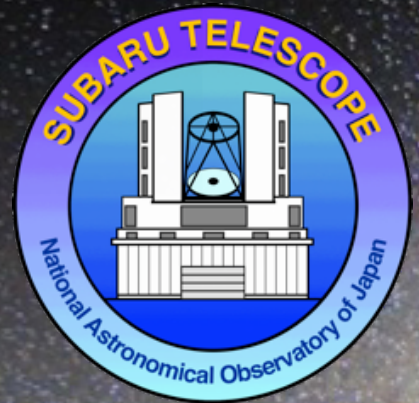


ULTIMATE-Subaru Collaboration Meeting 2018 (2018/1/16)

Discussion (II) - Science

Yusei Koyama (Subaru Telescope)

ULTIMATE-Subaru Science Team



Discussion Items

1. A/Is and (updated) timeline toward CoDR
2. Summarize requirements for ULTIMATE instruments
3. Identify weak points and missing science cases
4. How to establish network of our science team
5. Chair of Galactic team
6. Strong (and realistic) survey design for phase-I, II, III
7. Smaller issues (project website, science team wiki...?)

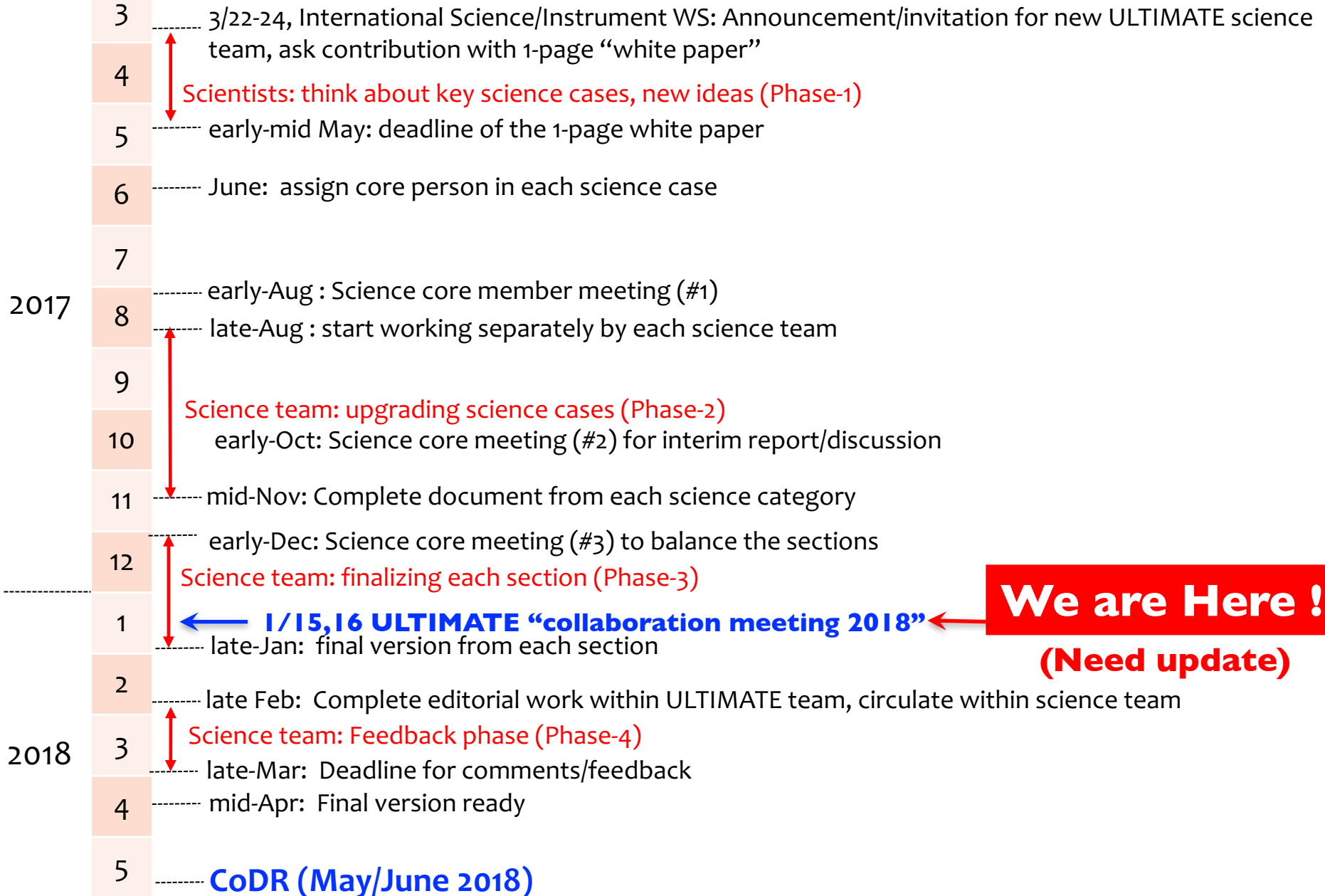
Anything else...?

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Anything else...?

Work schedule toward CoDR 2018



Updated work schedule (2018-2019)

11

12

----- Original submission deadline (from scientists to chair)

1

← **1/15,16 ULTIMATE “collaboration meeting 2018”** ←

We are Here !

2

----- **New submission deadline from each team (2/16) !!**

3

----- early March: Circulate the compiled draft of science sections within the team (3/10)

4

----- **Final/fixed version from each team (4/20) !!**

2018

5

----- May: Complete editorial work within ULTIMATE WG

6

----- early June: Circulate the final version of the draft (6/1)

----- **late June: Send the science document to reviewers (6/30?)**

7

----- **ULTIMATE GLAO CoDR (mid July) !!**

8

----- Polish science requirement on instruments following the CoDR

9

10

----- 1st announcement of ULTIMATE science WS (10/15?)

11

----- 2nd announcement of ULTIMATE science WS, registration open (11/15?)

12

----- Deadline for registration for science WS (12/15?)

TBD

2019

1

2

----- **Feb 2019: Next ULTIMATE Science workshop ?**

Science document to be prepared by CoDR (July 2018)

1. Executive Summary

2. Science cases

- i. High-z imaging ← Recent HSC results?
- ii. High-z MOS
- iii. IFU ← Recent SAMI/MaNGA/KMOS results?
- iv. Nearby Galaxies
- v. Galactic Science

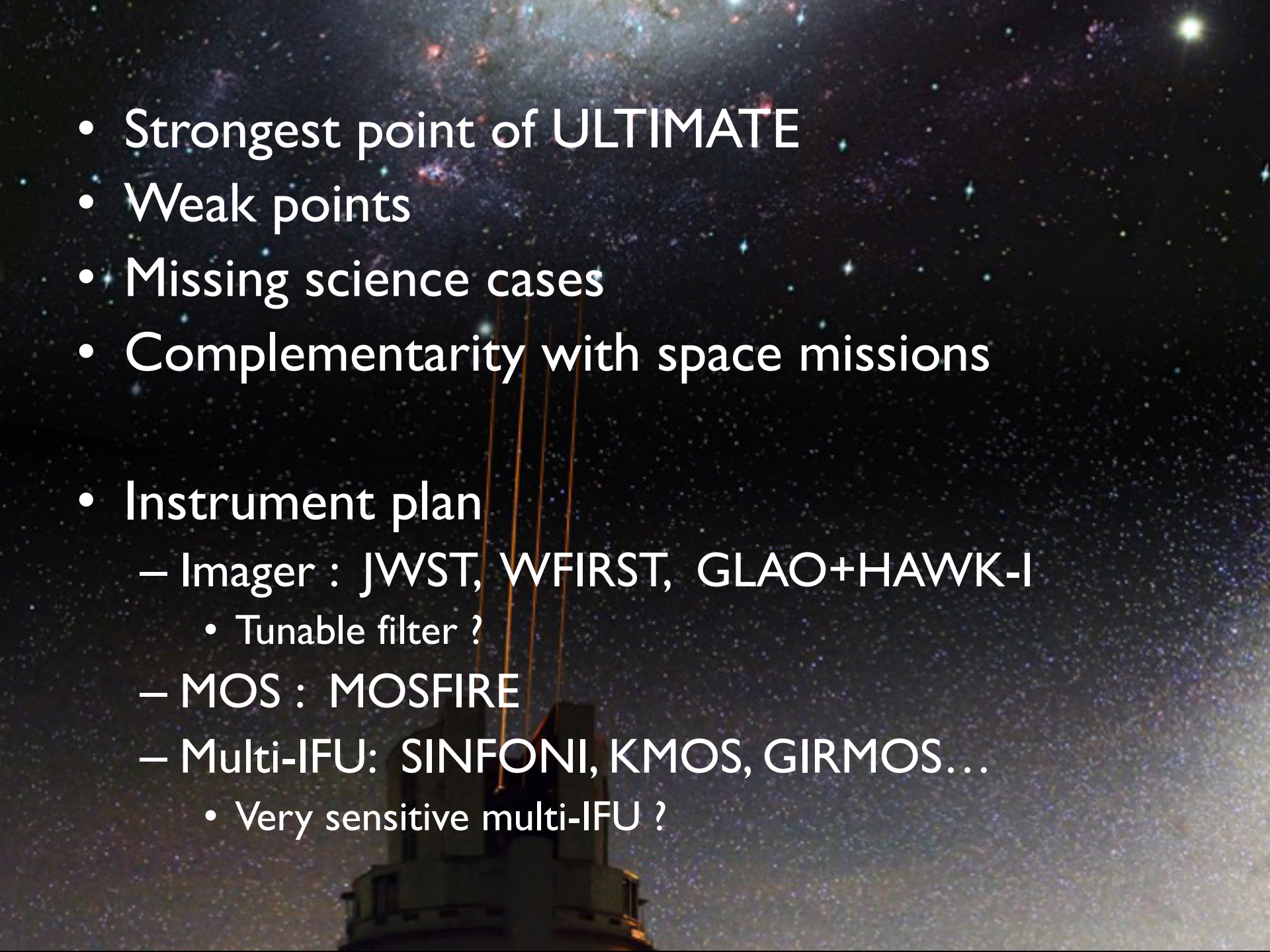
3. Science requirement for instruments

- i. Imaging: sensitivity, FoV, pixel scale, filter set, tunable filter
- ii. Spec: Wavelength coverage, resolution, multiplicity (spec)

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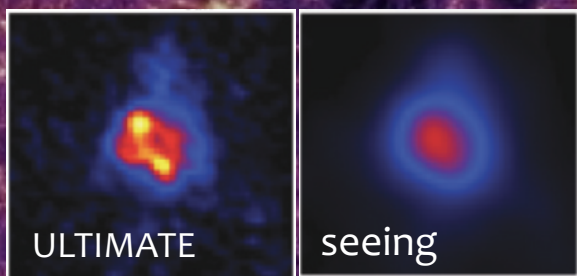
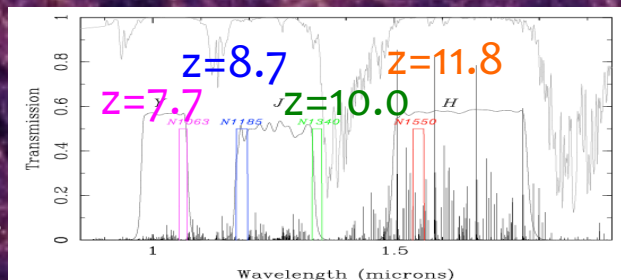
- 
- Strongest point of ULTIMATE
 - Weak points
 - Missing science cases
 - Complementarity with space missions
 - Instrument plan
 - Imager : JWST, WFIRST, GLAO+HAWK-I
 - Tunable filter ?
 - MOS : MOSFIRE
 - Multi-IFU: SINFONI, KMOS, GIRMOS...
 - Very sensitive multi-IFU ?

ULTIMATE-Subaru key science (mostly with high-z imaging)

“Birth, Life, Death” of galaxies in the cradle of large-scale structure

I. First galaxies (birth)

- Unprecedentedly deep NB imaging to detect galaxies a “cosmic dawn” ($z \gg 7$).
- Go beyond the depths of JWST.
- Extension of HSC optical NB survey

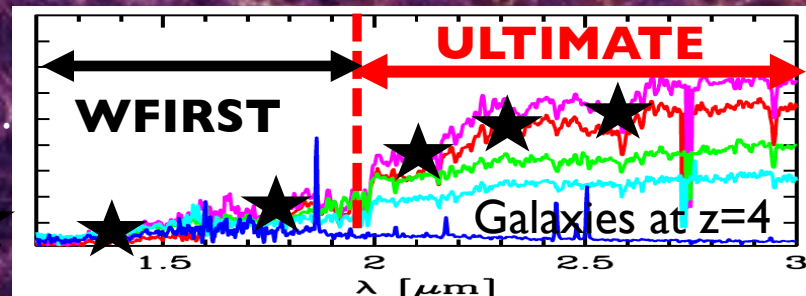


2. Stellar build-up (life)

- Origin of Hubble sequence: bulge, disk, and black hole growth
- Deep & sharp & panoramic NB imaging and 3-D spectroscopy of galaxies at “cosmic noon” ($z=0.5-3.5$)

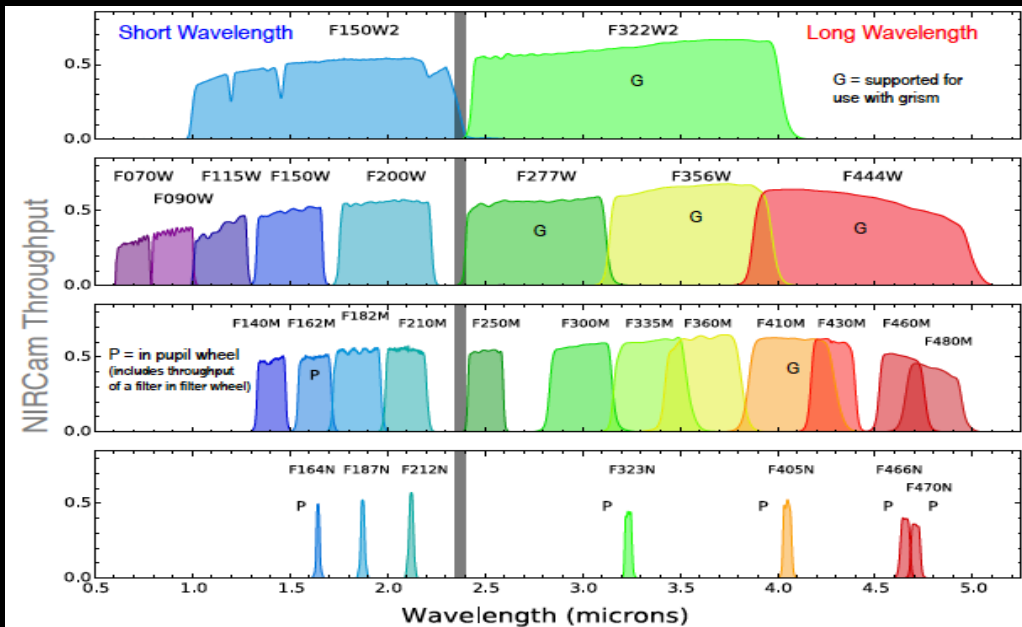
3. Quenching (death)

- Tracking down the “passive” galaxies to $z \sim 5$ with deep BB/MB imaging (in K-band).
- Environment of dead galaxies: do first galaxies die in isolation or in clusters? ★
- Great synergy with WFIRST.



Advantage/complementarity with other space/ground-based facilities

- JWST (NIRCAM):
 - Only 3 NB filters at $1.6 < \lambda < 2.2 \mu\text{m}$ (no NB at $\lambda < 1.6 \mu\text{m}$)
 - No medium-band filters.
 - ULTIMATE has ~ 20 times wider FoV.



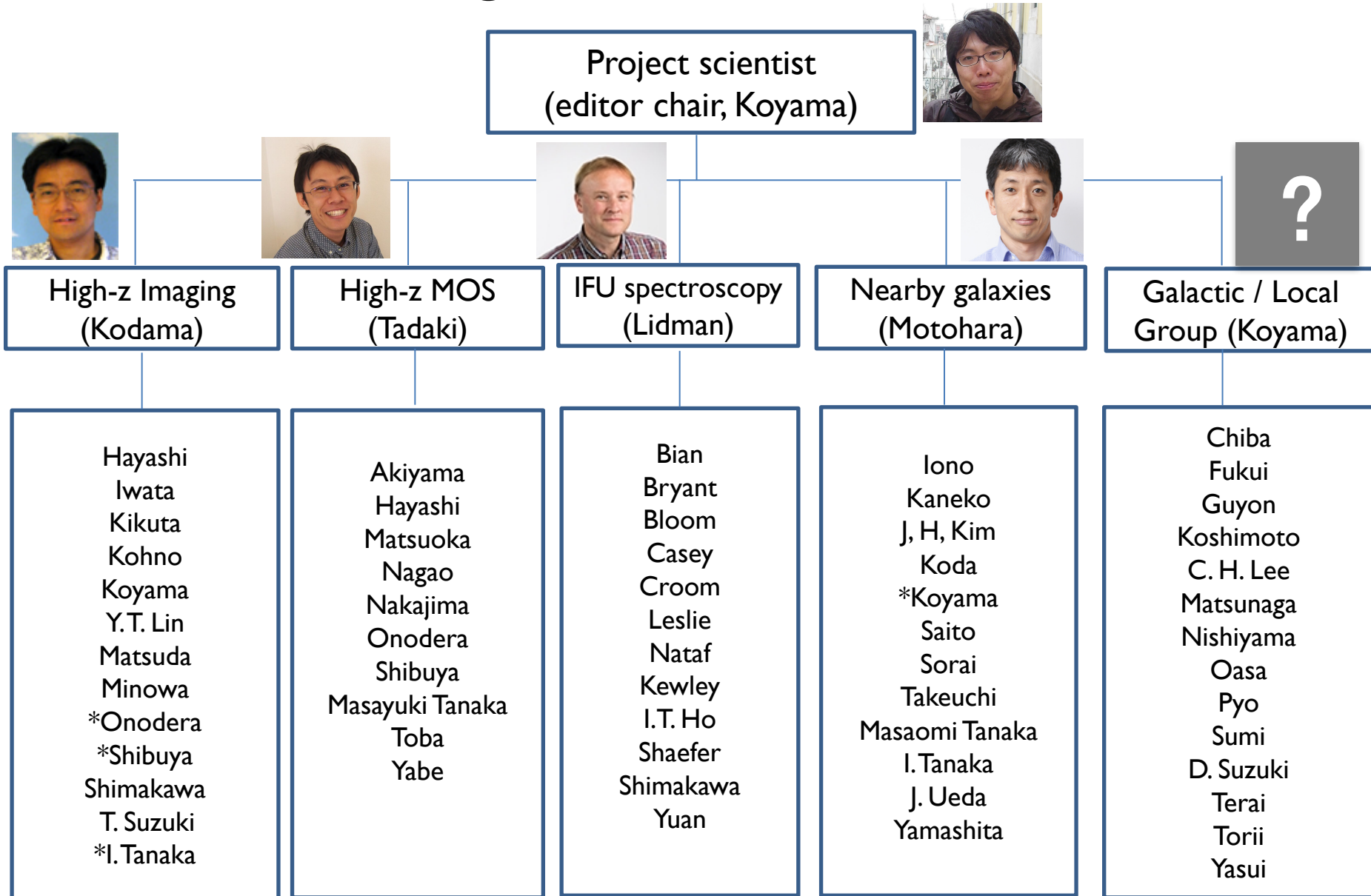
Filter	Sensitivity Point source S/N=10 in 10 ks	Saturation G2V star 80% full well 2 reads of 64x64 subarray
F070W	22.5 nJy	K ~ 9.0 Vega
F115W	13.2 nJy	K ~ 9.6 Vega
F200W	9.1 nJy	K ~ 9.3 Vega
F210M	14.9 nJy	K ~ 8.4 Vega
F212N	129 nJy	K ~ 5.6 Vega
F277W	14.3 nJy	K ~ 9.6 Vega
F322W2	9.1 nJy	K ~ 10.0 Vega
F356W	12.1 nJy	K ~ 8.9 Vega
F444W	23.6 nJy	K ~ 8.0 Vega

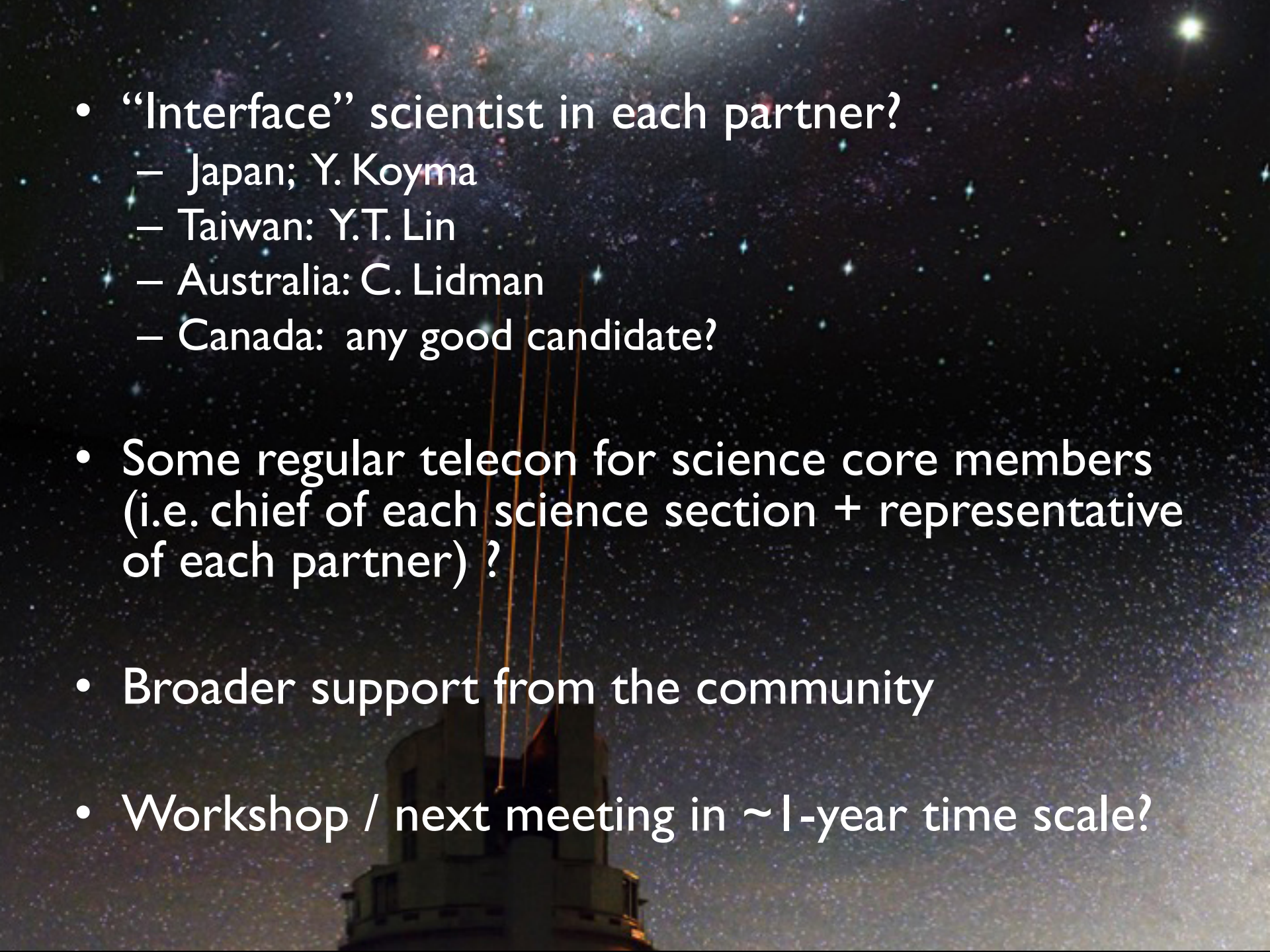
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Anything else...?

Science team organization toward CoDR2018



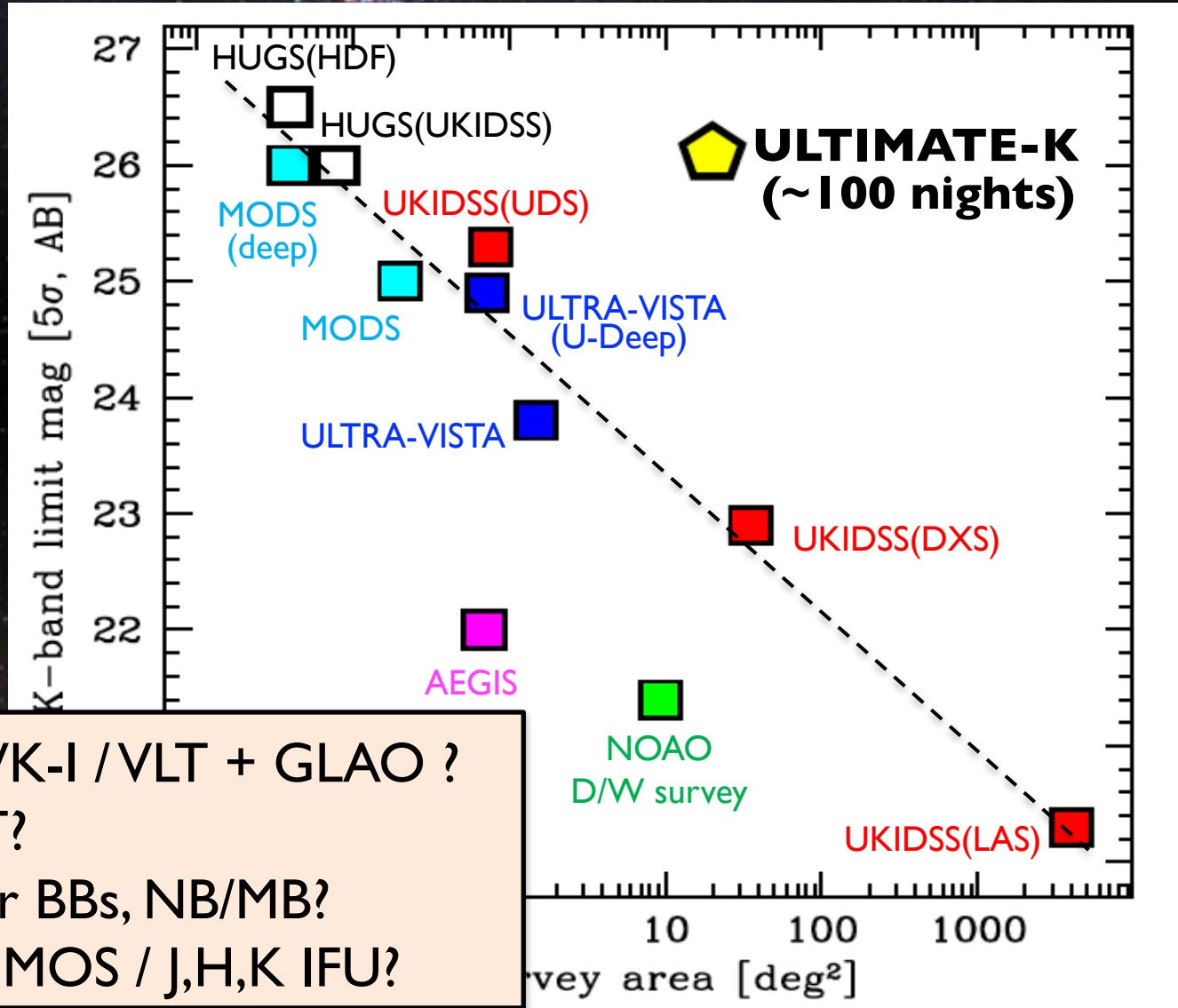
- 
- “Interface” scientist in each partner?
 - Japan; Y. Koyma
 - Taiwan: Y.T. Lin
 - Australia: C. Lidman
 - Canada: any good candidate?
 - Some regular telecon for science core members (i.e. chief of each science section + representative of each partner) ?
 - Broader support from the community
 - Workshop / next meeting in ~1-year time scale?

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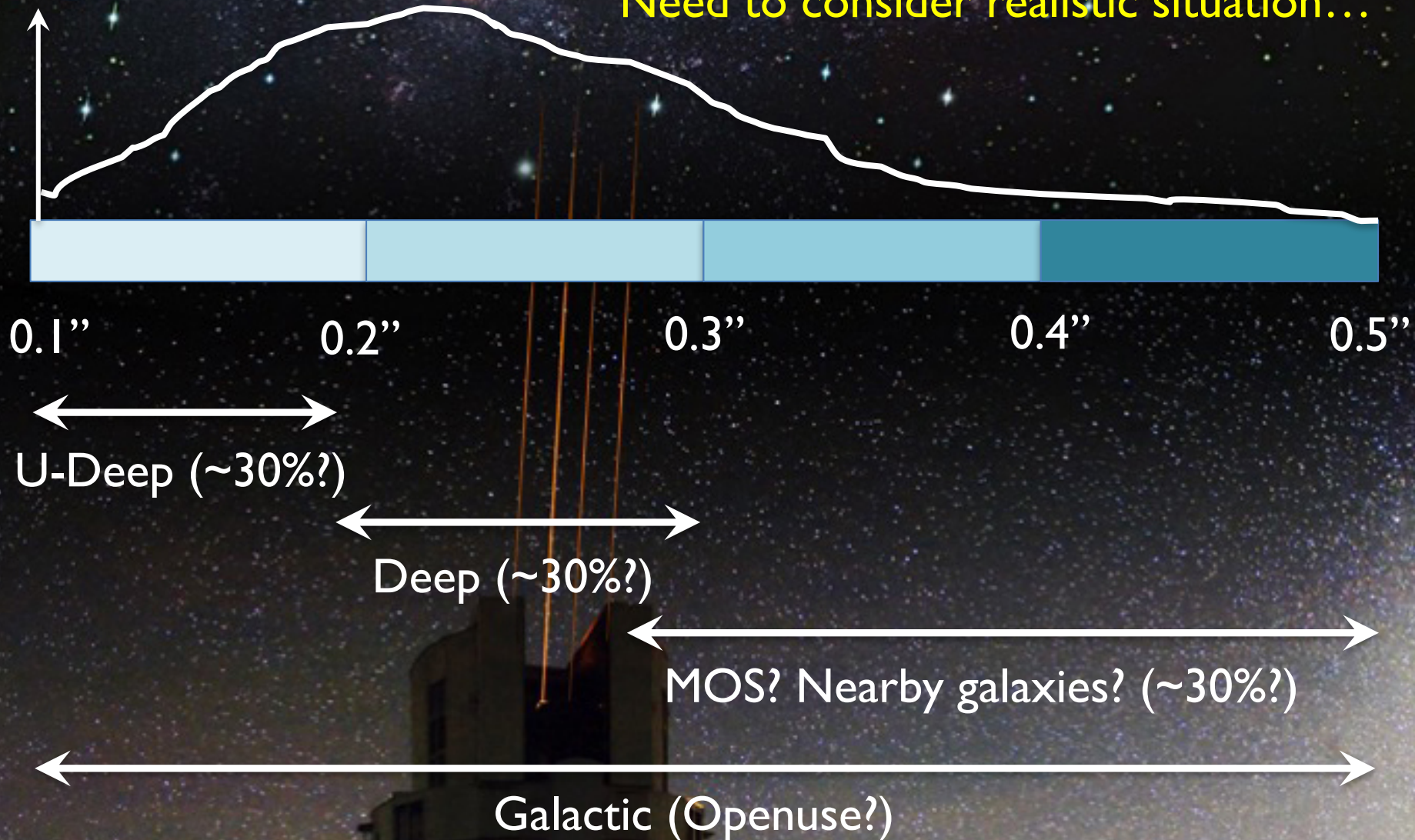
Anything else...?

K-band deep/wide survey



Survey design (SSP with WFC+GLAO)

Need to consider realistic situation...



Timeline of (large) programs?

GLAO 1st light




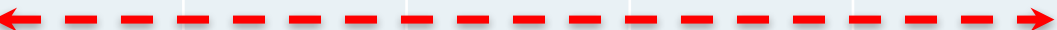

↓ ↓ WFIRST launch


WFI?



M-IFS?



	2025	2026	2027	2028	2029	2030	2031
High-z imaging	Precursor study with MOIRCS openuse		 SSP with WFI ?				
High-z MOS	 GLAO + MOIRCS Intensive?		...new MOS instrument? Imaging + MOS SSP ?				
IFU						 M-IFS intensive?	
Nearby	Precursor study with MOIRCS openuse		 SSP with WFI ?				
Galactic	 Openuse (or organize Galactic intensive?)						

 **PFS SSP**

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