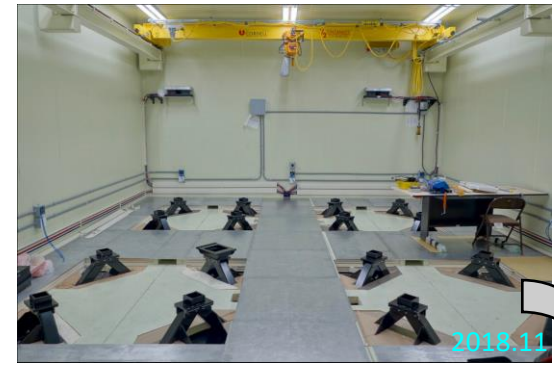


PFS Instrument Status

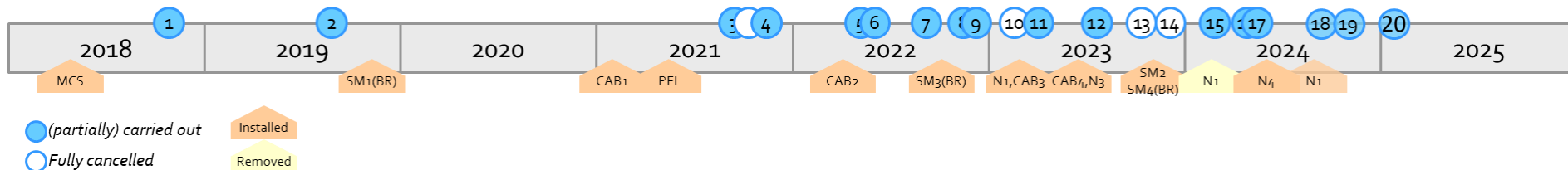
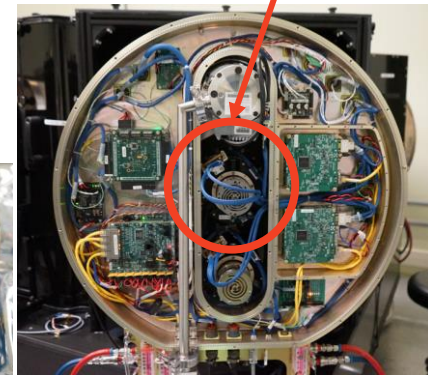
Yuki Moritani (Subaru Telescope/Kavli IPMU)

Completion of hardware delivery in 2024

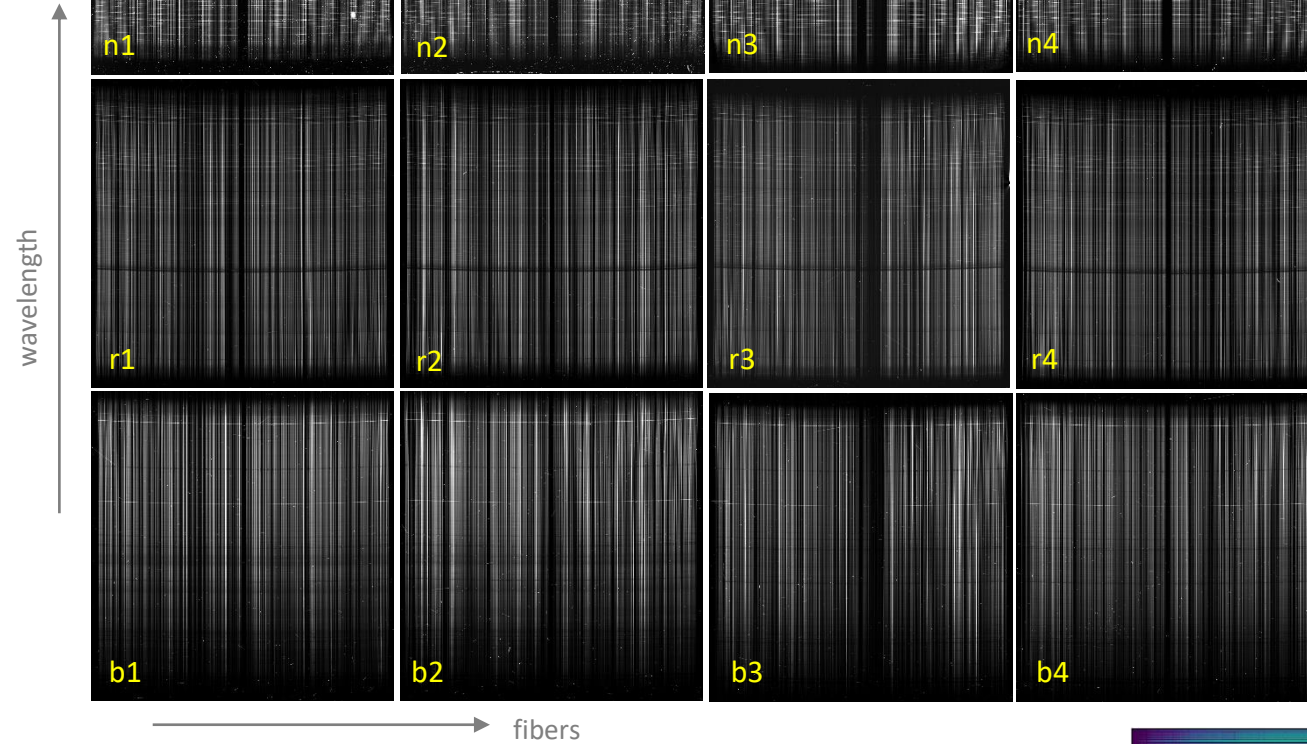
- It was ~6.3 years to complete delivery of the all hardware.
 - Start with MCS in April 2018
 - Complete with NIR camera of SM1 in August 2024
 - A lot of effort by PFS technical members and supports by Subaru staff and DCs.
- Engineering observations along with hardware deliveries.
 - PFS technical team and PFS A-project members carried out integration work and engineering observation one after another.
- Upgrade and modifications of the instruments
 - MCS M1 support, PFI E-box stabilization, AG focal position correction, VPHG orientation correction
- Troubleshoot of the instruments
 - As of today, one SpS camera (NIR for SM2) has trouble with cryocooler. A project members/Inst Div have been consulting the PFS collaborator (JHU) and vendor.
 - Cooler replacement is scheduled in Feb.



Cooler #1

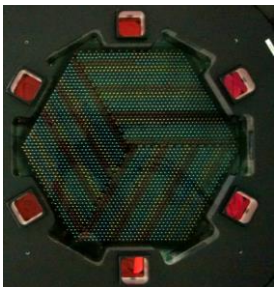
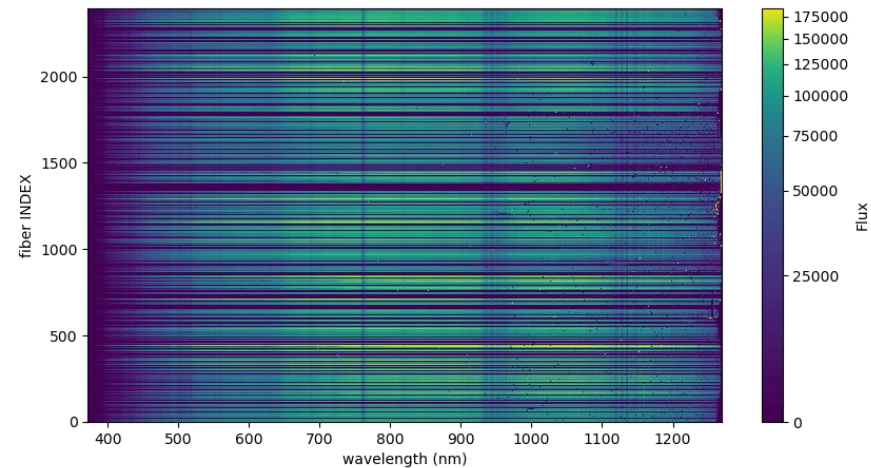


Spectra with the full system



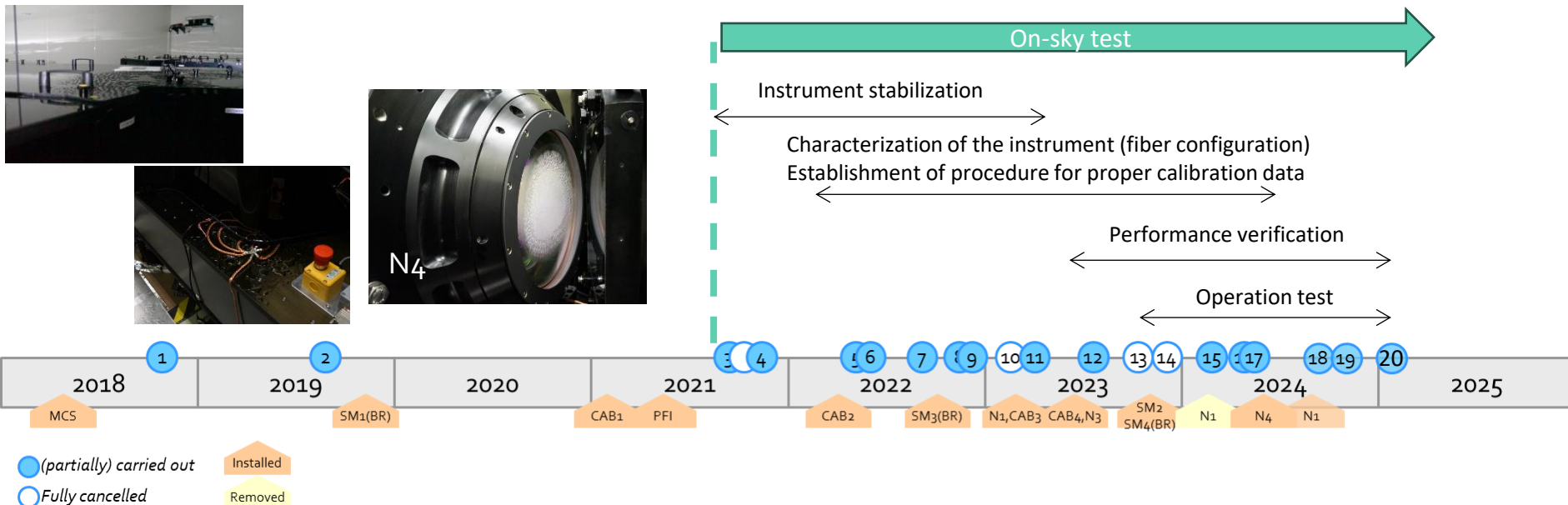
calExp images of the 12 cameras (left) and *pfsMerged* spectra for visit=115021 (a field of many flux-std stars)

pfsMerged 115021 (allfluxstd_field1_pa+90_run18_fluxstd_v3.3)
['u/kiyoyabe/processing/2024I127a']



Engineering Observations

- Engineering run started in 2018
 - “On-sky” test started in 2021
- 20 runs (including on-telescope test during the telescope down time)
 - 150 nights
 - 39 nights (26%) were cancelled (= no access to the summit):
 - 14 nights (9%) due to instrument (PFI) trouble
 - 11 nights (7%) due to bad weather
Note: it doesn't include time when the dome is closed
 - 14 nights (9%) due to telescope shutdown
 - On average, no on-sky data for 1 or 2 night per run due to bad weather
 - (Two big tropical storms “Calvin” and “Hone” came close to Hawaii. Those were the only one for each year.)



Fundamental Instrument Performance

• Fiber configuration

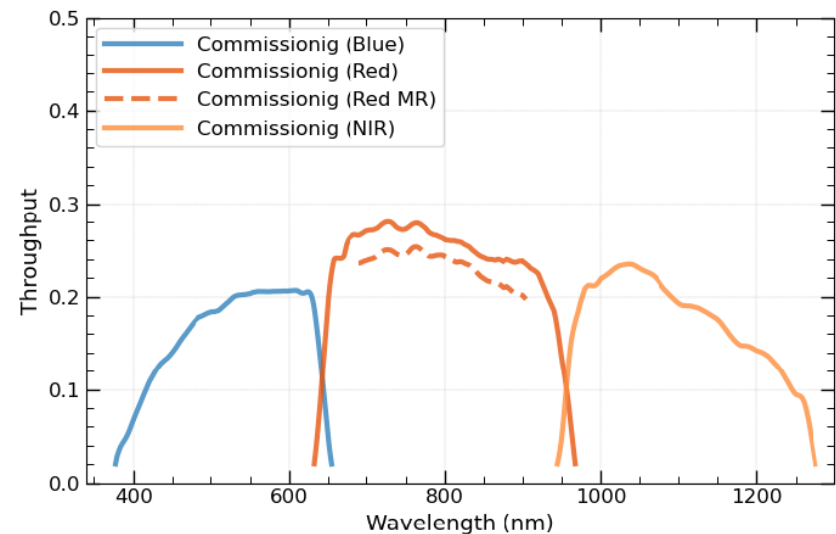
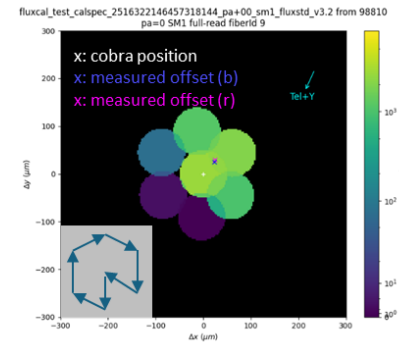
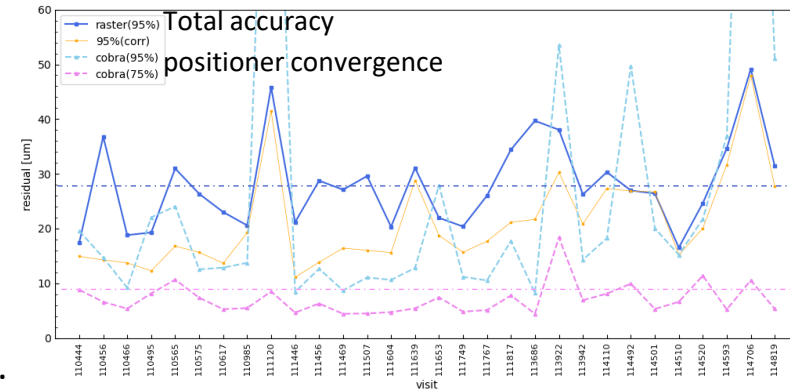
- Configuration time: 124s on average in the last two runs.
 - 8 iterations, skipping to move back to Home position first. 155s with Home position.
- Accuracy of positioners' convergence: 8.5 μm (75%-tile)/ 74.8 μm (95%-tile) in the last four runs.
 - Affected by dome seeing
- Total fiber configuration accuracy : $\sim 20\text{-}35 \mu\text{m}$ (95%-tile) on average
 - Rather stable between run by run. We are monitoring for further optimization.

• Limited range of instrument rotator angle.

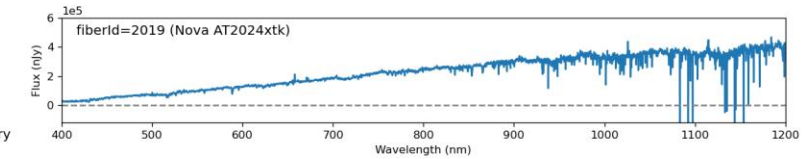
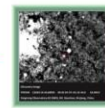
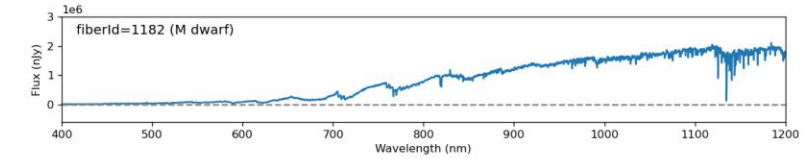
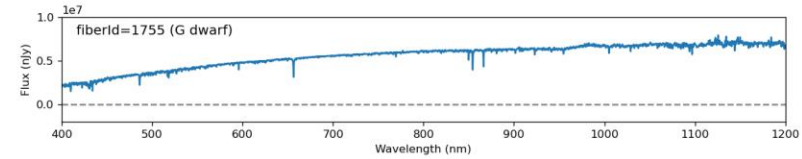
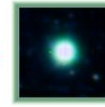
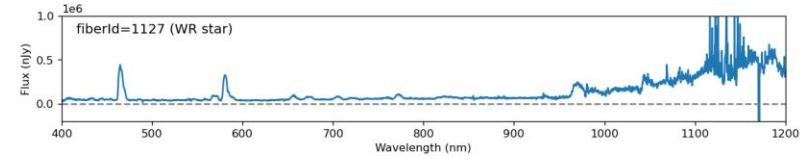
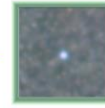
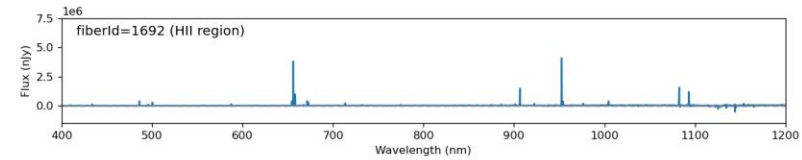
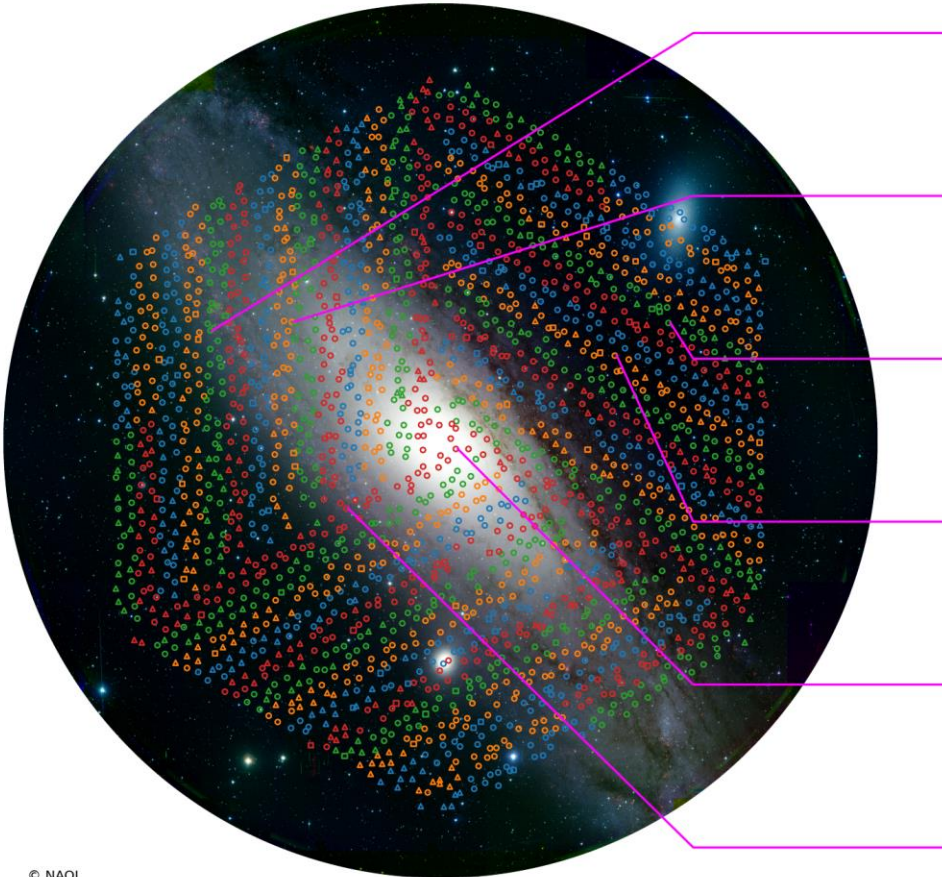
- Instrument rotator: $-174 \text{ -- } +174 \text{ deg}$.
- There is unobservable area for any given PA.

• Throughput

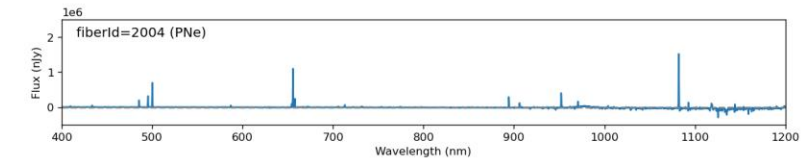
- The system throughput was measured by observing CALSPEC stars as well as F-stars
- It is comparable to expectation from the optics measurement and models (after VPHG correction).



Example of spectra

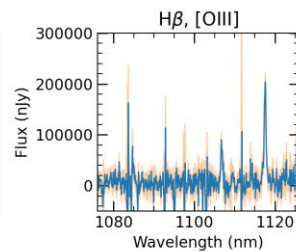
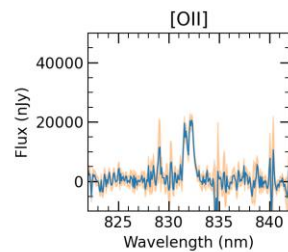
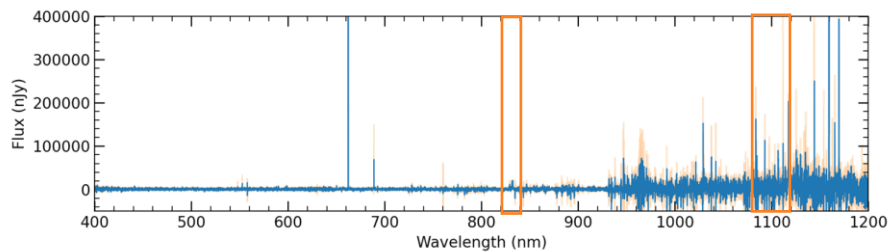


Xingming Observatory

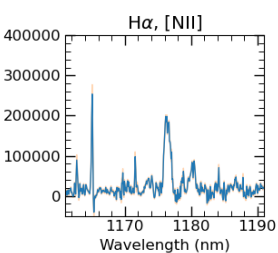
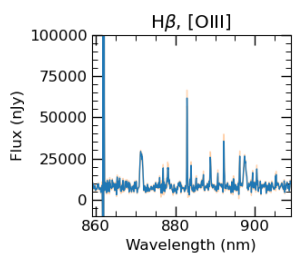
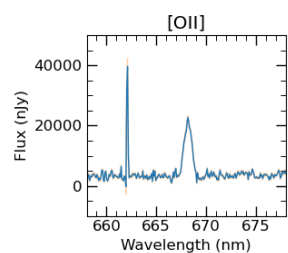
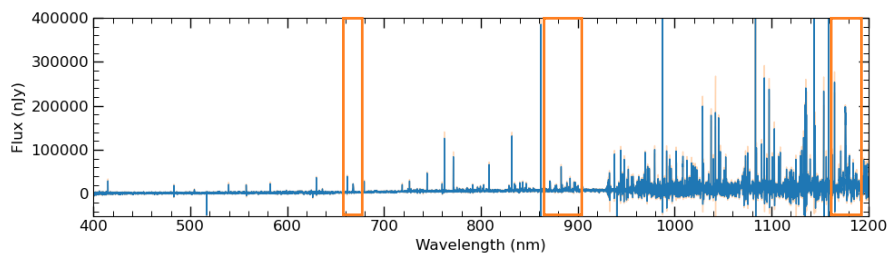
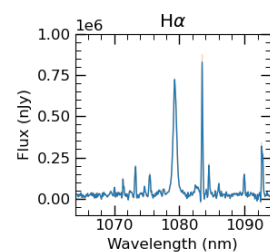
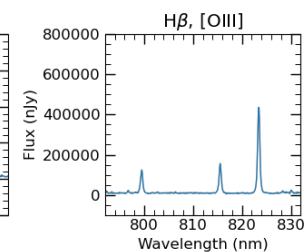
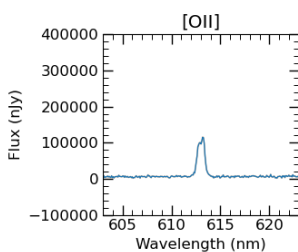
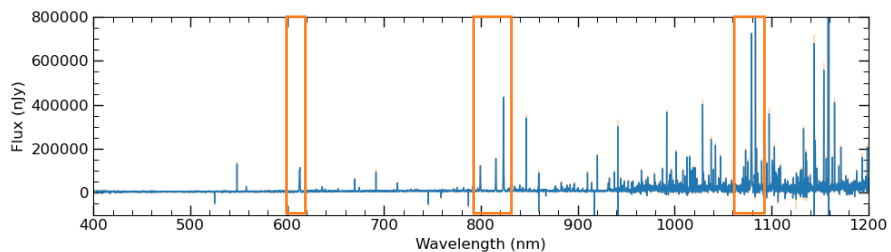
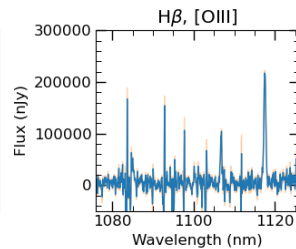
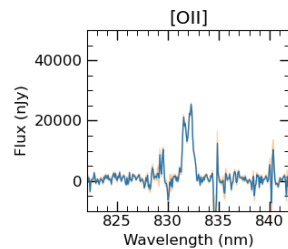
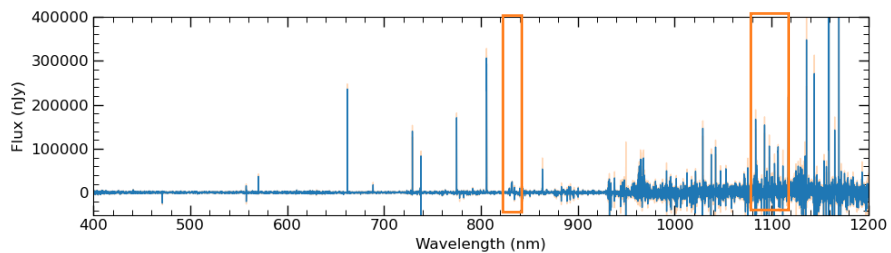


Example of spectra

900s

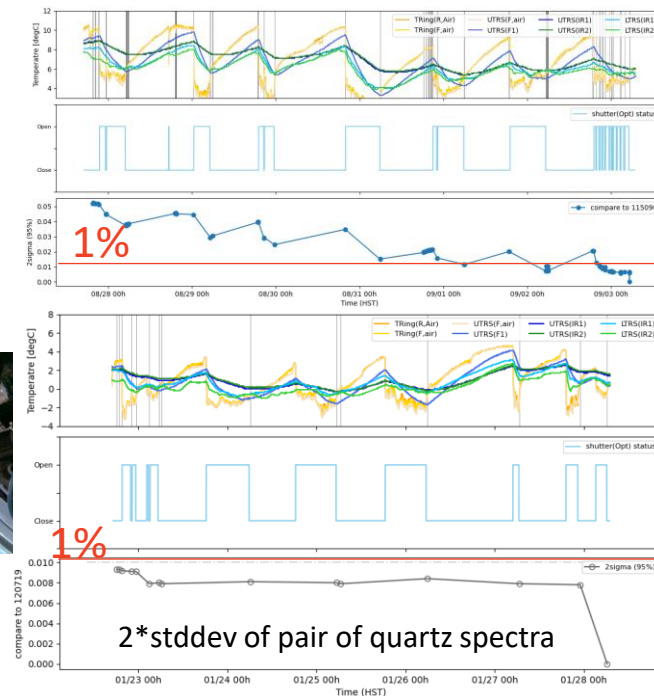


900sx4 (combined)



Fundamental Instrument Performance

- Fiber throughput variation (see P05 for details)
 - Relative throughput among the fibers changes rapidly (<night) which makes accurate sky subtraction impossible.
 - Significant change is seen when we open the dome.
 - To normalize the fiber throughput accurately, quartz data should be taken frequently
 - Worst case is every exposure.
 - High load on telescope and operation
 - In the last two runs, the effect of the index matching gel was tested.
 - If we apply the gel on the fiber connector on the telescope spider, the variation is improved.



• Software development status

- Development of Data Reduction Pipelines (2D & 1D: see P12, P23 and P27)
 - Basic functionality of the processing were developed.
 - The quality is being improved using the on-sky data in the engineering observations.
 - ← Development depends on instrument performance/stability
- The team is putting a lot of effort on better handling the NIR data now (non-linearity, defect, persistence etc)
- To Do:
 - Improvement of field acquisition for high-galactic-latitude area
 - Detailed performance is under analysis
 - etc.