

# Possible Subaru-TESS Synergy Campaign



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on behalf of Japanese TESS consortium

# Outline

- Introduction of TESS
  - Schedule, current status, extended mission
  - Orbit, survey plan, and expected sensitivity
- TESS Science Cases
  - Exoplanets: Targets, planet yield
  - Asteroseismology, supernovae, superflare, variable stars
  - Guest Investigator Program
- Possible Synergies of Subaru with TESS
  - Introduction of Japanese TESS Consortium

# TESS Overview

- TESS is the next flagship space mission for exoplanet studies after Kepler mission, expected to launch in 2017 August
- Almost-all-sky survey of “transiting” planets around nearby 500,000 stars
- 2 year observation with 4 cameras (10cm diameter) each having  $24^\circ \times 24^\circ$  FoV
- Providing **~20 Earth-like transiting planets around HZ** for further follow-up observations with ELTs and JWST



# TESS Timeline

2013: selected as Medium-Class Explores mission

2014: passed PDR, and Vehicle was selected as SpaceX Falcon 9

2015: Critical Design Review, lab test ongoing

2017: **planned to be launched in August**

- about two months to reach the High Earth Orbit

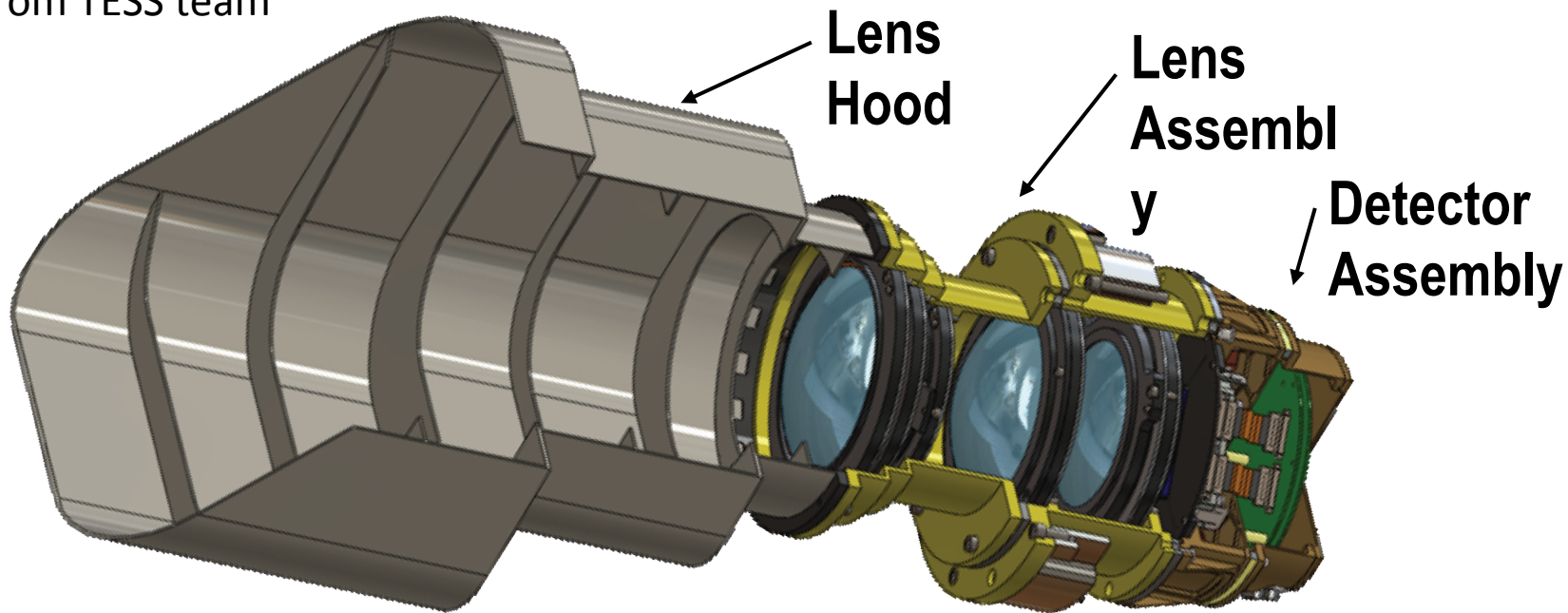
- first science data will be released in 18A semester**

2018: observing the first hemisphere

2019: observing the other hemisphere

2020: **possible extension (discussions ongoing)**

slide from TESS team

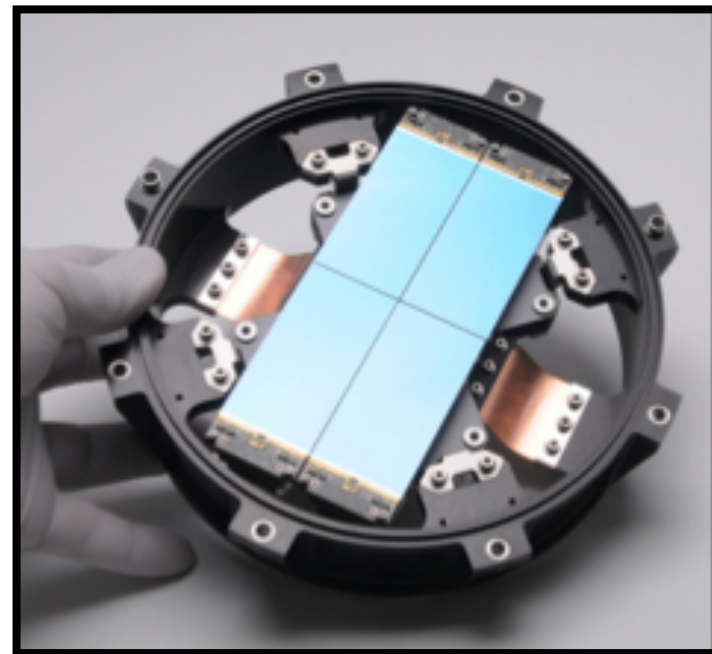


|                               |             |
|-------------------------------|-------------|
| Entrance pupil diameter       | 10.5 cm     |
| Bandpass                      | 600-1000 nm |
| Field of view                 | 24° x 24°   |
| Cadence for target stars      | 2 min       |
| Cadence for full frame images | 30 min      |

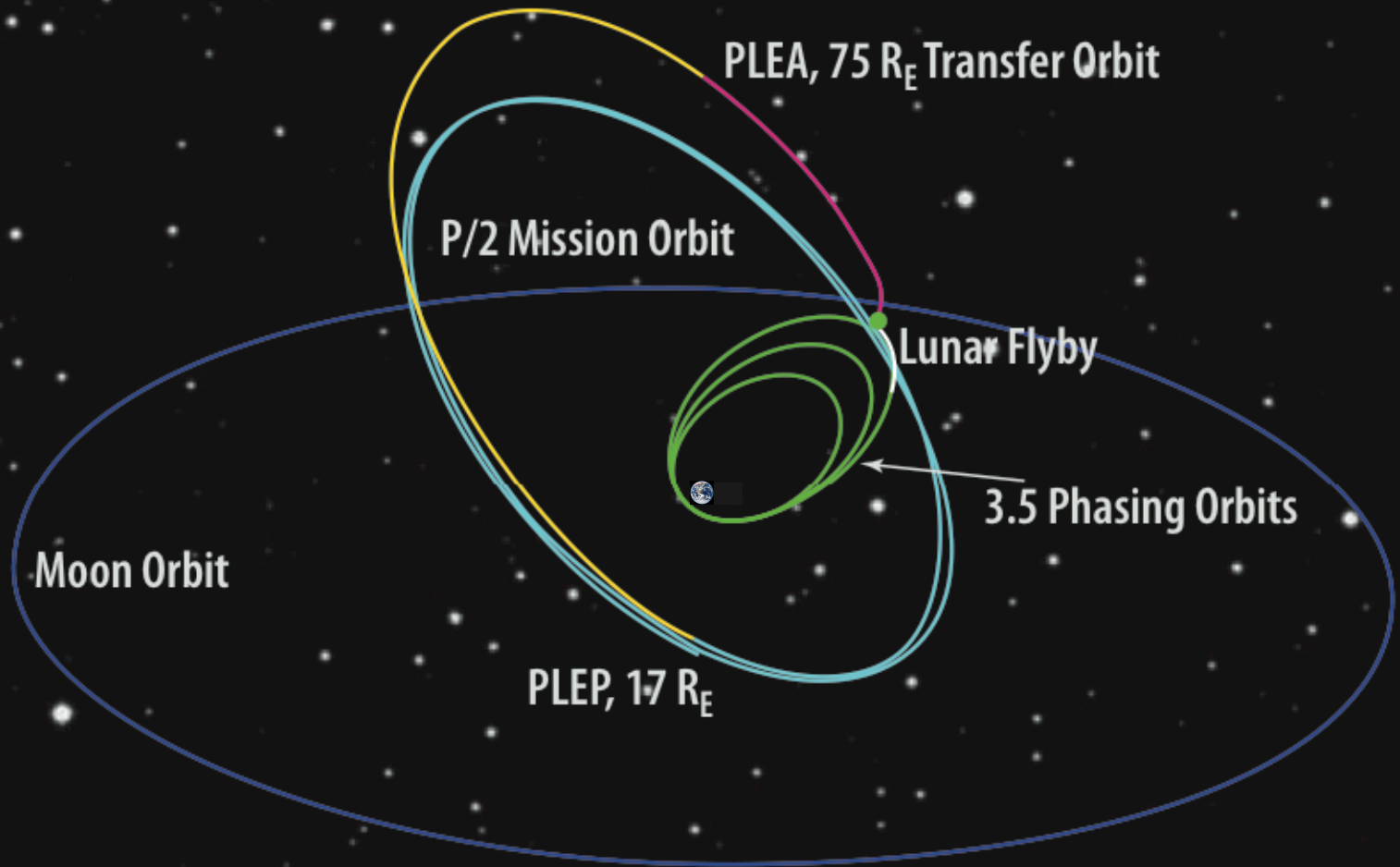
four 2048 x 4096 CCDs

each CCD  
= 2048 x 2048 imaging  
+ 2048 x 2048 storing

in total  
4096 x 4096 imaging

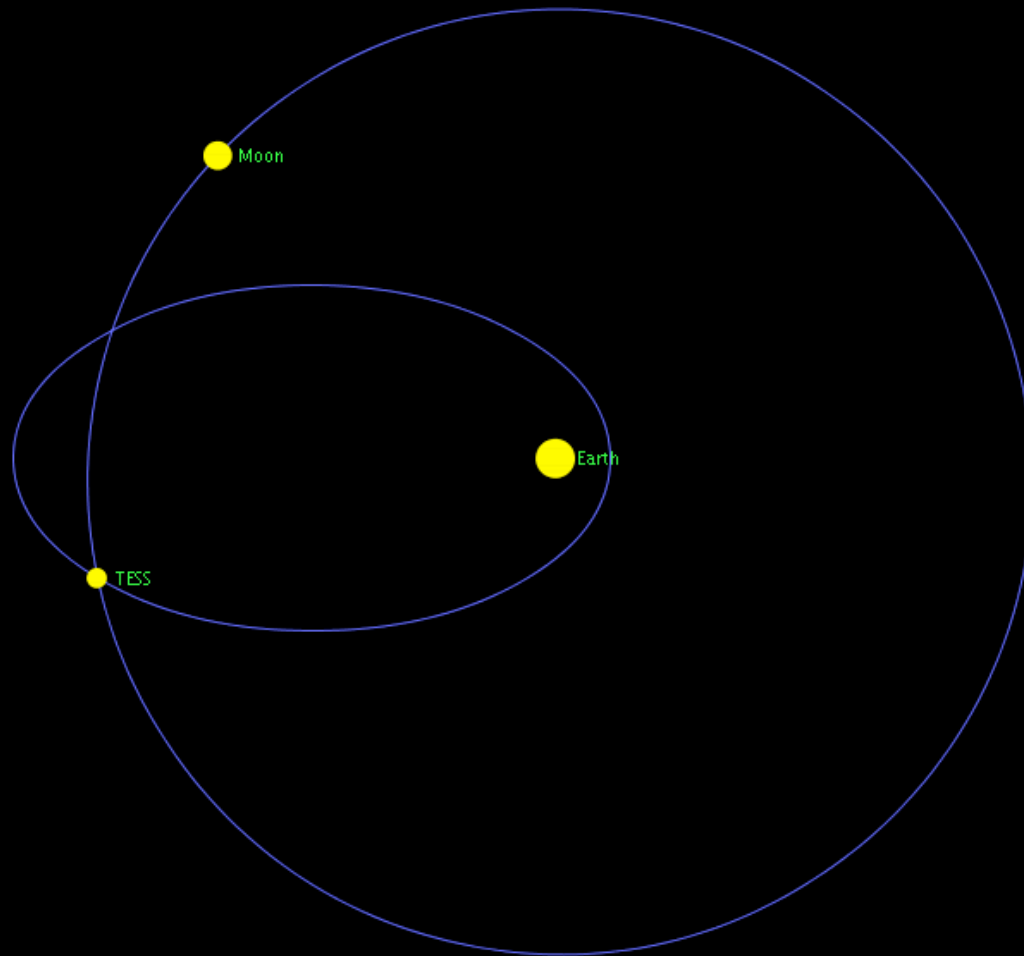


# TESS High Earth Orbit and Downlink



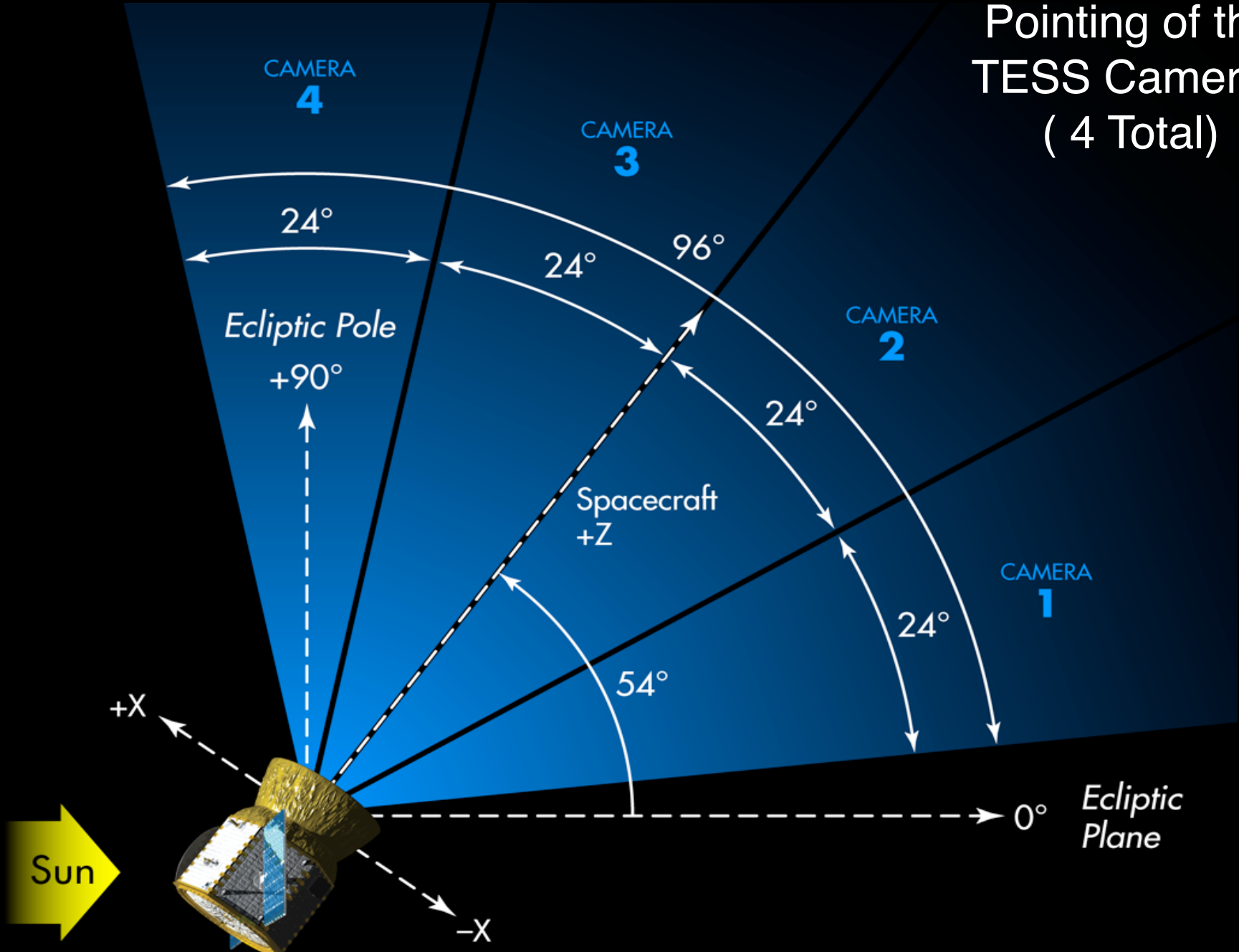
Uninterrupted  
viewing for >95%  
of time

Orbital Periods:  
TESS = 13.7 days  
Moon = 27.4 days  
➔ 2:1 Resonance  
➔ 90° Phasing

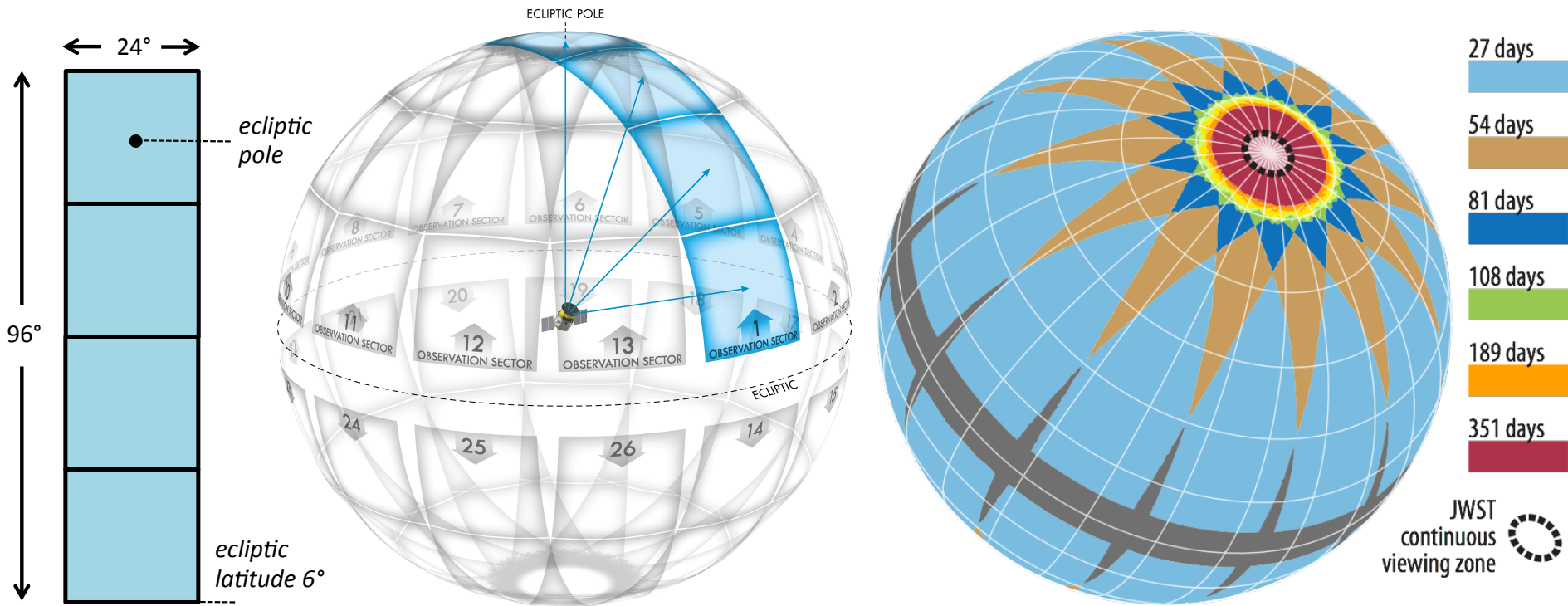


TESS Orbit is **Stable** for Decades (*no station keeping req'd*)

# Pointing of the TESS Cameras ( 4 Total)

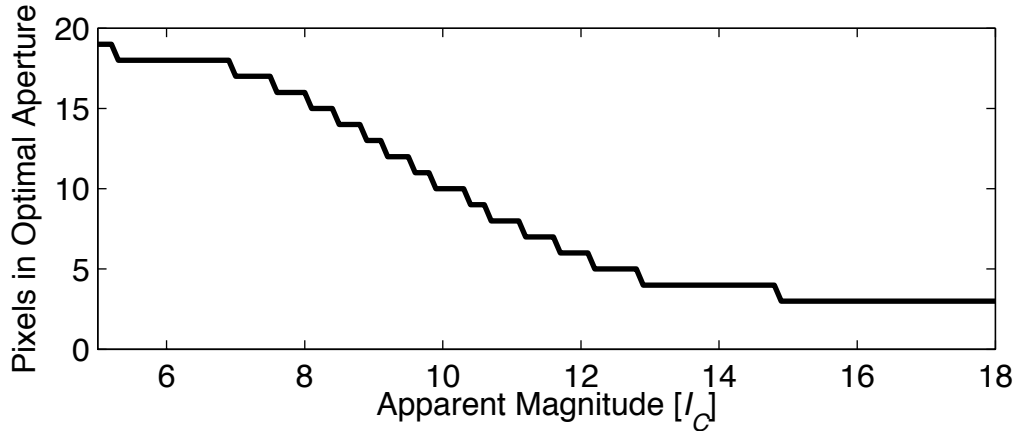
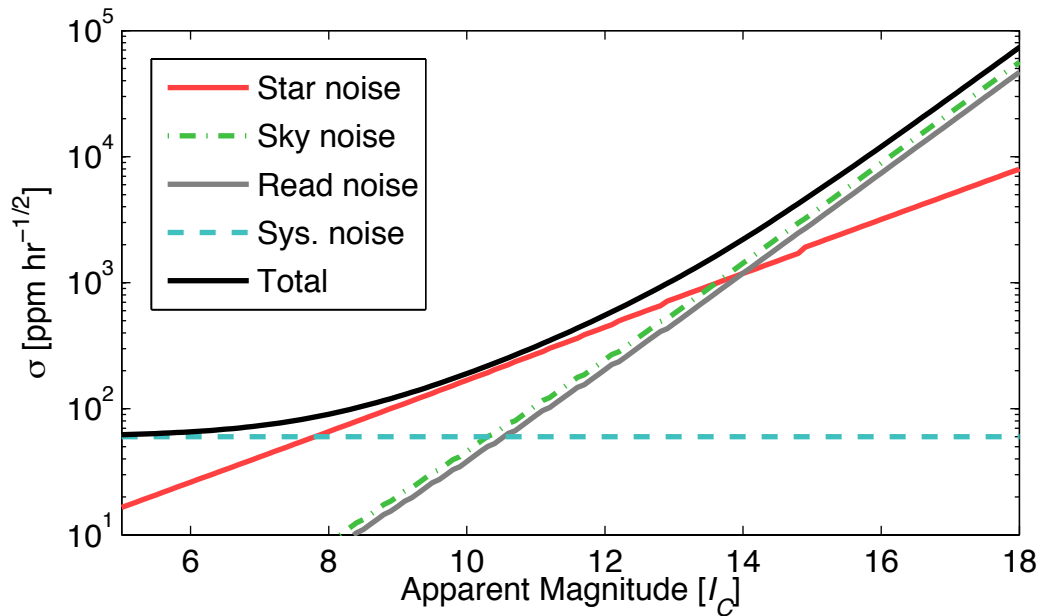


# TESS Field of View and Strategy



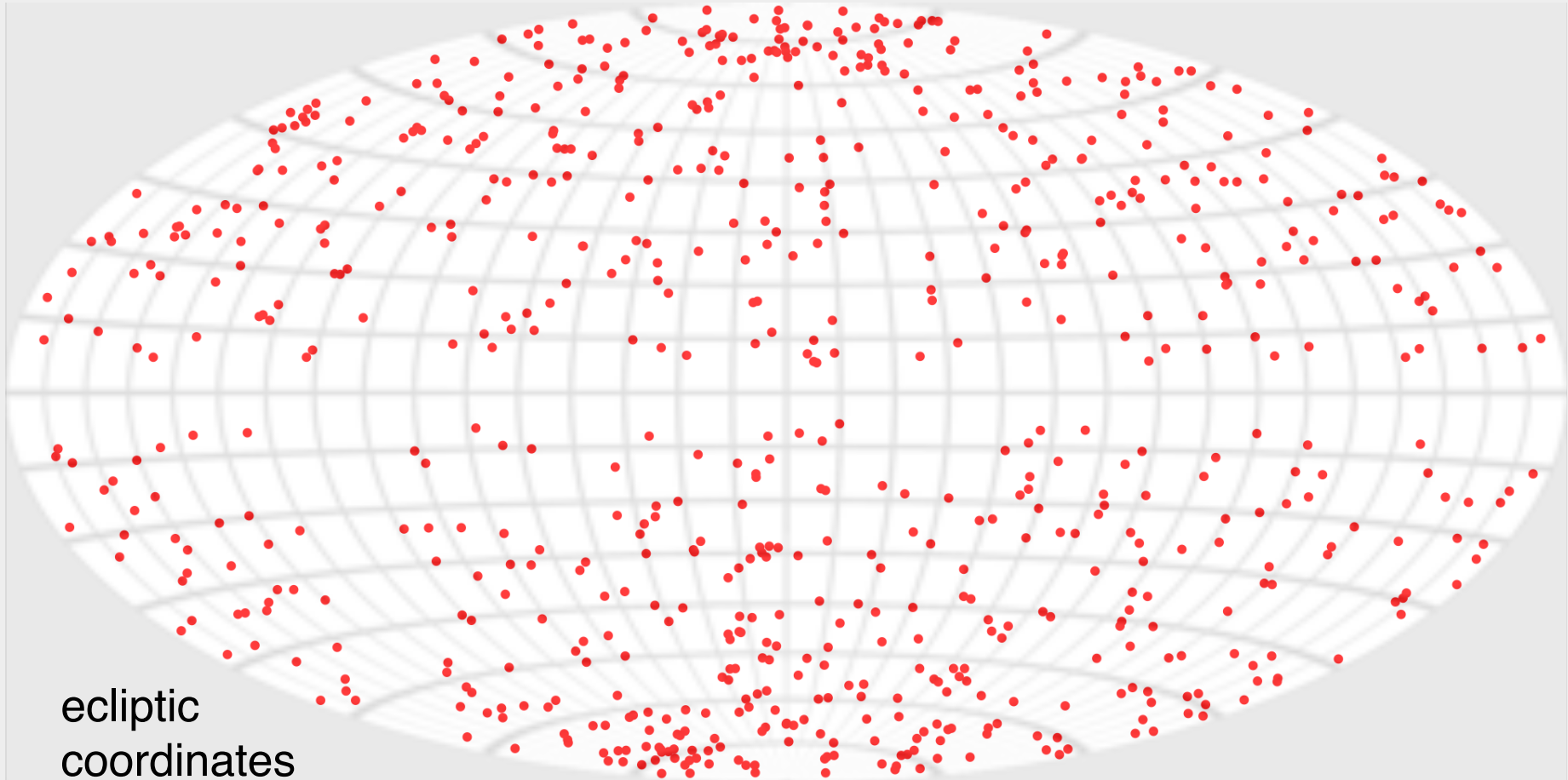
- TESS observes **27 days** for each rectangle sector (24° x 96°)
- Start with one of hemispheres (1 year) and then the other hemisphere (1 year), and possible extension.

# TESS Expected Sensitivity



- Better than 0.1% (per hr) photometric precision for  $I$  mag < 12
- Sufficient to detect (sub)Earth-size planets around M dwarfs
- 1% precision for  $I \sim 16$
- It may be also useful for other science cases

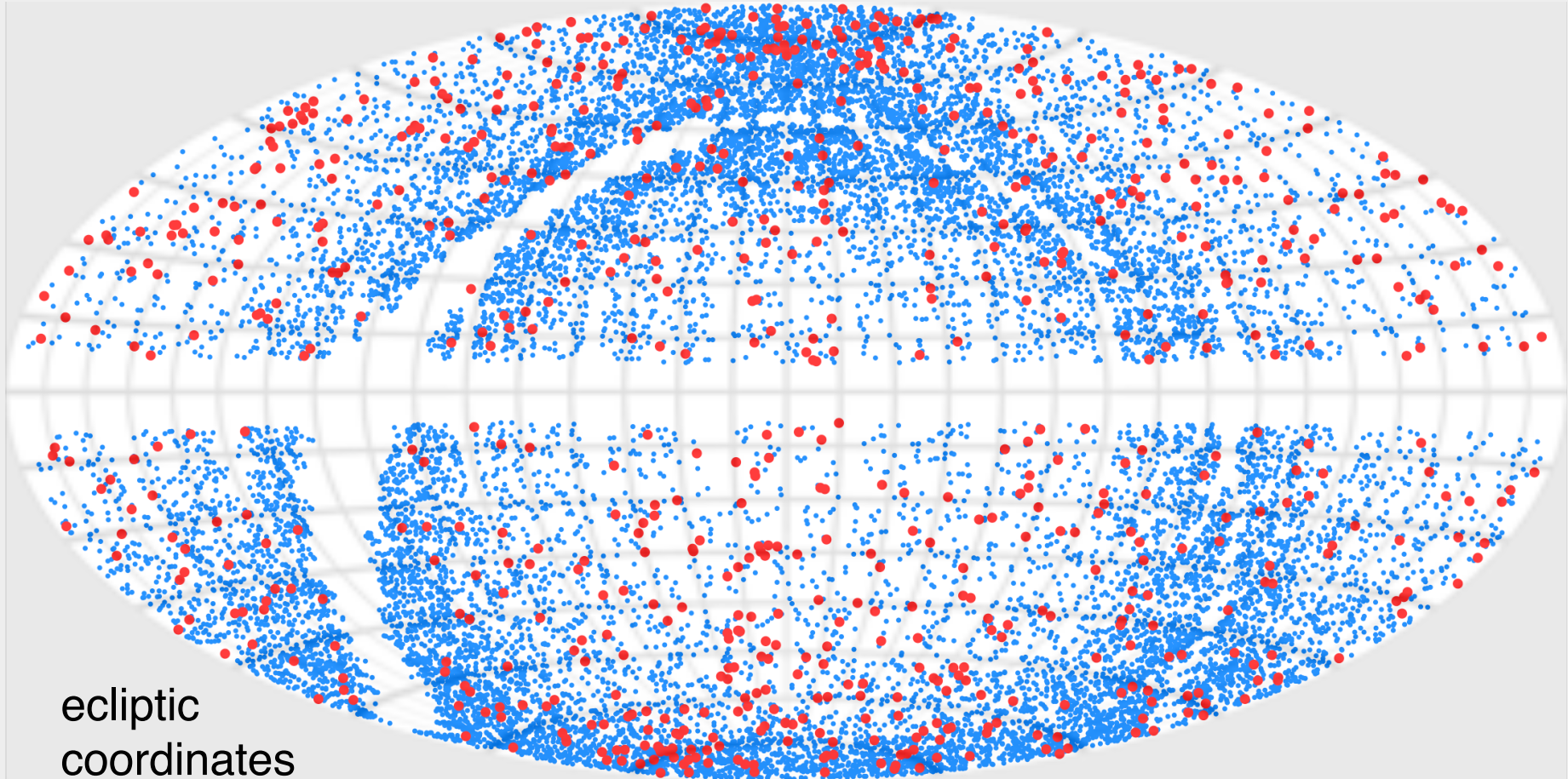
# Simulated TESS detections



Sullivan et al. (2015)

● detectable planets around pre-selected target stars

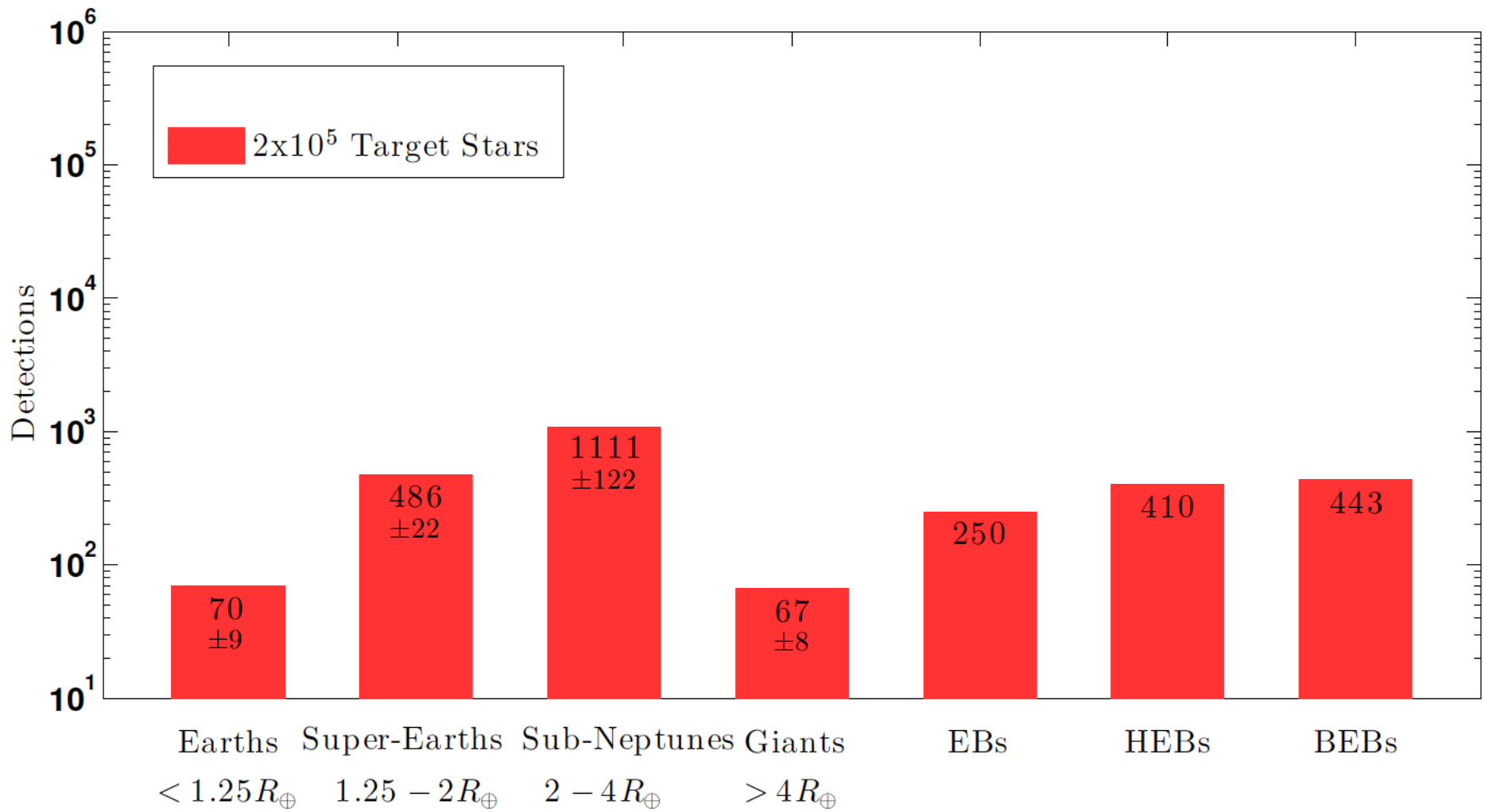
# Simulated TESS detections



ecliptic  
coordinates

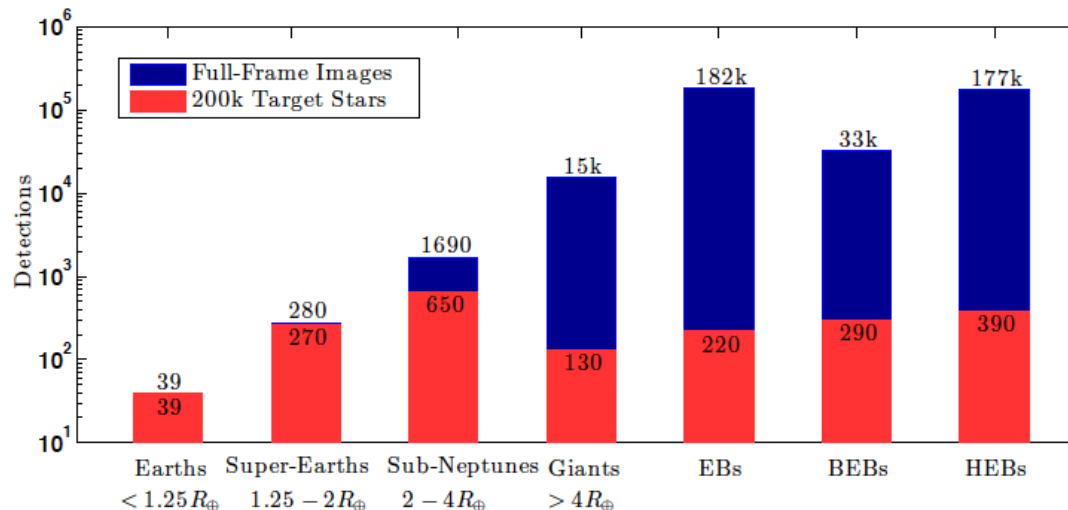
Sullivan et al. (2015)

- detectable planets around pre-selected target stars
- detectable planets around other stars in full-frame images

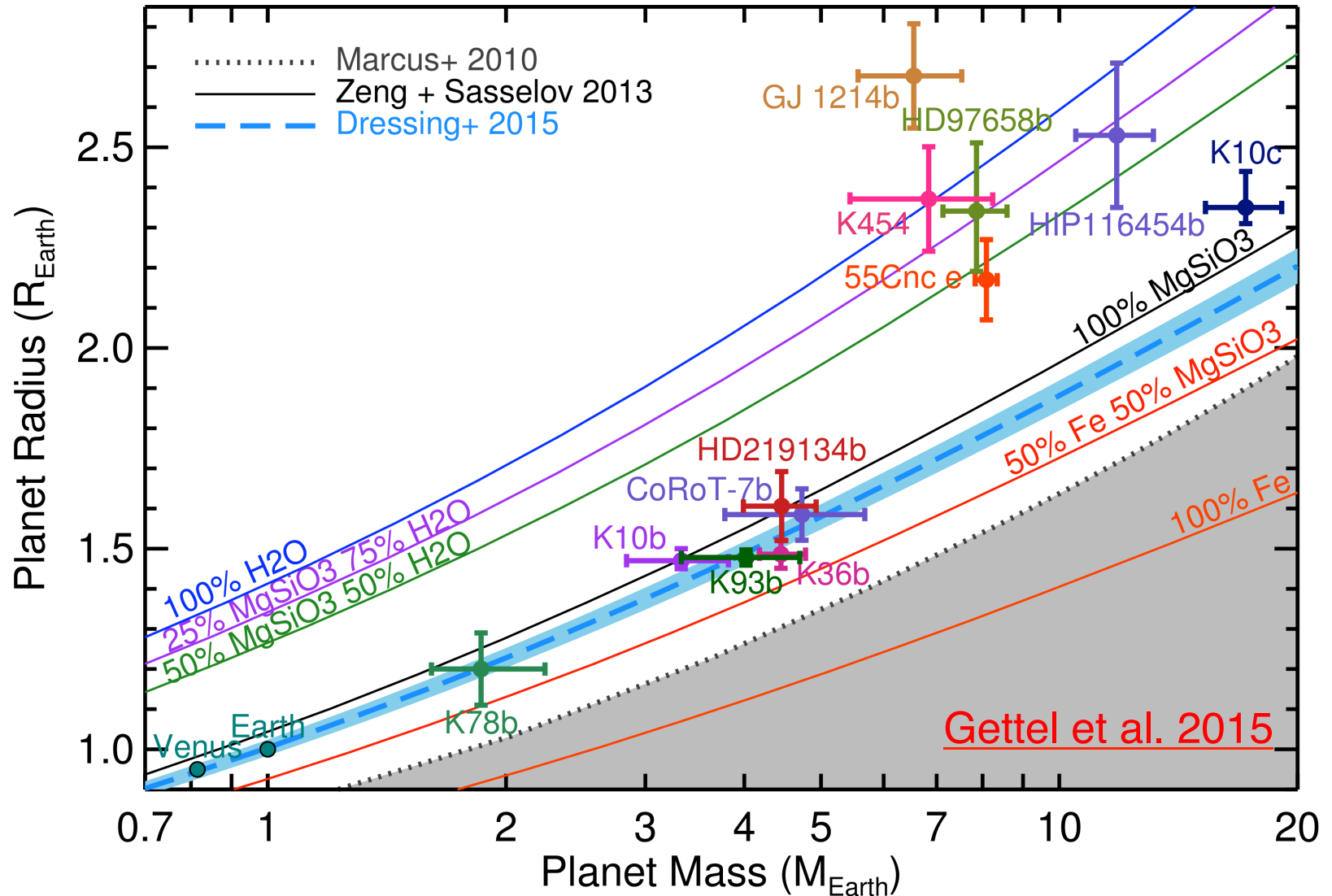


# TESS Planet Yield Simulation (Sullivan+2015)

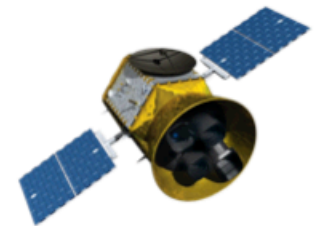
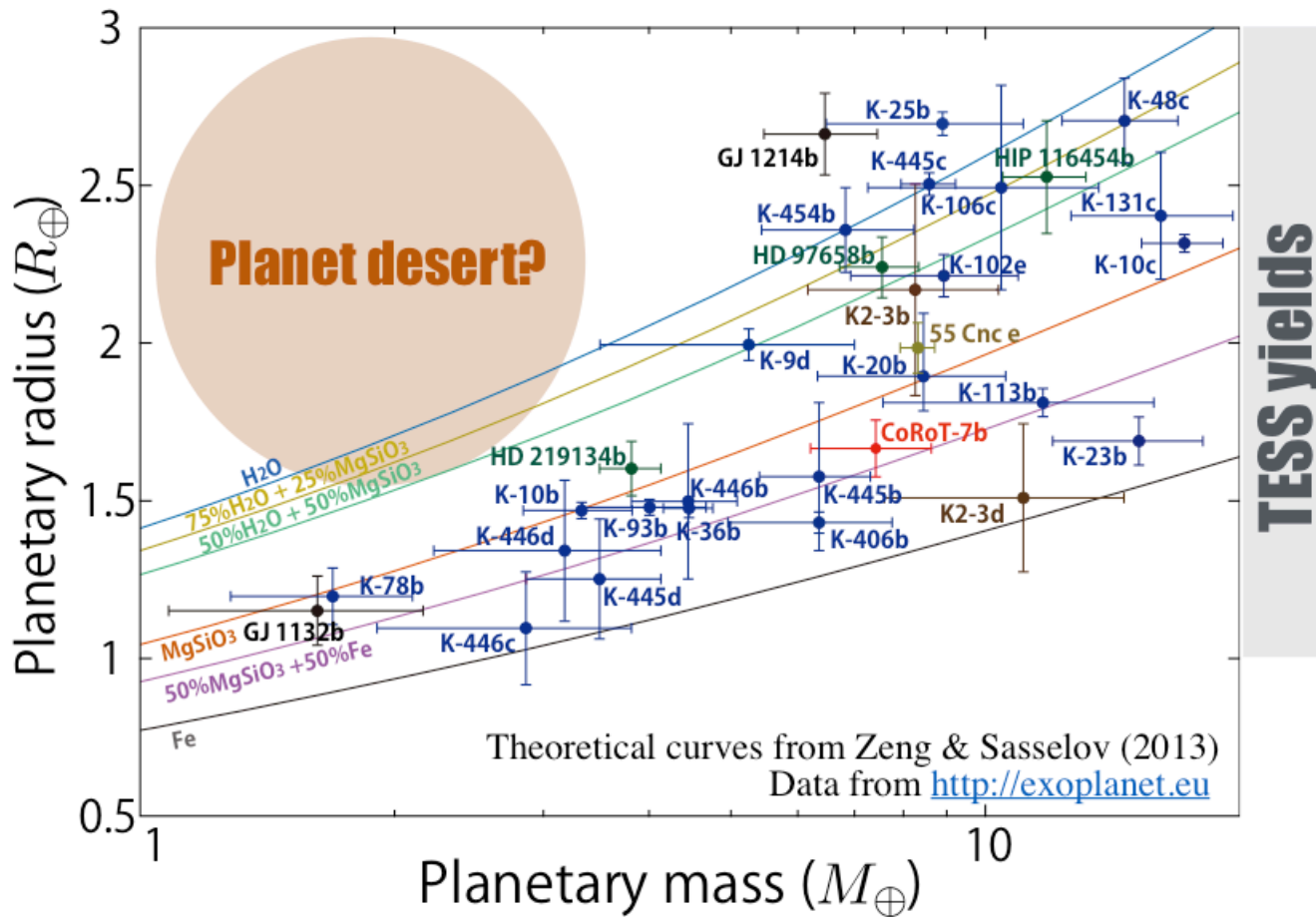
- about 300 planets: smaller than  $2 R_{\text{earth}}$ 
  - among those, ~165 planets are orbiting M dwarfs
  - ~100 planets are orbiting  $L < 10$  host stars
  - about 20 planets are orbiting within or near habitable zone
- about 650 planets: between  $2 R_{\text{earth}}$  and  $4 R_{\text{earth}}$ 
  - boundary of “super-Earth” and “mini-Neptune”
  - also important targets for planetary science



# TESS and Subaru/IRD can reveal mass-radius relationship of small planets



# To Be, or Not to Be Rocky Planets ?



- The paucity of planets with  $\lesssim 3 M_{\oplus}$  and  $\sim 1.5 - 3 R_{\oplus}$  , which corresponds to transition from rocky planets to gas dwarfs? (Weiss & Marcy, 2014; Rogers, 2014; Marcy+14; Buchhave+14; Wolfgang & Lopez, 2014)

- Earth-sized planets are “Earth-like” planets? — Small planets are dry, wet, or waterlogged?

Courtesy of Y. Hori

# Other Science Cases: Not Only Exoplanet!

- **Asteroseismology**
  - Stellar physics, Galactic archaeology
- Transient
  - Supernovae, GRBs, Gravitational wave counterparts
- Variable stars
  - Cepheid, Mira, Eclipsing binaries (young and dying)
- Stellar flares
- Solar system small bodies
- Variability of AGN/Blazars/Quasars
- Unknown fast transient (new time domain science)

# What Subaru can do with TESS?

- TESS plans to release TESS object of interest (TOI) data and Full Frame Image (FFI) data every 4 months
- Subaru can follow-up TOI targets (i.e. exoplanets) and other interesting objects (asteroseismology target, transient, etc) soon after the data releases
- Very high scientific productivity is expected

# Subaru-TESS Synergies in Exoplanet Studies

- **Validation** of very interesting planets
    - within 1 year after data release
    - AO imaging to eliminate false positives: [AO188/SCEXAO](#)
  - **Characterization** (mass, orbit, atmospheres)
    - within 2-4 years
    - [IRD](#) is especially useful to measure RVs of M dwarfs
    - [FOCAS/MOIRCS \(SWIMS?\)](#) are useful for studies of atmos.
- Providing the most interesting ~20 transiting planets for TMT's FL from Subaru -> toward TMT key program

# Subaru-TESS Synergies in Other Studies

- **Asteroseismology:**
  - spectroscopic (high resolution) follow-up
- **Galactic archeology:**
  - spectroscopic (high resolution) follow-up
- **Transient:**
  - spectroscopic (low, medium resolution) follow-up
- **Super-flare stars:**
  - spectroscopic (high resolution) follow-up

FOCAS/HDS/IRD will be useful for those studies

# Establishment of Japanese TESS Consortium

Aims:

## 1. Detailed considerations of Subaru-TESS synergetic campaign

- What kind of science?
- How many nights?
- Which instruments?

## 2. Providing latest information of TESS to Japanese community

- preparing Guest Investigator Proposal
- 4 Japanese researchers are joining the TESS team as collaborators and can access to latest information

# Establishment of Japanese TESS Consortium

Previous activities:

- **2015 Nov. 26-27: Japanese TESS Science Workshop**
  - about 60 attendees
  - 19 talks for not only exoplanets, but also asteroseismology, galactic archeology, supernovae, transient, variable stars, Wolf-Rayet stars, super-flare stars
  - Following the discussions, we established Japanese-TESS consortium here
- 2015 Dec. 17: First regular meeting
- 2015 Dec. 18: Web-ex meeting with TESS board members

# Upcoming Activities

- 2016 Jan. 21: Subaru UM FY 2015
  - Introduction of TESS mission and Japanese TESS consortium to the Subaru community
- 2016 Jan or Feb: Subaru Advisory Committee
  - introduction of Japanese TESS consortium and ask to start discussions for Subaru-TESS synergetic campaign
- 2016 Jan: Call for more members to the consortium

# Considering Subaru-TESS Synergetic Campaign

- Exoplanets
  - We have estimated **the number of necessary Subaru IRD nights** to determine mass/orbit of interesting TESS planet candidates
- Other Science Cases
  - We will start discussions with consortium members of other fields after call for members
  - **tennet, gopira ML**

# Summary

- TESS is the next flagship exoplanet mission to search for interesting exoplanets in the solar neighborhood
- Not only exoplanets, but also other fields (such as stars, transient, and galactic archeology) can receive benefits from TESS data
- Subaru-TESS synergetic campaign would benefit Japanese astronomical community